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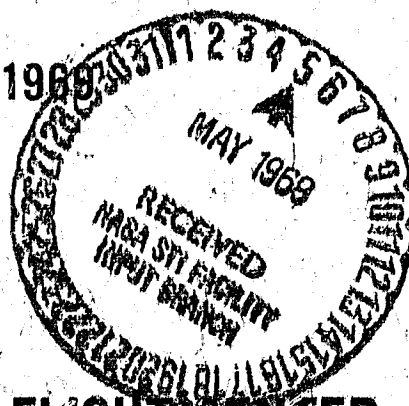
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**A COMPUTER PROGRAM  
FOR ELECTRON PROBE MICROANALYSIS  
IN THE  
FIELDS OF METALLURGY AND GEOLOGY**

**J. I. GOLDSTEIN  
P. A. COMELLA**

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A COMPUTER PROGRAM FOR ELECTRON PROBE MICROANALYSIS  
IN THE FIELDS OF METALLURGY AND GEOLOGY

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ABSTRACT

A comprehensive computer program has been written which should prove useful to investigators in both the fields of metallurgy and geology. The program is flexible in both input and output and is relatively easy for an inexperienced person to use. It is written in the Fortran language and can be used in remote terminal operation which is now available in many laboratories. The program is built up of a large number of subroutines which can be easily changed as the user sees fit or as new correction schemes become available. At present, the following corrections are incorporated; the absorption correction of Philibert-Duncumb<sup>1,2</sup> recently modified by Heinrich,<sup>3</sup> the atomic number correction of Duncumb and Reed,<sup>4</sup> and the fluorescence correction of Reed.<sup>5</sup>  $K_{\alpha}$  x-ray radiation from elements of atomic numbers 3-39 and  $L_{\alpha}$  radiation from elements of atomic numbers 20-92 are considered. The program operates in either of two modes; conversion of raw intensity data to composition using an iterative procedure, or conversion of compositional data to expected x-ray intensities. For problems in geology involving oxides, calculations are made with oxygen considered as a matrix element. Specimens and complex standards each containing up to nine elements can be treated and several standards can be used for each element measured in order to check the consistency of the answer. The program determines internally whether a fluorescence correction is necessary and also calculates the initial intensity ratios. The program does not correct for non-normal incidence or continuum excitation. An exhaustive print out of the various correction factors can be obtained if desired.

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The program is rather unique with respect to the method of data handling and ease of operation. All the critical data for a given element, such as x-ray wave lengths, fluorescence yields, atomic weights, absorption edge jump ratios and the necessary factors for the calculation of mass absorption coefficients from Heinrich<sup>6</sup> are stored on cards or are placed on file at the computer. These data are then used as needed for the elements under consideration in each problem. The compositional data for each of the element standards are also stored on cards or at the computer. The input data which specifies the problem of interest then requires a minimum of information and can be supplied in a non-rigid format.

## A COMPUTER PROGRAM FOR ELECTRON MICROPROBE ANALYSIS IN THE FIELDS OF METALLURGY AND GEOLOGY

### I. INTRODUCTION

The electron microprobe has been used to solve problems in many fields of scientific research such as metallurgy, geology and biology. In the area of quantitative microchemistry the results of microprobe analysis have not fully met the expectations of the early pioneers. Schemes which correct the x-ray data for the effect of atomic number, absorption and fluorescence have been used extensively. However at this time precision errors (Ziebold)<sup>7</sup> and errors in the input parameters to the various correction schemes (Heinrich and Yakowitz)<sup>8</sup> limit the precision and accuracy of the analysis to no better than about 2% relative. Comparisons between measured intensity data from well characterized standards and calculated intensity ratios (Ziebold and Ogilvie,<sup>9</sup> Goldstein, et al.<sup>10</sup> and Beaman<sup>11</sup>) support this conclusion.

To avoid using any corrections for microprobe analysis, standard samples with similar compositions to the analyzed sample must be obtained. In the vast majority of cases this is not practical, and the correction schemes must be used. As the number of unknown elements, standard samples and data points desired increases, the amount of data handling and calculation becomes excessive. The use of computer techniques allows rapid handling of data, accurate calculation of correction procedures and assimilation of output in a useful form. With the development of remote terminal operation and rapid microprobe readout devices, the need for computer programs for quantitative electron microprobe analysis has developed.

About 20 computer programs for electron probe microanalysis are now available. Descriptions of the various programs are presently being compiled by Beaman<sup>12</sup> and need not be discussed here. In general many of these programs are either inflexible, are written for specific computer systems or are only able to handle a small number of elements or standards. Since problems originating in various disciplines are often run in the same microprobe facility, a program was designed that could be used to treat both geologic and metallurgical problems. We desired a flexible program built on a modular design with minimal input and data handling, compatible with remote terminal operation, and incorporating the best correction procedures currently available.

This paper describes the program we have developed. The program design, calculation schemes and input-output are described in sufficient detail so that the computer program can be adopted for use in most laboratories. Several sample problems are discussed in order to illustrate the use of the program.

## II. CORRECTION PROCEDURES FOR ELECTRON MICROPROBE ANALYSIS

### A. Corrections to Raw Intensity Data

To obtain an analysis for each element in a sample of interest the following measurements are made: (1) the x-ray intensity from the element standard taken both before and after the analysis, (2) the x-ray intensity from the background standard or the continuum radiation above and below the x-ray peak, and (3) the x-ray intensity for each element from the sample of interest. Three changes are made in the raw data, so that it is in a suitable form for calculations of the x-ray corrections.

1. Detector Dead Time – Detector dead-time causes a discrepancy between the true count rate of x-rays incident on a detector and the measured count rate. This relation is expressed as:

$$N = n/(1 - n\tau) \quad (1)$$

where  $N$  is the true count rate,  $n$  is the measured count rate and  $\tau$  is the detector dead time. This correction is applied separately for each element.

2. Drift – Drift is the slow change in x-ray signal caused by instabilities in the electronic system of the microanalyzer. Normally this change is less than 1% unless unusually long analysis times are used. In this program, if the standard intensity for each element differs, from the beginning to the end of an analysis, by more than 6 times the square root of the average standard intensity, a drift correction is applied. It has been found in practice that if the change in a monitored signal is no more than 5% per hour, then it may be assumed that the drift is linear in time. A linear drift correction is therefore applied when necessary.

3. Background – Background intensity arises from the continuous x-ray spectrum generated in the sample. It is obtained for each element by either measurement of the element intensity from a background standard or at spectrometer settings away from the Bragg angle. When samples of interest contain



many elements it is difficult to obtain an ideal background standard, that is, one which contains all the elements except the element being measured. Therefore the background is commonly measured by the second method. In the computer program, the background intensity is subtracted directly after the data is corrected for dead time.

#### B. Correction Procedures for the Conversion of X-Ray Intensity to Composition

It was first demonstrated by Castaing<sup>13</sup> that the relationship between the mass concentration,  $C_A$ , of an element, A, in a specimen analyzed in the microprobe and the ratio ( $k_A = I_A/I(A)$ ) of the intensity of the characteristic x-ray line of element A measured from the specimen,  $I_A$ , to that measured from a standard,  $I(A)$ , is, to a first approximation, linear:

$$k_A = C_A/C(A) \quad (2)$$

where  $C(A)$  is the amount of A present in the standard. If the standard consists of pure A, then

$$k_A = C_A \quad (3)$$

Departures from the simple law of proportionality expressed in Equations 2, 3 can be large. Therefore if quantitative microchemical analysis is to be attempted with the microprobe, it is essential to apply a correction factor:

$$k_A = I_A/I(A) = \frac{C_A (\text{correction factor})_{\text{unknown}}}{C(A) (\text{correction factor})_{\text{std.}}} \quad (4)$$

This factor, is needed for the standard when it is not a pure element. The correction factor is placed on the right hand side of the equation because it is itself a function of  $C_A$ ,  $C(A)$  and of the concentrations of the other elements present.

Following Castaing, most authors have divided the total correction into 3 separate factors: atomic number, absorption and fluorescence. The atomic number correction takes into account the effect of composition or atomic number on the intensity of characteristic x-rays generated within a specimen. The absorption correction allows for the absorption of characteristic x-rays in emerging from the specimen. The fluorescence correction allows for the additional intensity

contributed by fluorescence radiation which is due to excitation by x-rays from other elements in the specimen. The measured intensity ratio is therefore given by:

$$k_A = \frac{I_A}{I(A)} = \frac{C_A [At.\#]_u [Abs]_u [fluor]_u}{C(A) [At.\#]_s [Abs]_s [fluor]_s} \quad (5)$$

where u represents the specimen of interest and s represents the standard. In most geologic problems C(A) is not equal to 1.0 and Equation 5 must be applied.

In the following sections the three corrections which are applied to unknown and standard intensities are described in detail. They were chosen as the best correction schemes available to date. It should be noted that none of these schemes correct for a non-normal electron beam. In the program the three correction schemes are incorporated in separate subroutines so that they can be easily changed if better corrections are developed.

1. Atomic Number Correction — The atomic number correction takes account of the fact that both the fraction of the total energy loss of an electron going into ionization of a particular shell (S) and the loss of ionization efficiency due to backscattering (R) are functions of atomic number. The atomic number correction is usually expressed as a ratio of backscattering to stopping power:

$$\begin{aligned} [At. No.]_u &= [\bar{R}/\bar{S}]_u \\ [At. No.]_s &= [\bar{R}/\bar{S}]_s \end{aligned} \quad (6)$$

Of the 3 major corrections, the atomic number correction appears to be the least developed. Uncertainties appear to lie in the accuracy of the R values available, in the validity of the Bethe equation for stopping power and in the value of J to be used in the Bethe stopping power. The correction according to Duncumb and Reed<sup>4</sup> provides a method for the correction that is easy to apply and appears to be achieving some acceptance. This formulation provides accuracies of the order of a few percent (Duncumb et al.).<sup>14</sup>

According to Duncumb-Reed,<sup>4</sup>

$$\bar{R} = \sum_i R_i C_i$$

for a complex sample where  $R_i$  is the backscatter coefficient for each element,  $C_i$  is the mass fraction of the  $i^{\text{th}}$  element and  $R_i$  is a function of  $1/U$ , the reciprocal of the overvoltage.  $\bar{S}$  may be obtained by using an expression of the form

$$\bar{S} = \sum_i C_i S_i$$

where:

$$S_i = Z_i/A_i \ln \left( \frac{1.166 \times 10^3 \bar{E}}{J_i} \right) \quad (7)$$

and  $\bar{E} = (E_0 - E_c)/2$  for each element in the sample.  $E_0$  is the operating voltage and  $E_c$  is the excitation voltage of the x-ray line of interest. The mean ionization potential  $J_i$  was determined empirically as a function of atomic number  $Z$ .<sup>4</sup> The atomic number correction  $\bar{R}/\bar{S}$  can therefore be calculated for both unknown and standard. It should be noted that the correction is the same for all the elements in the sample.

2. Absorption Correction – The absorption correction for both the analyzed sample and standard is given by the absorption function  $f(\chi)$ , where  $[Abs]_u = f(\chi)_u$  and  $[Abs]_s = f(\chi)_s$ . The absorption correction procedure of Philibert<sup>1</sup> modified by Duncumb and Shields<sup>2</sup> is widely accepted as satisfactory, especially if  $f(\chi)$  has a value between 0.3 and 1.0 (Yakowitz and Heinrich).<sup>8</sup>

The absorption function for each element analyzed in a sample is given by:

$$[Abs]_u = f(\chi)_u = 1 + h / \left( (1 + \chi/\sigma) [1 + h (1 + \chi/\sigma)]_u \right), \quad (8)$$

$$[Abs]_s = f(\chi)_s = 1 + h / \left( (1 + \chi/\sigma) [1 + h (1 + \chi/\sigma)]_s \right)$$

in which

1.  $\chi = (\mu/\rho) \csc \theta$  where  $\mu/\rho$  is the mass absorption coefficient of the target for the measured x-ray line;

$$\mu/\rho = \sum_i C_i \mu/\rho_i \quad (9)$$

and  $C_i$  = wt.% of element  $i$

$\mu/\rho_i$  = mass absorption coefficient of element  $i$  for the x-ray line of interest.

$\theta$  = is the take-off angle of the emitted radiation.

2.

$$h = 1.2 \sum_i \alpha_i A_i / \left( \sum_i \alpha_i Z_i \right)^2 \quad (10)$$

where  $\alpha_i$  is the atomic concentration;  $A_i$ , the atomic weight; and  $Z_i$ , the atomic number of element  $i$  in the multi-element sample. The atomic concentration,  $\alpha_i$ , is approximated by the weight fraction,  $C_i$ , within the program.

3.

$$\sigma = \frac{C \times 10^5}{E_0^n - E_c^n} \quad (11)$$

where  $E_c$  is the excitation potential of the element of interest and  $E_0$  is the electron beam operating voltage. Heinrich<sup>3</sup> proposed the values,  $C = 4.5$  and  $n = 1.65$ . Recent studies by Duncumb, et al.<sup>14</sup> Show these values of  $C$  and  $n$  to be about optimum. The absorption function,  $f(X)$ , must be recalculated when another element is analyzed or when  $C_A$  changes. This correction limits the general applicability of the calculation procedure because  $\mu/\rho$  values are not available for long wavelength  $L_\alpha$  radiation.

**3. Fluorescence Correction** – The fluorescence correction for characteristic x-rays is required when one or more of the elements present in the specimen or standard has a characteristic emission line with energy greater than the critical excitation potential of the line being measured. In this case a fraction of the intensity observed,  $I_f/I_p$ , is due to fluorescence excitation rather than direct excitation.  $I_f$  represents the intensity due to fluorescence and  $I_p$  represents the primary intensity from the radiation of interest. The fluorescence correction for characteristic x-rays is given by:

$$\begin{aligned} [\text{fluor}]_u &= \left[ 1 + \sum_i \frac{I_f}{I_p} \right]_u \\ [\text{fluor}]_s &= \left[ 1 + \sum_i \frac{I_f}{I_p} \right]_s \end{aligned} \quad (12)$$

where  $I_f/I_p$  is summed over all the elements which cause fluorescence. The secondary fluorescence correction procedure developed by Castaing<sup>13</sup> and modified by Reed<sup>5</sup> has received the widest acceptance and seems to be the most suitable for general usage. The modified version of Castaing's correction formula for K - K fluorescence involving L lines is given by:

$$\left(\frac{I_f}{I_p}\right)_A = 0.5 P_{ij} C_B \left(\frac{r_A - 1}{r_A}\right) W(B) \frac{A}{B} \left(\frac{U_B - 1}{U_A - 1}\right)^{1.67} \frac{\mu_B^A}{\mu_B} \left(\frac{\ln(1+X)}{X} + \frac{\ln(1+Y)}{Y}\right) \quad (13)$$

A is the element fluoresced

B is the element whose characteristic radiation excites element A radiation

$C_B$  is the concentration of element B

$r_A$  is the absorption edge jump ratio of element A

$W(B)$  is the K or L shell fluorescence yield of element B

A and B are the atomic weights of elements A and B.

$U_B$  is the overvoltage ratio,  $E_0/E_c$ , for element B

$U_A$  is the overvoltage ratio,  $E_0/E_c$ , for element A

$\mu_B^A$  is the mass absorption coefficient of element A for radiation from element B.

$\mu_B$  is the mass absorption coefficient of the specimen for radiation from element B.

The absorption parameters X and Y which appear in the above expression are given by:

$$X = (\mu_A/\mu_B) \csc \theta, \quad Y = \sigma/\mu_B \quad (14)$$

where  $\mu_A$  is the mass absorption coefficient of the specimen for A radiation,  $\sigma$  is the electron mass absorption coefficient and  $\theta$  is the take-off angle of the emitted radiation.  $P_{ij}$  takes account of the extra factor required in K-L and L-K fluorescence, where j represents the excited line ( $K_\alpha, K_\beta, L_\alpha$ ) and i represents the excitation line ( $K_\alpha, K_\beta, L_\alpha$ ). When only certain lines of K or L spectrum produce fluorescence excitation special precautions must be taken. The value of  $(r_A - 1)/r_A$  for L radiation is equal to 0.75 (Reed).<sup>5</sup>

The correction for fluorescence by the x-ray continuum has been developed by Henoc<sup>15</sup> and by Springer.<sup>16</sup> Generally it has been thought to be a small correction. Recently however, Brown, et al.<sup>17</sup> have shown that this correction is larger than often considered in the past. At this time our computer program does not include this correction.

### III. GENERAL DESCRIPTION OF THE PROGRAM

A. Flow Diagram – The microprobe correction program is designed to run in either of 2 modes. Mode 1 allows one to compute the composition of the sample of interest from measured data ( $I_A$  and  $I(A)$ ). Mode 2 allows one to calculate the expected intensity ratios ( $I_A/I(A)$ ) from the known composition of the sample. The flow diagram for the correction program is given in Figure 7.

1. Mode 1 – The input data contains the parameters necessary for the calculation steps and the raw intensity data from both sample and standards. The raw intensity data is then compiled and the background, dead time and drift corrections are made. The mass absorption coefficients for all the elements analyzed and for all the elements which are present in both samples and standards are then calculated using the method of Heinrich.<sup>6</sup> The corrections, ZAF, (Z is the atomic number correction, A is the absorption correction and F is the fluorescence correction) are then computed for the elements measured in each of the standards in turn (Equations 6, 8 and 12). If any of the standards are pure elements ( $C(A) = 1.0$ ) no secondary fluorescence correction, F, is made.

The intensity data from the first sample data point is divided by the intensity from the corresponding element standard to obtain the measured intensity ratio,  $I_A/I(A)$ . This procedure is repeated for all the elements measured in the first sample data point. Then an initial estimate of the composition of all the elements measured in the sample,  $C_{\text{unknown}}^0$ , is made by assuming the linear relation

$$C_{\text{unknown}}^0 = C_A = (I_A/I(A))_{\text{measured}} C(A). \quad (15)$$

If there are any elements in the sample which are not measured, their compositions are read into the program as input. If the sample contains oxygen, a corresponding oxygen content is calculated from the assumed cation content of each element in the sample. The total amount of oxygen in the sample is then obtained to be used in the following calculations. The ZAF corrections are applied to the values of  $C_{\text{unknown}}^0$  and a calculated value of  $I_A$  is obtained,  $(I_A)_{\text{calc}}$ :

$$(I_A)_{\text{calc.}} = (ZAF) C_{\text{unknown}}^o \quad (16)$$

The calculated intensity ratio for each element of interest is then obtained,  $k_{\text{calc}} = (I_A / I(A))_{\text{calc.}}$ . A new estimate of the composition  $C'_A$  can then be determined for each element from the equation:

$$C'_{\text{unknown}} = \frac{(I_A / I(A))_{\text{measured}}}{(I_A / I(A))_{\text{calculated}}} \times C_{\text{unknown}}^o \quad (17)$$

A test for the convergence of each element is made in the following fashion:

$$\text{if } \left| \frac{(I_A / I(A))_{\text{calculated}}}{(I_A / I(A))_{\text{measured}}} - 1 \right| < 5 \times 10^{-3} \quad (18)$$

sufficient convergence has occurred. If any of the elements do not converge, another iteration with new values of  $C'_{\text{unknown}}$  is made. Iterations are continued until convergence is achieved for all elements.

When convergence is achieved, the calculated compositions are read out. The next point on the sample or another sample is selected and the calculations are repeated. It is only necessary to use new standards and re-calculate standard corrections when a new problem is introduced.

2. Mode 2 — In this mode the compositions of both the standard and sample are input as data. One iteration of the correction procedure is made in order to calculate the intensity ratios,  $(I_A / I(A))_{\text{calculated}}$  for each element in the sample.

#### B. Structure of the Computer Program

The program is written in the Fortran IV language according to specifications for its use on IBM S/360 computers. It is built up in modular form using subroutines each of which performs one task. A listing of the program and a glossary of terms used in the conversion program are given in the Appendices I and II respectively. Detailed flow charts and descriptions of the most important subroutines are found in Appendix III. These can be used when the program is modified or initially set up on another type of computer. The subroutines are described briefly in the following paragraphs.

**MAIN** – Organizes the input data, directs execution of the subroutines, calculates  $\sigma$  (Equation 11), which is used in the absorption and fluorescence corrections, and coordinates the output of the final results.

**DATPTS** – Associates for each of the measured elements, the standard background and sample intensity values as well as the time for analysis. It also associates the intensity values with the spectrometers used and with the compositions of the standards used for each element. DATPTS inputs the intensities measured by the probe to MAIN after calculating intensity in counts per second and is the one subroutine which is most dependent on the equipment which records the microprobe data (scalers, data translator, typewriter or tape readout). Therefore this is the routine which users will probably wish to tailor to their individual equipment.

**MODIFY** – Computes the drift correction, if necessary, to be applied to the measured data. It also calculates dead time and background corrections for the measured standards and samples.

**ABSCO** – Computes the mass absorption coefficients, according to Heinrich,<sup>6</sup> for  $K_\alpha$ ,  $K_\beta$  and  $L_\alpha$  radiation from all the elements present in both the standards and samples as absorbed by each of these elements. Some of these cannot be determined<sup>6</sup> and are calculated as zero.

**ATNCOR** – Computes the atomic number correction for the standards and samples.

**INTERP** – Interpolates linearly from the table for R given by Duncumb and Reed<sup>4</sup> to obtain the value of the backscatter coefficient. The coefficient R is a function of both the atomic number of the element of interest, Z, and the reciprocal of the overvoltage,  $1/U$ .

**ABSCOR** – Computes the absorption correction for any measured element in either sample or standard.

**FLUCOR** – Computes the secondary fluorescence correction for K or L radiation produced by either  $K_\alpha$ ,  $K_\beta$  or  $L_\alpha$  radiation from any element present in either sample or standard. The total fluorescence correction (Equations 12-14)

$$\left[ 1 + \sum_i I_f / I_p \right]$$

is obtained by the summation of the effect from the various elements in the sample.



**TEST** – Determines whether secondary fluorescence by a sample matrix element can excite the element of interest. This is done by determining whether the  $K_\alpha$ ,  $K_\beta$  and  $L_\alpha$  line produced by each element present in the sample is less than the  $K$  or  $L$  edge of the measured element.

**CHANGE** – Calculates a new estimate for the composition of the unknown (Equation 17) at the end of each iteration.

**ENDIT** – Tests for convergence of each element in the sample at the end of each iteration (Equation 18). If convergence is achieved for all elements the composition of the sample is printed out.

**OXYGEN** – Calculates the amount of oxygen associated with each cation  $i$  if oxygen is present in the sample. The following equation is used to calculate the oxygen content:

$$C_{\text{oxy}} \text{ for cation } i = \frac{16 C_i}{A_i OX_i} \quad (19)$$

where  $OX_i = N/M$  if the oxide has the stoichiometric formula  $A_N O_M$ .

$C_i$  = wt. % of element  $i$

$A_i$  = atomic weight of element  $i$

The total amount of oxygen present in the sample is given by

$$C_{\text{oxy}} (\text{total}) = 16 \sum_{i=1}^N \frac{C_i}{A_i OX_i} + C_o \quad (20)$$

where  $N$  is the number of measured cations in the sample and  $C_o$  is the amount of oxygen associated with cations in the sample which are not measured.  $C_o$  is given as input data.

#### IV. INPUT TO THE PROGRAM

The input data for the program includes all the necessary physical parameters and compositional information for calculation of the corrections. If Mode 1 is used the measured intensity data are also included as input. The method of input used for the program is a compromise between incorporating many of the parameters into the basic program, which uses up computer storage, and developing very flexible input which increases the time required for the analyst

to organize and run a problem. The input data are broken up into four files which are considered separately.

- A. Data File – This file contains the parameters which describe the individual problem and operating conditions of the electron probe.
- B. Element File – This file contains the parameters which are a function of a particular element. This file is already available and the elements which will be considered in the problem are selected by the analyst before the run is made.
- C. Composition File – This file contains the composition of the standards used for the analysis. The standards are usually available well in advance of a particular run. The data contained in the composition file are often stored on cards or at the computer.
- D. Intensity Data File – This file contains the intensity readings from the probe. In many cases, punched tape or cards which are suitable for direct submission to the computer are obtained as output from the microprobe.

The Element file and the Composition file are available beforehand, and in remote terminal usage individual file members can be selected and compiled into the data set for a problem just prior to executed time. The intensity readings from the probe are available or can be made available in suitable form. Therefore, the only input cards that need be punched, before a computer run can be made, are the DATA cards. Sets A, B, C are input in the MAIN program while subroutine DATPTS inputs set D. Appendix IV considers each of the 4 types of input data separately and describes how the input must be set up to make a successful run.

## V. OUTPUT FROM THE PROGRAM

We have selected the following output as most important. The element (E) file and selected parameters from the Data File are printed out first. Next the matrix of mass absorption coefficients are given, followed by the raw probe data as it was submitted to the computer.

The correction calculations for the standards are given next: (1) Atomic number correction for each standard, (2) Absorption correction for each element in each standard, (3) Fluorescence correction for each element in each standard. For the first point on each sample, the calculated element compositions are

printed for each iteration step and the ZAF corrections are printed for the final iteration step. For subsequent data points, only the composition values from the first and last iterations are written out.

In Mode 2 problems, the calculated intensity ratios and the ZAF corrections are printed out for each point. In Appendix V the statements which control the output will be discussed as well as methods to change the output statements if desired.

# APPENDIX I Program Listing

MAIN PROGRAM		S.N.
C*****		100
C		200
C		300
C	MICROPROBE CORRECTION PROGRAM.....MARCH,1969	400
C		500
C		600
C	J.I.GOLDSTEIN	700
C	DEPARTMENT OF METALLURGY & MATERIALS SCIENCE	800
C	LEHIGH UNIVERSITY	900
C	RETHLEHEM,PENNSYLVANIA	1000
C		1100
C	P.A. COMELLA	1200
C	MATHEMATICS AND COMPUTING BRANCH, CODE 642	1300
C	LABORATORY FOR THEORETICAL STUDIES	1400
C		1500
C		1600
C	GODDARD SPACE FLIGHT CENTER	1700
C	GREENBELT,MARYLAND	1800
C		1900
C		2000
C*****		2100
3	FORMAT('1ELEMENT' 3X,A2,11(8X,A2))	2200
4	FORMAT(' ATOM.NO',12F10.4)	2300
5	FORMAT(' ATOM.WT',12F10.4)	2400
6	FORMAT(' K ALPHA',12F10.4)	2500
7	FORMAT(' K BETA ',12F10.4)	2600
8	FORMAT(' K EDGE ',12F10.4)	2700
9	FORMAT(' L ALPHA',12F10.4)	2800
10	FORMAT(' L3 EDGE',12F10.4)	2900
11	FORMAT(' CK ',12F10.4)	3000
12	FORMAT(' NK ',12F10.4)	3100
13	FORMAT(' CKL ',12F10.4)	3200
14	FORMAT(' NKL ',12F10.4)	3300
15	FORMAT(' L1 ',12F10.4)	3400
16	FORMAT(' CL1 ',12F10.4)	3500
17	FORMAT(' L2 ',12F10.4)	3600
18	FORMAT(' CL2 ',12F10.4)	3700
19	FORMAT(' CLM ',12F10.4)	3800
20	FORMAT(' M1 ',12F10.4)	3900
21	FORMAT(' CM1 ',12F10.4)	4000
22	FORMAT(' M2 ',12F10.4)	4100
23	FORMAT(' CM2 ',12F10.4)	4200
24	FORMAT(' M3 ',12F10.4)	4300
25	FORMAT(' CM3 ',12F10.4)	4400
26	FORMAT(' M4 ',12F10.4)	4500
27	FORMAT(' M5 ',12F10.4)	4600
28	FORMAT(' CMN ',12F10.4)	4700
29	FORMAT(' N1 ',12F10.4)	4800
30	FORMAT(' VC(Z,K)'12F10.4)	4900
31	FORMAT(' VC(Z,L)'12F10.4)	5000
32	FORMAT(' J1(Z) '12(18,2X))	5100
33	FORMAT(' RK(Z) '12F10.4)	5200
34	FORMAT(' OXY(Z) '12F10.4)	5300
35	FORMAT(' OMEGA-K'12F10.4)	5400
36	FORMAT(' OMEGA-L'12F10.4)	5500
37	FORMAT(' LINE '3X,A2,11(8X,A2))	5600
38	FORMAT(' VO '12F10.4)	5700
39	FORMAT(' SPEC.NO',12(17,3X))	5800
40	FORMAT(20A4)	5900
41	FORMAT(I4,10(2X,A2))	6000

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42  FORMAT(4X,10F4.2)
43  FORMAT(' THERE ARE ',I3,' ELEMENTS IN THIS SAMPLE : ',5X,A2,9(4X,A2
1  ))
44  FORMAT(' OATOMIC NUMBER CORRECTION= ',F13.5)
45  FORMAT(
1  / ' OABSORPTION CORRECTION FOR ',A2,' .....CH)= ',F13.5,7X,
2  ' F(CH)= ',F13.5 )
46  FORMAT(' WEIGHT PERCENTAGES /100 FOR THE ELEMENTS: ',F7.4,9F8.4)
47  FORMAT('////',10',20A4)
48  FORMAT('////'OFINAL CORRECTIONS'//)
49  FORMAT(' THETA ',F10.4 /
1  ' TAU ' 3E13.5/
2  ' N ' 17/
2  ' NM ' 17/
2  ' NRGS ' 17/
2  ' MODF ' 17/
2  ' NPROR ' 17)
50  FORMAT('1')
51  FORMAT('OELEMENT ',A2,' IS NOT ACCOUNTED FOR--NOT INPUT, NOT
1  MEASURED, NOT OXYGEN')
67  FORMAT(1X,A2,1X,
1  12,2X,F5.1,2(1X,F6.3),1X,14,1X,F5.2,1X,F6.4,1X,F5.3,3(1X
1  ,F6.3),1X,F7.3/4X,F4.2,1X,F6.3,1X,F6.2,1X,F4.2,1X,F6.4,1X,F6.3,5(1
2  X,F6.2)/4X,F5.1,2(1X,F6.3,1X,F4.1),2(1X,F6.3),1X,F5.2,1X,F5.3,1X,
3  F6.5)
705  FORMAT('O ATOMIC NUMBER CORRECTION(ANC) -- ANC=ATN/ATD=(RRII/SR
1  I)*(SRS/RRS)-----('A2,')-----')
706  FORMAT(' ',6X,'ABS = ',F10.5/7X,'ARN = ',F10.5/7X,'ABD = ',F10.5)
707  FORMAT('O ABSORPTION CORRECTION(ABS) -- ABS=ARN/ABD-----('A2,
1  ')-----')
710  FORMAT(7X,'ATN= ',F10.5/7X,'ATD= ',F10.5/7X,'RRII= ',F14.5/7X,'SRII= ',F14.5
1  /7X,'RBS= ',F14.5/7X,'SRS= ',F14.5/7X)
711  FORMAT(7X,'ANC= ',F10.5)
713  FORMAT('O FLUORESCENCE CORRECTION(FLU) -- FLU=FLN/FLD-----('A
1  2,')-----')
714  FORMAT(' ',6X,'FLU = ',F10.5/7X,'FLN = ',F10.5/7X,'FLD = ',F10.5)
715  FORMAT('////' ELEMENT CHART -- WEIGHT PERCENT/ 1X,64(' - '))
716  FORMAT(' DATA PT NO ITER ',3X,A2,10(7X,A3))
717  FORMAT(15,5X,15,3X,11(2PF10.3))
718  FORMAT(15,6X,A4,3X,11(2PF10.3))
719  FORMAT('OMODE 2 PROBLEM: INPUT-COMPOSITION RATIOS...OUTPUT-INTENSI
1  TY RATIOS/' PROBLEM#,15/' ELEMENT.....INTENSITY RATIO-SAMPLE TO
3  STANDARD-ATN*ARN*FLN*(U)/ATD*ABD*FLD*(S)'/ (4X,A4,10X,F13.5))
720  FORMAT('OIN STANDARD COMBINATION #',15,2X,A2,' USES STANDARD #',13)
1  DIMENSION OXY(100),STD(9,10),AK(9),
1  AK1(10),UC(10),SC(10),SCON(9,10),ATDS(10),ABDS(9,10),FLDS(9,10),
1  RBS(10),SRS(10)
1  INTEGER*4 STD,II,RSTD
1  REAL*4 NK,NKL,L1,L2,L3,M1,M2,M3,M4,M5,N1,KALPHA,KBETA,KEDGE,LALPHA
C
1  NAMELIST /DATA/ THETA,TAU,STD,N,NEU,MODF,VO,LINE,ISPEC,NPROR,
1  NSTCOM,NM,NRGS,RSTD,M
1  COMMON/EV/VC(100,2)/JJJ/J1(100)/RVAL/RK(100)/WVAL/OMEGA(100,2)
A  /F9/ ARCO(50,50,3),SIG,THETA
1  COMMON /BACK/ RGINT(9,9),RKGR(9),RSTD(9),NRGS ,RTITLE(20,9)
X  /STAND/ ESTD(10,10),STINT(9,9,10),NESTD(10),
Y  CS(10,10),STINTD(9,10,8),COMPST(9,10)
X  /SCRIPT/ N,NU(8,3),NM,M ,INDIC
X  /INFO/ E(50),Z(50),A(50),KALPHA(50),KBETA(50),KEDGE(50),
1  LALPHA(50),L3(50),CK(50),NK(50),CKL(50),NKL(50),
2  L1(50),CL1(50),L2(50),CL2(50),CLM(50),M1(50),
3  CM1(50),M2(50),CM2(50),M3(50),CM3(50),M4(50),
4  M5(50),CMN(50),N1(50)
5  /ARGS/ LINE(10),VO,MP,I,LL,LLL,J,NCPINT
X  /CORREC/DRIFT(9,10,8), TAU(3),JSPEC(9)
X  /UNKWN/ IINT(75,9,8),NEU,C(10)

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	X	/XXX/ CHI	12700
	A	/INOUT/ IOUT,IN	12800
		REAL*4 STITLE(20,10),UTITLE(20)	12900
		SIGMA(Q,W)= 4.5F5/(O**1.65-W**1.65)	13000
		IOUT=6	13100
		IN=5	13200
		DATA LAST,L,O,SUM/'LAST','L','O','SUM'/	13300
100		READ(IN,DATA,FND=10000)	13400
		NU(1,1)=1	13500
		INDIC=1	13600
		NP1=N+1	13700
		IS=0	13800
		DO 125 I=1,M	13900
		READ(IN,67,END=10000) F(I),NO,A(I),VC(I,1),VC(I,2),J1(I),RK(I),	14000
		1 OXY(I),OMEGA(I,1)	14100
		1,KALPHA(I),KBETA(I),LALPHA(I),CK(I),NK(I),KEDGE(I),CKL(I),NKL(I),	14200
		2L1(I),CL1(I),L2(I),CL2(I),L3(I),CLM(I),M1(I),CM1(I),M2(I),CM2(I),	14300
		3 M3(I),CM3(I),M4(I),M5(I),CMN(I),N1(I),OMEGA(I,2)	14400
		Z(I)=FLOAT(NO)	14500
125		CONTINUE	14600
		MP=M	14700
		L=1	14800
		U=12	14900
130		IF(M-U) 135,135,160	15000
135		U=M	15100
160		WRITE(IOUT,3) (F(I),I=L,U)	15200
		WRITE(IOUT,4) (Z(I),I=L,U)	15300
		WRITE(IOUT,5) (A(I),I=L,U)	15400
		WRITE(IOUT,6) (KALPHA(I),I=L,U)	15500
		WRITE(IOUT,7) (KBETA(I),I=L,U)	15600
		WRITE(IOUT,8) (KEDGE(I),I=L,U)	15700
		WRITE(IOUT,9) (LALPHA(I),I=L,U)	15800
		WRITE(IOUT,10) (L3(I),I=L,U)	15900
		WRITE(IOUT,11) (CK(I),I=L,U)	16000
		WRITE(IOUT,12) (NK(I),I=L,U)	16100
		WRITE(IOUT,13) (CKL(I),I=L,U)	16200
		WRITE(IOUT,14) (NKL(I),I=L,U)	16300
		WRITE(IOUT,15) (L1(I),I=L,U)	16400
		WRITE(IOUT,16) (CL1(I),I=L,U)	16500
		WRITE(IOUT,17) (L2(I),I=L,U)	16600
		WRITE(IOUT,18) (CL2(I),I=L,U)	16700
		WRITE(IOUT,19) (CLM(I),I=L,U)	16800
		WRITE(IOUT,20) (M1(I),I=L,U)	16900
		WRITE(IOUT,21) (CM1(I),I=L,U)	17000
		WRITE(IOUT,22) (M2(I),I=L,U)	17100
		WRITE(IOUT,23) (CM2(I),I=L,U)	17200
		WRITE(IOUT,24) (M3(I),I=L,U)	17300
		WRITE(IOUT,25) (CM3(I),I=L,U)	17400
		WRITE(IOUT,26) (M4(I),I=L,U)	17500
		WRITE(IOUT,27) (M5(I),I=L,U)	17600
		WRITE(IOUT,28) (CMN(I),I=L,U)	17700
		WRITE(IOUT,29) (N1(I),I=L,U)	17800
		WRITE(IOUT,30) (VC(I,1),I=L,U)	17900
		WRITE(IOUT,31) (VC(I,2),I=L,U)	18000
		WRITE(IOUT,32) (J1(I),I=L,U)	18100
		WRITE(IOUT,33) (RK(I),I=L,U)	18200
		WRITE(IOUT,34) (OXY(I),I=L,U)	18300
		WRITE(IOUT,35) (OMEGA(I,1),I=L,U)	18400
		WRITE(IOUT,36) (OMEGA(I,2),I=L,U)	18500
		WRITE(IOUT,37) (LINE(I),I=L,U)	18600
		WRITE(IOUT,39) (ISPEC(I),I=1,N)	18700
		IF(U.GE.M) GO TO 175	18800
		L=U+1	18900
		U=U+12	19000
		WRITE(6,41) U	19100
		WRITE(6,41) M	19200
		GO TO 130	19300

175	CONTINUE	19400
	WRITE(10UT,38) VO	19500
	WRITE(10UT,49) THETA,(TAU(I),I=1,3),N,NM,NBGS,MODE	19600
	1,NPRNB	19700
	THETA=THETA*3.1415926/180.0	19800
	IOX1=0	19900
	DO 200 I=1,NEU	20000
200	IF (E(I).EQ.0) GO TO 300	20100
	GO TO 400	20200
300	IOX1=I	20300
400	CONTINUE	20400
	READ(IN,40) (UTITLE(I),I=1,20)	20500
	READ(5,2) (C(I),I=1,NEU)	20600
2	FORMAT(8F10.2)	20700
	DO 500 I=1,NEU	20800
500	C(I)=.01*C(I)	20900
	IF (IOX1.GT.0) COXY= C(IOX1)	21000
	DO 600 K=1,NM	21100
	READ(IN,40) (STITLE(I,K),I=1,20)	21200
	READ(IN,41) NEST,(ESTD(I,K),I=1,NFST)	21300
	READ(IN,2) (CS(I,K),I=1,NFST)	21400
	DO 550 I=1,NEST	21500
550	CS(I,K)=.01*CS(I,K)	21600
	NESTD(K)=NEST	21700
600	CONTINUE	21800
	IF(NBGS.EQ.0) GO TO 900	21900
	DO 800 K=1,NBGS	22000
	READ(IN,40) (BTITLE(I,K),I=1,20)	22100
800	CONTINUE	22200
900	CONTINUE	22300
	IF(NP1.GT.NEU) GO TO 920	22400
	DO 910 I=NP1,NEU	22500
	IF(E(I).EQ.0) GO TO 910	22600
	IF(C(I).GE.1.E-5) GO TO 910	22700
	WRITE(10UT,51) F(I)	22800
	STOP	22900
910	CONTINUE	23000
920	DO 930 I=1,M	23100
	IJK=1	23200
	IF(LINE(I).EQ.L) IJK=2	23300
930	LINE(I)=IJK	23400
	CALL ABSCO	23500
	IF(MODE.EQ.1) CALL DATPTS	23600
	NCOUNT=1	23700
	DO 4000 K=1,NM	23800
	NEST=NESTD(K)	23900
	CALL ATNCOR(RBS(K),SBS(K),NEST,CS(1,K),ATDS(K),ESTD(1,K))	24000
	WRITE(10UT,47) (STITLE(I,K),I=1,20)	24100
	WRITE(10UT,43) NEST,(ESTD(I,K),I=1,NFST)	24200
	WRITE(10UT,46) (CS(I,K),I=1,NFST)	24300
	WRITE(10UT,44) ATDS(K)	24400
	DO 3500 I=1,M	24500
	DO 3500 J=1,NFST	24600
	IF (E(I).NE. ESTD(J,K)) GO TO 3500	24700
	LLL=Z(I)	24800
	DO 3200 JJ=1,NSTCOM	24900
3200	IF(STD(I,JJ).EQ.K) GO TO 3300	25000
	GO TO 3500	25100
3300	LL=LINE(I)	25200
	IF(LLL.GE.37) LL=2	25300
	SIG=SIGMA(VO,VC(I,LL))	25400
	SCON(I,K)=CS(J,K)	25500
	CALL ABSOR(NEST,CS(1,K),ARDS(I,K),ESTD(1,K))	25600
	WRITE(10UT,45) E(I),CHI,ARDS(I,K)	25700
	FLDS(I,K)=1.	25800
	IF(NEST.GT.1)	25900
	1 CALL FLUCOR(CS(1,K),FLDS(I,K),ESTD(1,K),NEST)	26000

3500	CONTINUE	26100
4000	CONTINUE	26200
	DO 7000 KKKK=1,NSTCOM	26300
	WRITE(IOUT,50)	26400
	DO 7000 IGROUP=1,INDIC	26500
	NUP=NU(IGROUP,1)	26600
	DO 7000 NC=1,NUP	26700
	GO TO (4050,4100),MODE	26800
4050	NCOUNT=0	26900
	GO TO 4200	27000
4100	NCOUNT=1	27100
4200	DO 5050 I=1,N	27200
	K=STD(I,KKKK)	27300
	SC(I)=SCON(I,K)	27400
	IF(NC.EQ.1) WRITE(IOUT,720) KKKK,E(I),K	27500
	IF(MODE.EQ.2) GO TO 5050	27600
	DR=DRIFT(I,K,IGROUP)*NC	27700
	STINT=STINTD(I,K,IGROUP)+DR	27800
	AK(I)=UINT(NC,I,IGROUP)/STINT	27900
	C(I)=AK(I)*COMPST(I,K)	28000
5050	CONTINUE	28100
5075	IF (IOX1.NE.0) CALL OXYGEN(OXY,C,A, N,IOX1,COXY)	28200
	SUMM=0.0	28300
	DO 5150 J=1,NEU	28400
	UC(J)=C(J)	28500
5150	SUMM=SUMM+UC(J)	28600
	INDEX=0	28700
	IF(NC.GT.1) GO TO 5200	28800
	WRITE(IOUT,715)	28900
	WRITE(IOUT,716) (E(I),I=1,NEU),SUM	29000
5200	CONTINUE	29100
	WRITE(IOUT,717)NC,INDEX,(UC(I),I=1,NEU),SUMM	29200
5250	DO 5500 INDEX=1,10	29300
472	CONTINUE	29400
	CALL ATNCOR(RRU,SRU,NEU,UC,ATN,E)	29500
	DO 5400 I=1,N	29600
	K=STD(I,KKKK)	29700
5325	LLL=Z(I)	29800
	LL=LINE(I)	29900
	IF(LLL.GE.37) LL=2	30000
	SIG=SIGMA(VO,VC(I,LL))	30100
	CALL ARSCOR(NEU,UC,ARN,E)	30200
	CALL FLUCOR(UC,FLN,E,NEU)	30300
	ANC=ATN/ATDS(K)	30400
	ARS=ARN/ARDS(I,K)	30500
	FLU=FLN/FLDS(I,K)	30600
	AAA=ANC*ARS*FLU	30700
	AK1(I)=AAA*UC(I)/SC(I)	30800
	IF(NCOUNT.EQ.0) GO TO 5400	30900
	WRITE(IOUT,713) E(I)	31000
	WRITE(IOUT,714) FLU,FLN,FLDS(I,K)	31100
	WRITE(IOUT,707) E(I)	31200
	WRITE(IOUT,706) ARS,ARN,ARDS(I,K)	31300
	WRITE(IOUT,705) E(I)	31400
	WRITE(IOUT,711) ANC	31500
	WRITE(IOUT,710) ATN,ATDS(K),RRU,SRU,RBS(K),SBS(K)	31600
5400	CONTINUE	31700
	IF(MODE.EQ.2) GO TO 5550	31800
	IF(NCOUNT.EQ.1) GO TO 5575	31900
	CALL CHANGE(N,AK1,AK,UC)	32000
	IF (IOX1.NE.0) CALL OXYGEN(OXY,UC,A, N,IOX1,COXY)	32100
	IF(NC.NE.1) GO TO 5475	32200
	SUMM=0.	32300
	DO 5450 I=1,NEU	32400
5450	SUMM=SUMM+UC(I)	32500
	WRITE(6,717) NC,INDEX,(UC(I),I=1,NEU),SUMM	32600



5475	CALL ENDIT(N,AK1,AK,65525)	32700
5500	CONTINUE	32800
5525	IF(NC.NE.1) GO TO 5600	32900
	NCOUNT=1	33000
	WRITE(IOUT,48)	33100
	GO TO 472	33200
5550	WRITE(6,719) NPROR,(F(I),AK1(I),I=1,N)	33300
	GO TO 7000	33400
5575	IF(NC.NE.1) GO TO 5600	33500
	WRITE(IOUT,715)	33600
	WRITE(IOUT,716) (E(I),I=1,NEU),SUM	33700
5600	IF (IOX1.NE.0) CALL OXYGEN(OXY,UC,A, N,IOX1,COXY)	33800
	SUMM=0.0	33900
	DO 5625 I=1,NEU	34000
5625	SUMM=SUMM+UC(I)	34100
	INDEX=LAST	34200
5650	WRITE(IOUT,718) NC,INDEX,(UC(I),I=1,NEU),SUMM	34300
7000	CONTINUE	34400
	GO TO 100	34500
10000	RETURN	34600
	END	

	<b>SUBROUTINE ABSCO</b>	100
	REAL*4 KALPHA,KRETA,KEDGE,LALPHA,LAMBDA	200
	REAL*4 NK,NKL,L1,L2,L3,M1,M2,M3,M4,M5,N1	300
	COMMON /SCRIPT/ N,NU(8,3),NM,M	400
X	/INFO/ E(50),Z(50),A(50),KALPHA(50),KRETA(50),KEDGE(50),	500
1	LALPHA(50),L3(50),CK(50),NK(50),CKL(50),NKL(50),	600
2	L1(50),CL1(50),L2(50),CL2(50),CLM(50),M1(50),	700
3	CM1(50),M2(50),CM2(50),M3(50),CM3(50),M4(50),	800
4	M5(50),CMN(50),N1(50)	900
X	/F9/ ABCO(50,50,3) /IOUT/ IOUT	1000
	DIMENSION LAMBDA(50)	1100
C	THIS SUBROUTINE COMPUTES AND ASSOCIATES THE MASS ABSORPTION COEFFI	1200
	EQ(X1,Y1,Z1)=X1*Y1**Z1	1300
	WRITE(IOUT,701)	1400
	DO 801 JUMP=1,3	1500
	GO TO (841,844,842),JUMP	1600
841	DO 851 J=1,M	1700
851	LAMBDA(J)=KALPHA(J)	1800
	GO TO 814	1900
842	DO 852 J=1,M	2000
852	LAMBDA(J)=KRETA(J)	2100
	GO TO 814	2200
844	DO 854 J=1,M	2300
854	LAMBDA(J)=LALPHA(J)	2400
814	DO 801 I=1,M	2500
	DO 801 J=1,M	2600
	ABCO(I,J,JUMP)=0.0	2700
	IF(LAMBDA(J).EQ.0.) GO TO 801	2800
	IF(LAMBDA(J).LT.KEDGE(I)) GO TO 1	2900
	IF(L1(I).EQ.0.0) L1(I)=50.0	3000
	IF(LAMBDA(J).LT.L1(I)) GO TO 2	3100
	IF (LAMBDA(J).LT.L2(I)) GO TO 3	3200
	IF(LAMBDA(J).LT.L3(I)) GO TO 4	3300
	IF(LAMBDA(J).LT.M1(I)) GO TO 5	3400
	IF(LAMBDA(J).LT.M2(I)) GO TO 6	3500
	IF(LAMBDA(J).LT.M3(I)) GO TO 7	3600
	IF(LAMBDA(J).LT.M4(I)) GO TO 8	3700
	IF(LAMBDA(J).LT.M5(I)) GO TO 9	3800
	IF(LAMBDA(J).LE.N1(I)) GO TO 10	3900
	GO TO (1000,1100,1000),JUMP	4000
1000	WRITE(IOUT,12) LAMBDA(J),F(J),N1(I),E(I)	4100
	GO TO 5	4200

1100	WRITE(IOUT,13) LAMBDA(J),E(J),N1(I),F(I)	4300
	GO TO 5	4400
1	ABCO(I,J,JUMP)=CK(I)*LAMBDA(J)**NK(I)	4500
	GO TO 801	4600
2	ABCO(I,J,JUMP)=CKL(I)*LAMBDA(J)**NKL(I)	4700
	GO TO 801	4800
3	ABCO(I,J,JUMP)=CL1(I)*LAMBDA(J)**NKL(I)	4900
	GO TO 801	5000
4	ABCO(I,J,JUMP)=CL2(I)*LAMBDA(J)**NKL(I)	5100
	GO TO 801	5200
5	ABCO(I,J,JUMP)=CLM(I)*LAMBDA(J)**2.60	5300
	GO TO 801	5400
6	ABCO(I,J,JUMP)=CM1(I)*LAMBDA(J)**2.60	5500
	GO TO 801	5600
7	ABCO(I,J,JUMP)=CM2(I)*LAMBDA(J)**2.60	5700
	GO TO 801	5800
8	ABCO(I,J,JUMP)=CM3(I)*LAMBDA(J)**2.60	5900
	GO TO 801	6000
9	WRITE(6,13) LAMBDA(J)	6100
	GO TO 801	6200
10	ABCO(I,J,JUMP)=CMN(I)*LAMBDA(J)**2.22	6300
801	CONTINUE	6400
	WRITE(IOUT,701)	6500
	DO 901 JUMP=1,3	6600
	GO TO (1200,1300,1400),JUMP	6700
1200	WRITE(IOUT,702)	6800
	GO TO 1500	6900
1300	WRITE(IOUT,706)	7000
	GO TO 1500	7100
1400	WRITE(IOUT,705)	7200
1500	L=1	7300
	MJ=MINO(14,M)	7400
1550	WRITE(IOUT,703) (F(I),I=L,M)	7500
	DO 1600 I=J,M	7600
1600	WRITE(IOUT,704) F(I),(ABCO(I,J,JUMP),I=L,M)	7700
	IF (MJ.GE.M) GO TO 901	7800
	L=MI+1	7900
	MJ=MINO(MJ+14,M)	8000
	GO TO 1550	8100
901	CONTINUE	8200
	701 FORMAT('CORRECTION VALUES (FIRST DATA POINT , LAST ITERATION	8300
	1)'/1X,57(' - '))	8400
	702 FORMAT(' MASS ABSORPTION COEFFICIENTS (EMITTER AT TOP) -- K	8500
	1ALPHA RADIATION')	8600
	703 FORMAT(14(7X,A2))	8700
	704 FORMAT(1X,A2,2X,14(F7.1,2X))	8800
	705 FORMAT(' MASS ABSORPTION COEFFICIENTS (EMITTER AT TOP) -- K	8900
	1BETA RADIATION')	9000
	706 FORMAT(' MASS ABSORPTION COEFFICIENTS (EMITTER AT TOP) -- L	9100
	1ALPHA RADIATION')	9200
12	FORMAT(' LAMBDA-K(' ,F6.3,' ) OF 'A4,' IS GREATER THAN N1(' ,F6.3,	9300
1	' ) OF 'A4)	9400
13	FORMAT(' LAMBDA-L(' ,F6.3,' ) OF 'A4,' IS GREATER THAN N1(' ,F6.3,	9500
1	' ) OF 'A4)	9600
	END	9700
 SUBROUTINE ABCOR (NEL,C,AB,EL)		100
	REAL*4 EL(10),C(10)	200
	COMMON/XXX/X	300
1	/INFO/ E(50),Z(50),A(50)	400
3	/F9/ ABCO(50,50,3),SIG,THETA	500
4	/ARGS/ LINE(10),VO,M,I,KK	600
	AAVE=0.0	700
	ZAVE=0.0	800
	X=0.0	900

	DO 2 J=1,NEL	1000
	DO 2 K=1,M	1100
	IF(E(K).NE.EL(J)) GO TO 2	1200
	AAVE=AAVE+C(J)*A(K)	1300
	ZAVE=ZAVE+C(J)*Z(K)	1400
	X=X+C(J)*ARCO(K,I,KK)	1500
2	CONTINUE	1600
	H=1.2*AAVE/ZAVE**2	1700
	X=X/SIN(THETA)	1800
	AR=(1.0+H)/((1.0+X/SIG)*(1.0+H*(1.0+X/SIG)))	1900
	RETURN	2000
	END	2100
	<b>SUBROUTINE ATNCOR (RB,SB,NEL,C,AT,EL)</b>	100
	DIMENSION EL(10),C(10)	200
	COMMON/EV/VC(100,2)/JJJ/J(100)	300
1	/INFO/ E(50),Z(50),A(50)	400
2	/ARGS/LINE(10),VO,M	500
	RR=0.0	600
	SR=0.0	700
	DO 1000 I=1,M	800
	DO 500 K=1,NEL	900
500	IF(E(I).EQ.EL(K)) GO TO 600	1000
	GO TO 1000	1100
600	L=Z(I)	1200
	LL=LINE(I)	1300
	IF(L.GE.37) LL=2	1400
	V=.5*(VO+VC(I,LL))	1500
	RECII=VC(I,LL)/VO	1600
	CALL INTERP(RECII,L,RAB)	1700
	RR=RR+C(K)*RAB	1800
	S=Z(I)/A(I)*ALOG((1.166F3*V/J(I)))	1900
	SR=SR+C(K)*S	2000
1000	CONTINUE	2100
	AT=RR/SR	2200
	RETURN	2300
	END	2400
	<b>SUBROUTINE CHANGE (N,AK1,AK,UC)</b>	100
	DIMENSION AK1(10),AK(10),UC(10)	200
C	THIS SUBROUTINE GIVES NEW VALUES FOR THE WEIGHT PERCENTS	300
	DO 200 I=1,N	400
	J=I	500
100	RAT=AK(I)/AK1(I)	600
	UC(J)=RAT*UC(I)	700
200	CONTINUE	800
	RETURN	900
	END	1000
	<b>SUBROUTINE DATPTS</b>	100
	INTEGER*4 RSTD ;SAME	200
	REAL*4 TBG(9,3),TSTD(9,10,3),XINT(3)	300
	COMMON/SS/S(3)	400
	COMMON /BACK/ BGINT(9,9),BKGR(9),RSTD(9),NRGS	500
X	/SCRIPT/ N,NII(8,3),NM,M,INDIC	600
X	/STAND/ ESTD(10,10),STDINT(9,9,10),NFSTD(10)	700
X	/UNKWN/ UINT(75,9,8)	800
X	/INOUT/ IOUT,IN	900
X	/INFO/ E (50)	1000
	WRITE(IOUT,701)	1100
	DO 100 K=1,NM	1200
	DO 50 L=1,9	1300
	DO 40 J=1,3	1400
	TSTD(L,K,J)=0,	1500

40	CONTINUE	1600
	DO 50 J=1,N	1700
50	STDINT(L,J,K)=0.	1800
100	CONTINUE	1900
	IF(NBGS.EQ.0) GO TO 275	2000
	DO 250 K=1,NBGS	2100
	DO 200 J=1,3	2200
	TBG(K,J)=0.	2300
200	CONTINUE	2400
	DO 250 J=1,N	2500
	BGINT(J,K)=0.	2600
250	CONTINUE	2700
275	CONTINUE	2800
	DO 280 L=1,8	2900
	DO 280 I=1,3	3000
280	NU(L,I)=0	3100
	DO 20000 ICT=1,2000	3200
3000	READ(IN,1) TIME,XINT,SAME,NUMRDG,INDIC,NOST,IUNK,LAST	3300
	GO TO (3025,3050,3075),NUMRDG	3400
3025	IT=1	3500
	INDMAX=MIN0(N,3)	3600
	GO TO (4000,5000,6000),SAME	3700
3050	IT=2	3800
	NUMRDG=4	3900
	INDMAX=MIN0(N,6)	4000
	GO TO (4000,5000,6000),SAME	4100
3075	IT=3	4200
	NUMRDG=7	4300
	INDMAX=MIN0(N,9)	4400
	GO TO (4000,5000,6000),SAME	4500
4000	NU(INDIC,IT)=NU(INDIC,IT)+1	4600
	WRITE(IOUT,702)	4700
	1 S(SAME),NU(INDIC,IT),TIME,XINT	4800
	2 ,NUMRDG,INDIC ,(E(J),J=NUMRDG,INDMAX)	4900
	K=0	5000
	DO 4100 J=NUMRDG,INDMAX	5100
	K=K+1	5200
	UINT(NU(INDIC,IT),J,INDIC)=XINT(K)/TIME	5300
4100	CONTINUE	5400
	GO TO 20000	5500
5000	TSTD(INDIC,NOST,IT)=TSTD(INDIC,NOST,IT)+TIME	5600
5050	K=0	5700
	DO 5100 J=NUMRDG,INDMAX	5800
	K=K+1	5900
	STDINT(INDIC,J,NOST)=STDINT(INDIC,J,NOST)+XINT(K)	6000
5100	CONTINUE	6100
	WRITE(IOUT,702)S(SAME),NOST,TIME,XINT,NUMRDG,INDIC,	6200
	1 (E(J),J=NUMRDG,INDMAX)	6300
	GO TO 20000	6400
6000	TBG(NOST,IT)=TBG(NOST,IT)+TIME	6500
6050	K=0	6600
	DO 6100 J=NUMRDG,INDMAX	6700
	K=K+1	6800
	BGINT(J,NOST)=BGINT(J,NOST)+XINT(K)	6900
6100	CONTINUE	7000
	WRITE(IOUT,702)S(SAME),NOST,TIME,XINT,NUMRDG,INDIC,	7100
	1 (E(J),J=NUMRDG,INDMAX)	7200
20000	IF (LAST.NE.0) GO TO 21000	7300
21000	DO 21500 K=1,NM	7400
	DO 21400 I=1,INDIC	7500
	DO 21200 J=1,N	7600
	GO TO (21010,21010,21010,21020,21020,21020,21030,21030,21030),J	7700
21010	IT=1	7800
	GO TO 21040	7900
21020	IT=2	8000
	GO TO 21040	8100
21030	IT=3	8200

21040	IF (TSTD(I,K,IT).NE.0) STDINT(I,J,K)=STDINT(I,J,K)/TSTD(I,K,IT)	8300
21200	CONTINUE	8400
21400	CONTINUE	8500
21500	CONTINUE	8600
	IF(NRGS.EQ.0) GO TO 23000	8700
	DO 22000 K=1,NRGS	8800
	DO 21600 J=1,N	8900
	IF(RSTD(J).NE.K) GO TO 21600	9000
	GO TO (21510,21510,21510,21520,21520,21520,21530,21530,21530),J	9100
21510	IT=1	9200
	GO TO 21540	9300
21520	IT=2	9400
	GO TO 21540	9500
21530	IT=3	9600
21540	RGINT(J,K)=RGINT(J,K)/TRG(K,IT)	9700
21600	CONTINUE	9800
22000	CONTINUE	9900
23000	INDIC=INDIC-1	10000
	GO TO (25000,25000,25000,23100,23100,23100,23300,23300,23300),N	10100
23100	DO 23200 I=1,INDIC	10200
23200	NU(I,1)=MIN0(NU(I,1),NU(I,2))	10300
	GO TO 25000	10400
23300	DO 24000 I=1,INDIC	10500
24000	NU(I,1)=MIN0(NU(I,1),NU(I,2),NU(I,3))	10600
25000	CALL MODIFY	10700
	RETURN	10800
701	FORMAT('INPUT INFORMATION'/1X,19(' ')/'OTYPE NUMBER TIME ',	10900
	13('INTENSITY ', 'NUMRDG',3X,'INDIC',5X,'ELEMENT',5X,'ELEMENT',	11000
	2 5X,'ELEMENT')	11100
702	FORMAT('0',A2,3X,I5,2X,F7.2,3I2X,F11.1),I9,I8,8X,A2,10X,A2,10X,A2)	11200
1	FORMAT(F6.0,3F8.0,1X ,4I1,12,I1)	11300
7	FORMAT(' ',A5)	11400
	END	11500
	SUBROUTINE ENDIT (N,AK1,AK,*)	100
	DIMENSION AK1(10),AK(10)	200
C	THIS SUBROUTINE TELLS WHETHER OR NOT TO STOP ITERATING	300
	ICON=0	400
	DO 36 I=1,N	500
	IF(ABS(AK1(I)/AK(I)-1.).GT.5.0-3) ICON=ICON+1	600
36	CONTINUE	700
	IF(ICON.EQ.0) RETURN 1	800
	RETURN	900
	END	1000
	SUBROUTINE FLUCOR (C,FL,EL,NEL)	100
	COMMON/WVAL/OMEGA(100,2)/RVAL/RK(100)/EV/VC(100,2)	200
	COMMON/XXX/X	300
1	/INFO/ E(50),Z(50),A(50)	400
5	/INOUT/ IOUT	500
3	/F9/ ABCQ(50,50,3),SIG,THETA	600
4	/ARGS/ LINE(10),VO,M,L,LL,LLL,I,NCOUNT	700
	REAL*4 LORK(2) /'K','L'/	800
	REAL*8 KAKBLA(6)/2*'K ALPHA ',2*'K BETA ',2*'L ALPHA '/	900
	DIMENSION EL(10),C(10),ITEST(6),P(6)	1000
	DATA P/1.0,0.24,0.1,0.024,4.2,1.0/	1100
	UA=VO/VC(L,LL)	1200
	AR=RK(L)	1300
	FL=1.0	1400
	DO 5 J=1,M	1500
	DO 5 IJ=1,NEL	1600
	IF(E(J).NE.EL(IJ)) GO TO 5	1700
	CALL TEST(J,L,ITEST)	1800
	SFK=0.0	1900
	DO 4 K=LL,6,2	2000
	IF(ITEST(K).NE.1) GO TO 4	2100
	GO TO (125,125,150,150,175,175),K	2200

C	KALPHA RADIATION	2300
125	K1=1	2400
	KK=1	2500
	GO TO 200	2600
C	KBETA RADIATION	2700
150	K1=3	2800
	KK=1	2900
	GO TO 200	3000
C	LALPHA RADIATION	3100
175	K1=2	3200
	KK=2	3300
200	KAY=Z(J)	3400
	IF(VC(J,KK).EQ.0.) GO TO 4	3500
	IF(ABCO(J,J,K1).EQ.0.0) GO TO 4	3600
	UR=V0/VC(J,KK)	3700
	DEN=0.0	3800
	DO 3 K2=1,M	3900
	DO 3 JJ=1,NFL	4000
	IF(F(K2).NE.FL(JJ)) GO TO 3	4100
	DEN=DEN+C(JJ)*ABCO(K2,J,K1)	4200
3	CONTINUE	4300
	XF=X/DEN	4400
	Y=SIG/DEN	4500
	IF(LL.EQ.2)AR=4.0	4600
	FK=.5*P(K)*(1.-1./AR)*OMFGA(J,KK)*(A(L)/A(J))*((UR-1.)/(UA-1.))	4700
	1*1.67*(ABCO(L,J,K1)/DEN)*((ALOG(1.+XF)/XF)+(ALOG(1.+Y)/Y))	4800
	FK=C(IJ)*FK	4900
	SFK=SFK+ FK	5000
	ITEST(K)=0	5100
	IF(NCOUNT.EQ.1) WRITE(1OUT,712) F(J),KAKBLA(K),F(L),LORK(LL),FK,	5200
1	F(J),C(IJ)	5300
4	CONTINUE	5400
	FL=FL+SFK	5500
5	CONTINUE	5600
	IF(NCOUNT.NE.1) RETURN	5700
	IF( FL.NE.1) GO TO 400	5800
	WRITE(1OUT,713) F(L)	5900
	RETURN	6000
400	WRITE(1OUT,714) F(L),FL	6100
	RETURN	6200
712	FORMAT(' ',A2,'-',A8,' FLOURESCED ',A2,'-',A2,'.',5X,'KF=',F13.5,	6300
1	7X,'C( ',A2,' )=',F13.5)	6400
713	FORMAT(' NO ELEMENT FLOURESCED ',A2,'...KF(TOTAL)= 1.0' )	6500
714	FORMAT(' TOTAL FLOURESCENCE CORRECTION FOR ',A2,'=',F13.5)	6600
	END	6700

	SUBROUTINE INTERP (RECU,IZ,RAB)	100
	COMMON/RA/R(11,11)	200
	DO 1 K=1,10	300
	IF(IZ.LE.K*10) GO TO 3	400
1	CONTINUE	500
3	I1=K	600
	I2=I1+1	700
	IF(RECU.LE..1) GO TO 7	800
	DO 6 L=2,10	900
	AL=L	1000
	IF(RECU.LE.AL*.1) GO TO 8	1100
6	CONTINUE	1200
7	J1=1	1300
	R=.01	1400
	GO TO 9	1500
8	J1=L	1600
	R=(AL-1.)*.1	1700

9	J2=J1+1	1800
	IF(J1.EQ.1) GO TO 10	1900
	X=.10	2000
	GO TO 11	2100
10	X=.09	2200
11	Y=10.0	2300
	RAB1=R(I1,J1)+((RFGU-R)/X)*(R(I1,J2)-R(I1,J1))	2400
	RAB2=R(I2,J1)+((RFGU-R)/X)*(R(I2,J2)-R(I2,J1))	2500
	RAB=RAB1+((I2-10*(K-1))/Y)*(RAB2-RAB1)	2600
	RETURN	2700
	END	2800
 SUBROUTINE MODIFY		100
	INTEGER*4 RSTD	200
	COMMON /BACK/ BGINT(9,9),BKGR(9),RSTD(9),NBGS,RTITLE(20,9)	300
X	/STAND/ ESTD(10,10),STDINT(9,9,10),NESTD(10),	400
X	CS(10,10),STINTD(9,10,8),COMPST(9,10)	500
X	/UNKWN/ UINT(75,9,8),NFU,C(10)	600
3	/INFO/ F(25) /INOUT/ IOUT	700
4	/SCRIPT/ N,NU(8,3),NM,M,INDIC	800
5	/CORREC/ DRIFT(9,10,8),TAU(3),ISPEC(9)	900
	DEADT(X,Y)=X/(1-X*Y)	1000
	DO 4000 I=1,N	1100
	BKGR(I)=0.	1200
	IF (NBGS.EQ.0) GO TO 1000	1300
C	BACKGROUND CORRECTIONS	1400
	K=RSTD(I)	1500
	IF(K.EQ.0) GO TO 1000	1600
	BKGR(I)=BGINT(I,K)	1700
	WRITE(IOUT,1) (RTITLE(J,K),J=1,20),BKGR(I)	1800
1	FORMAT(' ',20A4,5X,E13.5,' COUNTS/SECOND')	1900
1000	JJ=ISPEC(I)	2000
	DO 2000 K=1,NM	2100
	NEST=NESTD(K)	2200
	DO 1100 J=1,NEST	2300
	IF(E(I).EQ.ESTD(J,K)) GO TO 1150	2400
1100	CONTINUE	2500
	GO TO 2000	2600
C	DRIFT CORRECTIONS	2700
1150	DO 1600 L=1,INDIC	2800
	DRIFT(I,K,L)=0.	2900
	L1=L+1	3000
	IF(ABS(STDINT(L1,I,K)-STDINT(L,I,K)).GE.6.*SQRT(STDINT(L,I,K)))	3100
	DRIFT(I,K,L)=(STDINT(L1,I,K)-STDINT(L,I,K))/NU(L,1)	3200
C	STANDARD INTENSITY CORRECTIONS	3300
1500	CONTINUE	3400
	STINTD(I,K,L)=DEADT(STDINT(L,I,K),TAU(JJ))-BKGR(I)	3500
1600	CONTINUE	3600
	COMPST(I,K)=CS(J,K)	3700
2	FORMAT(3I4,9E13.5)	3800
2000	CONTINUE	3900
C	DEAD TIME AND BACKGROUND CORRECTIONS TO UNKNOWN	4000
	DO 3500 L=1,INDIC	4100
	NUP=NU(L,1)	4200
	DO 3000 J=1,NUP	4300
3000	UINT(J,I,L)=DEADT(UINT(J,I,L),TAU(JJ))-BKGR(I)	4400
3500	CONTINUE	4500
4000	CONTINUE	4600
	RETURN	4700
	END	4800

	<b>SUBROUTINE OXYGEN (OX,C,A,N,IOX,COX)</b>	100
	DIMENSION OX(100),C(10),A(10),Z(10)	200
	CT=0.	300
	DO 100 I=1,N	400
100	CT=CT+C(I)/(OX(I)*A(I))	500
	C(IOX)=COX+16.*CT	600
	RETURN	700
	END	800
	 <b>SUBROUTINE TEST (KKK,LLL,ITEST)</b>	100
	REAL*4 KALPHA,KBETA,KEDGE,LALPHA,LEDFE	200
	COMMON /INFO/ SPACE(150),KALPHA(50),KBETA(50),KEDGE(50),	300
	LALPHA(50),LEDFE(50)	400
	DIMENSION ITEST(6)	500
	DO 100 I=1,6	600
100	ITEST(I)=0	700
	IF(KALPHA(KKK).LT.KEDGE(LLL)) ITEST(1)=1	800
	IF(KALPHA(KKK).LT.LEDFE(LLL)) ITEST(2)=1	900
	IF(KBETA (KKK).LT.KEDGE(LLL)) ITEST(3)=1	1000
	IF(KBETA (KKK).LT.LEDFE(LLL)) ITEST(4)=1	1100
	IF(LALPHA(KKK).LT.KEDGE(LLL)) ITEST(5)=1	1200
	IF(LALPHA(KKK).LT.LEDFE(LLL)) ITEST(6)=1	1300
	RETURN	1400
	END	1500
	 <b>BLOCK DATA</b>	100
	COMMON /RA/R(11,11)	200
	COMMON/SS/S(3)	300
	DATA R/1.00,.934,.856,.786,.735,.693,.662,.635,.611,.592,.578,	400
	A1.00,.944,.873,.808,.760,.718,.688,.663,.630,.613,.606,1.00,.953,	500
	B.888,.828,.782,.741,.713,.687,.665,.639,.634,1.00,.961,.903,.847,	600
	C.804,.764,.737,.713,.691,.665,.661,1.00,.968,.917,.867,.827,.789,	700
	D.764,.740,.718,.695,.691,1.00,.975,.933,.888,.851,.817,.793,.770,	800
	E.750,.730,.725,1.00,.981,.948,.911,.878,.847,.825,.805,.785,.767,	900
	F.763,1.00,.988,.963,.935,.907,.881,.862,.844,.826,.811,.806,1.00,	1000
	G.993,.977,.959,.938,.919,.904,.889,.874,.862,.858,1.00,.997,.990,	1100
	H.981,.970,.959,.950,.941,.932,.924,.921,11*1.00/	1200
	DATA S/'U','S','R'/	1300
	END	1400



## APPENDIX II

### Glossary of Terms Used in the Microprobe Correction Program

$A(I), 1 \leq I \leq M$	Atomic weight of the element, $E(I)$ .
$AK(I), 1 \leq I \leq N \leq M$	Ratio of the measured intensity readings for $E(I)$ in the unknown to the average corrected intensity reading of $E(I)$ in standard $K$ . (Equation 2).
$AK1(I), 1 \leq I \leq M$	Value of $k_A$ (calculated) at the $i$ th iteration; $AK1(I) = AAA * UC(I)/SC(I)$ .
AAA	Product of ZAF. This is calculated separately for each element $E(I)$ .
AB	Dummy argument for absorption correction in subroutine ABSCOR.
ABN	Absorption correction for the unknown. This is computed for each element $E(I)$ .
$ABDS(I, K), 1 \leq I \leq N$ $1 \leq K \leq NM$	Absorption correction for $E(I)$ in standard $K$ .
ABS	Combined absorption correction $ABS = ABN/ABDS(I, K)$ .
AT	Dummy argument for the atomic number correction in subroutine ATNCOR.
ATN	Atomic number correction for the unknown.
$ATDS(K), 1 \leq K \leq NM$	Atomic number correction for standard $K$ .
ANC	Combined atomic number correction $ANC = ATN/ATDS(K)$
$ABC\phi(I, J, K)$ $1 \leq I \leq M$ $1 \leq J \leq N$ $1 \leq K \leq 3$	Matrix of mass absorption coefficients with $E(I)$ the absorber, $E(J)$ the emitter, at the $K_\alpha$ -line ( $K = 1$ ), $K_\beta$ -line ( $K = 2$ ) or $L_\alpha$ -line ( $K = 3$ ).

BSTD(I), $1 \leq I \leq N$	If BSTD(I) = K, $1 \leq K \leq 9$ , then background standard K is used for E(I). If BSTD(I) = 0, no background correction is necessary.
BKGR(I), $1 \leq I \leq N$	Average background intensity for E(I).
BGINT(I, K) $1 \leq I \leq N$ $1 \leq K \leq \text{NBGS}$	Average background intensity readings for each element E(I) using background standard K in units of counts per second.
BTITLE(J, K), J=1,2 ... 20 $1 \leq K \leq \text{NBGS}$	Identification card for background standard K.
C(I), $1 \leq I \leq \text{NEU}$	Composition (wt%) of E(I) present in the unknown. C is used as a dummy argument denoting the composition of a sample in subroutines FLUCOR, ATNCOR, ABSCOR, OXYGEN
CK(I) CKL(I) CL1(I) CL2(I) CLM(I) $1 \leq I \leq M$ CM1(I) CM2(I) CM3(I) CMN(I)	Constants for computing mass absorption coefficients in subroutine ABCØ where $\mu/\rho = C\lambda^n$ .
CS(I,K), $1 \leq I \leq \text{NESTD}(K)$ $1 \leq K \leq \text{NM}$	Composition (wt%) of each element (ESTD(I,K)) present in standard K.
CØXY	Wt% oxygen present in the unknown which is given as input, CØX is dummy argument for CØXY in subroutine OXYGEN.
CØMPST(I, K), $1 \leq I \leq N$ $1 \leq K \leq \text{NM}$	Wt% of E(I) present in standard K.
CHI	$\chi$ - used in computing absorption correction (Equation 8, 9).
DRIFT(I,K,L), $1 \leq I \leq N$ $1 \leq K \leq \text{NM}$ $1 \leq L \leq \text{INDIC}-1$	Total Drift correction to intensity readings for E(I) in standard K. L refers to INDIC-1, the number of times the standards have been measured.

EL	Dummy argument used to specify chemical symbol of each element in a sample. EL is used in subroutines ABSCOR, FLUCOR and ATNCOR.
DR	Drift correction for unknown calculated for each element E(I).
DATA	NAMELIST name which inputs the following parameters: THETA, TAU, STD, N, NEU, MØDE, VØ, LINE, ISPEC, NPRØB, NSTCØM, NM, NBGS, BSTD.
E(I)	Array of chemical symbols of all elements used in the problem; this table is ordered as described in Appendix IV (E File).
ESTD(I,K) 1 ≤ I ≤ NESTD(K) 1 ≤ K ≤ NM	Array of chemical symbols for the elements E(I) present in the standard K.
FL	Dummy argument for fluorescence correction in subroutine FLUCOR.
FLN	Fluorescence correction for the unknown. If necessary, this is computed for each element E(I) in the unknown.
FLDS(I,K), 1 ≤ I ≤ N 1 ≤ K ≤ NM	Fluorescence correction (if any) for E(I) in standard K.
FLU	Combined fluorescence correction FLU = FLN/FLDS(I,K) FLU = 1.0, if no correction.
IN	Specifies which computer input unit is used.
IØUT	Specifies which computer output unit is used.
IØXI	Subscript used to denote whether oxygen is present in the problem. If IØ X 1 = 0, no oxygen If IØ X 1 > 0, oxygen is present

ISPEC(I), I = 1,N	Denotes which spectrometer was used to measure E(I).
INDIC	The number of times the standards as a group were measured (see Appendix IV, Data set D).
J1(I), I = 1,M	Mean Ionization Potential - Input to MAIN.
J	Mean Ionization Potential - as used in subroutine ATNCOR.
KEDGE (I), I = 1,M	K-edge for E(I)
KALPHA(I), I = 1,M	Wavelength of $K_{\alpha}$ -line for E(I).
KBETA(I), I = 1,M	Wavelength of $K_{\beta}$ -line for E(I).
LALPHA(I), I = 1,M	Wavelength of $L_{\alpha}$ -line for E(I).
M	The number of different elements used in any problem, $1 \leq M \leq 50$ . This includes elements in standards, and unknowns.
N	Number of elements in the unknown measured by the probe.
L1(I), I = 1,M L2(I), I = 1,M L3(I), I = 1,M	Wavelengths associated with the 3 L-edges, $L_I, L_{II}, L_{III}$ ; used in subroutine ABCØ to obtain mass absorption coefficients.
M1(I), M2(I), M3(I), M4(I), M5(I), I = 1,M	Wavelengths associated with the 5 M-edges, used in subroutine ABCØ to obtain mass absorption coefficients.
NK(I), I = 1,M	Exponent of $\lambda$ for each element (I) in the expression for mass absorption coefficient ( $\mu/\rho = C \lambda^n$ ) in the wavelength region from 0.7Å to the K-edge of element I.
NKL(I), I = 1,M	Exponent of $\lambda$ for each element (I) in the expression for mass absorption coefficient ( $\mu/\rho = C \lambda^n$ ) in the wavelength region from the K-edge to the $L_{III}$ -edge of element I.

NI(I), I = 1, M	Wavelength associated with the first N-edge; used in subroutine ABCØ to obtain mass absorption coefficients.
NEU	Total number of elements in the unknown.
NM	Number of standards used in a problem.
NU	Number of readings taken on an unknown sample. NU is obtained by the computer internally.
NC	The number of the unknown sample being analyzed. NC is obtained by the computer internally; $1 \leq NC \leq NU$ .
NBGS	Specifies the number of background standards used in the analysis.
NESTD(K) K = 1, NM	Number of elements in standard K
NEST	NEST = NESTD(K)
NSTCØM	See Explanation of input data set A (Appendix IV).
ØXY(I), I = 1, M	ØXY(I) = N/M if the oxide is of the form E(I) <sub>N</sub> O <sub>M</sub> .
ØMEGA(I,K), I = 1, M K = 1, 2	Fluorescence yield $\Omega$ (I, 1) for K radiation. Fluorescence yield $\Omega$ (I, 2) for L radiation.
RB	Dummy argument for backscatter coefficient in ATNCOR.
RBS(K), $1 \leq K \leq NM$	Backscatter coefficient for standard K.
RBØ	Backscatter coefficient for unknown.
RK(I), $1 \leq I \leq M$	Absorption jump ratio, r, for K lines in each element E(I).
R(I, J), I = 1, 11 J = 1, 11	Tabulated values of R. I indicates the atomic number and J indicates the reciprocal of $U = VC/VO$ .

RECU	Reciprocal of $U = VC/VO$ .
STDINT (J, I, K), J = 1, INDIC $1 \leq I \leq N$ $1 \leq K \leq NM$	Intensity of each element E(I) in standard K measured in counts per second. J gives the number of times the standards as a whole have been measured.
STINTD(I, K, L), $1 \leq I \leq N$ $1 \leq K \leq NM$ $1 \leq L \leq INDIC-1$	Intensity readings in counts per second for each element E(I) in standard K after correction for dead time and background.
STITLE(K, J) $1 \leq K \leq NM$ J = 1, 20	Identification card for standard K. See Appendix IV, Input file C.
SIGMA	$\sigma$ (Equation 11).
SB	Dummy argument for stopping power in ATNCOR.
SBS(K), $1 \leq K \leq NM$	Calculated stopping power for each standard K.
SBU	Calculated stopping power for each unknown.
STD(I, J) $1 \leq I \leq M$ $1 \leq J \leq NSTC\emptyset M$	See Appendix IV, input data set A.
STINT	The fully corrected intensity for each element in a standard.
SCON(I, K) $1 \leq I \leq M$ $1 \leq K \leq NM$	Weight % of element E(I) present in standard K.
TAU(J), J = 1,3	Dead time for spectrometers 1, 2, and 3 respectively.
TBG(K, I), K = 1, NBGS $1 \leq I \leq 3$	Counting time on background standard K. I indicates which group of elements is being measured.

TSTD(I, K, J)  
I = 1, INDIC  
K = 1, NM  
J = 1, 3

Counting time on standard K. I gives the number of times all of the standards have been measured. J indicates which group of elements is being measured.

THETA

Take off angle of the electron probe.

UINT(NC, J, L)  
 $1 \leq NC \leq NU(L)$   
 $1 \leq J \leq N$   
 $1 \leq L \leq INDIC-1$

Intensity reading for an element E(J) in counts per second in the unknown. NC indicates the number of the sample and L indicates the number of times the standards as a whole have been measured.

UTITLE (J),  
J = 1, 20

Identification card for the unknown.

UC(J),  $1 \leq J \leq N$

Calculated wt. % of element E(J) present in the unknown.

VO

Operating voltage.

VC(I, LL)  
 $1 \leq I \leq M$   
LL = 1, 2

Excitation energy of each element E(I) for either the K-line (LL = 1) or the L-line (LL = 2).

X

Dummy variable for CHI in subroutine FLUCOR and ABSCOR

Z(I),  $1 \leq I \leq M$

Atomic number of E(I).

### APPENDIX III

#### Detailed Descriptions of Subroutines

A. MAIN – This subroutine inputs the data files (A,B,C) and calls DATPTS to input data set (D). For those elements whose compositions are specified in the unknown, it is assumed that these compositions will not vary from unknown sample to unknown sample. MAIN calls for the calculation of mass absorption coefficients, and for each standard calls for the computation of atomic number, absorption and fluorescence corrections. To each element in the unknown MAIN assigns a standard and then calculates in turn an estimate for the composition of the unknown, the corrections associated with the estimated composition and then re-estimates the composition. This is done for each set of intensity readings on the unknown. If MODE = 2, the composition of the sample is known, and the intensity ratios of unknown to standard are calculated for each element. Finally, the MAIN routine outputs a detailed account of the computations to assist the experimenter in his analysis. Figure 1 presents a flow chart of MAIN.

B. DATPTS – DATPTS, as developed for the GSFC microprobe, reads and processes data in the following fashion: Each card, which DATPTS inputs, contains a time and 3 intensity counts, one for each spectrometer, as well as other information outlined for the Intensity Data File (Appendix IV).

If the readings are taken on the sample, the intensities are divided by the time to obtain the number counts per second for each measured element. The total number of data points taken on the sample are then computed. If the readings are taken on the standards or backgrounds, the average intensities and average counts per second are obtained.

When all the readings for the problem have been input, DATPTS calls subroutine MODIFY which computes corrections to the raw data. The flow chart of DATPTS is shown in Figure 2.

C. MODIFY – To each measured element MODIFY assigns a background correction according to the values of NBGS and the BSTD array (Appendix IV). It computes the drift correction for each element standard, storing the corrections in the DRIFT matrix. The count rates for each standard are corrected for dead-time and background and stored in the STINTD matrix. Then the unknown sample count rates are corrected for dead-time and background and restored in the UNT matrix. Figure 3 illustrates the flow of MODIFY.



D. ATNCOR – for each sample this subroutine computes the atomic number correction according to Equations 6-7. The (dummy) arguments of the subroutine are:

RB – the backscatter coefficient  
 SB – the stopping power  
 NEL – the number of elements in the sample  
 AT – the atomic number correction  
 EL – the element array for the sample.

Figure 4 gives a flow chart of ATNCOR.

E. ABSCOR – computes the absorption correction for each element in each sample (Equations 8-11). NEL, C, EL, AB are the arguments of the subroutine, ABSCOR. The first 3 have the same meaning as in ATNCOR and AB is the absorption correction for the element E(I). The MAIN program selects E(I) and transmits the parameter, I, in the ARGS COMMON block. Figure 5 illustrates the flow of ABSCOR.

F. FLUCOR – computes the fluorescence correction (if any) for each element E(L) in each sample (Equations 12-14). The arguments EL, NEL, C are the same as in ATNCOR; FL is the fluorescence correction and L is transmitted in the ARGS COMMON block. Figure 6 is a flow chart of FLUCOR.

G. TEST – is a subroutine called by FLUCOR to determine if an element, E(J), in a sample has fluoresced the element, E(L), measured for that sample. Whenever fluorescence occurs, the ITEST argument is set to 1; otherwise it remains 0:

If  $K_{\alpha} E(J) < K_{\text{edge}} E(L)$ , ITEST (1) = 1  
 $K_{\alpha} E(J) < L_{\text{edge}} E(L)$ , ITEST (2) = 1  
 $K_{\beta} E(J) < K_{\text{edge}} E(L)$ , ITEST (3) = 1  
 $K_{\beta} E(J) < K_{\text{edge}} E(L)$ , ITEST (4) = 1  
 $L_{\alpha} E(J) < K_{\text{edge}} E(L)$ , ITEST (5) = 1  
 $L_{\alpha} E(J) < L_{\text{edge}} E(L)$ , ITEST (6) = 1

J and L are also transmitted as arguments.

H. INTERP – linearly interpolates from the tabulated values for R given by Duncumb and Reed.<sup>4</sup> This subroutine is used by ATNCOR to compute the value of the backscatter coefficient. The arguments are IZ, the atomic number of E(I); RECU, (1/U); and RAB, the value of R. The values of  $R^4$  are contained in the block data at the end of MAIN.

I. ENDIT -- Tests for convergence for each measured element,  $E(I)$  in the unknown sample. If sufficient convergence is achieved, Equation 18, then ENDIT directs MAIN to end the iterative procedure and output the final composition of the unknown. If not, ENDIT returns to the iterative portion of MAIN.

## APPENDIX IV

### Description of Input Files

**A-DATA FILE** – The Fortran IV namelist feature directs the input of DATA File. This option enables data to be identified by name; for example,  $I = 1$  will cause a 1 to be placed in the location labeled I;  $E(3) = 'Fe'$  causes the third item in the E-array to be identified as iron. The parameters initialized in this fashion are described below. The statement in the MAIN program

```
100 READ (IN,DATA,END = 10000)
```

accomplishes the initialization. Figure 8 shows an example of the sample NAMELIST cards for one of the problems presented in Appendix VI.

M is the total number of elements present in both the unknown and standards

NEU is the total number of elements in the unknown

N is the number of elements measured by the probe

NM is the number of standards used in an analysis

NSTCOM is the number of ways in which the standards were combined to determine the composition or intensity ratios for the unknown. This allows different standards to be used for calculation of the same element. For example, if element A appears in both standard #1 and standard #2, the experimenter may wish to use standard #1 first for A to obtain the amount of A present and then standard #2 for the same purpose. Up to 10 such combinations may be made for a problem.

$STD(I,J)$ ,  $I = 1, N$ ;  $J = 1, NSTCOM$  specifies which standard is to be used for  $E(I)$  in standard combination J; for example if  $STD(4,2) = 3$ , use standard #3 for  $E(4)$  in the second analysis of the data. The standard is designated by number which corresponds to its order in the Composition File.

NBGS specifies the number of background standards used in the analysis:  
 $0 \leq NBGS \leq 9$ . This program assumes one background standard per element.

$BSTD(I)$ ,  $I = 1, N$  – If  $NBGS > 0$  there is at least 1 element measured for which the background is non-zero; hence, the BSTD array must be initialized for  $I = 1, 2, \dots, N$ , in the following manner. If the background for  $E(I)$  is zero, set  $BSTD(I) = 0$ ; if the background is not 0, set  $BSTD(I) = K$  where K is the number of the background standard used for  $E(I)$ .

MODE – If MODE = 1, intensity readings from the probe (Intensity Data File) are input and the weight % present in the unknown is determined for the measured elements. If MODE = 2 the composition of the unknown is known and the output consists of the intensity ratios of unknown intensity to standard intensity for each element in the unknown.

LINE(I), I = 1, M specifies whether the element, E(I), was measured with the K or L x-ray line.

ISPEC(I), I = 1, N specifies on which spectrometer (first, second, or third, etc.) the element, E(I), was measured.

VO gives the operating voltage for the probe ( $E_0$ ) in kilovolts.

TAU(I), I = 1, 2, 3 specifies the dead-times for spectrometers, 1,2,3 respectively.

Note: If the probe has more than 3 spectrometers, the dimension on TAU should be increased accordingly.

THETA is the take-off angle given in degrees.

#### B-Element File

For elements 3-92, file cards are provided for each element. Each of these sets of 3 cards contains the element's chemical symbol, atomic number (Z), atomic weight (A), excitation energies for the K (elements Z = 5-39) and L (elements Z = 12-92) lines in kilovolts (Bearden),<sup>18</sup> mean ionization potential ( $J_i$ ) in electron volts (Duncumb and Reed),<sup>4</sup> the K absorption jump ratio ( $r_A$ ) (Colby),<sup>19</sup> the cation to anion ratio if oxide is formed (OX), the K shell fluorescence yield (W(B)) and the mean L shell fluorescence yield (W(B)) (Fink, et al.).<sup>20</sup> the  $K_\alpha$  and  $K_\beta$  wavelengths for elements Z = 3-39 and the  $L_\alpha$  wavelengths for elements Z = 20-92. For elements less than atomic number 20, the  $L_\alpha$  wavelength is set to 50Å. Also given in the E file are the K,  $L_I$ ,  $L_{II}$ ,  $L_{III}$ ,  $M_I$ ,  $M_{II}$ ,  $M_{III}$ ,  $M_{IV}$ ,  $M_V$ ,  $N_I$  wavelength edges (Bearden)<sup>18</sup> in Å units and the CK, NK, CKL, NKL, CL1, CL2, CLM, CM1, CM2, CM3, CMN parameters used by Henrich<sup>6</sup> to calculate the mass absorption coefficients.

To enable the user to calculate approximate mass absorption coefficients for L lines of elements of atomic number 24-29, values of CL1 and CL2 were determined by linear extrapolation from the values given by Heinrich<sup>6</sup> for elements Z = 30 and above. The complete Element File compiled to date is given in the following Table.

TABLE I  
Element File

FL	NO.	WGT.-	VC(K)-	VC(L)-	-J--	-RK--	-OXY--	OMGAK	KALPHA	KBETA-	LALPHA	--CK---
'LI'	3.	007.0	0.0548	00.000	0000	00.00	1.0000	.0000	00.000	00.000	50.000	000.135
		2.88	226.5									
												.00375
'RE'	4.	009.0	00.111	00.000	0000	00.00	1.0000	.0000	00.000	00.000	50.000	000.350
		2.86	111.00									
												.00500
'R'	5.	010.8	00.000	00.000	0000	15.73	0.6667	.0000	67.600	00.000	50.000	000.740
		2.85										
												.00625
'C'	6.	012.0	00.284	00.000	0146	15.14	1.0000	.0009	44.700	00.000	50.000	001.350
		2.84	43.680									
												.00750
'N'	7.	014.0	00.400	00.000	0135	14.59	1.0000	0.002	31.600	00.000	50.000	002.210
		2.83	30.990									
												.00875
'O'	8.	015.0	00.532	00.000	0127	14.08	1.0000	0.004	23.700	00.000	50.000	003.800
		2.82	23.320									
												.01000
'F'	9.	019.0	00.692	00.000	0123	13.60	1.0000	0.007	18.300	00.000	50.000	004.900
		2.81										
												.01125
'NF'	10.	020.2	00.867	00.000	0123	13.16	1.0000	0.015	14.600	14.500	50.000	006.770
		2.80	14.300									
												.01250
'NA'	11.	023.0	01.070	00.000	0126	12.74	2.0000	0.020	11.910	11.617	50.000	009.050
		2.79	11.480	0.62	2.73							
												.01375
'MG'	12.	024.3	01.300	00.049	0133	12.35	1.0000	0.030	09.889	09.558	50.000	011.750
		2.79	09.512	000.89	2.73							
												.01500
'AL'	13.	027.0	01.560	00.075	0142	11.99	0.6667	0.042	08.337	07.981	50.000	014.870
		2.78	07.951	001.18	2.73							
												.01625
'SI'	14.	028.1	01.840	00.100	0154	11.64	0.5000	0.056	07.126	06.768	50.000	018.500
		2.77	06.738	001.54	2.73							
												.01750
'P'	15.	031.0	02.140	00.132	0166	11.31	0.4000	0.075	06.155	05.804	50.000	022.500
		2.76	05.784	002.43	2.73							
												.01875
'S'	16.	032.1	02.470	00.000	0180	11.00	1.0000	0.096	05.373	05.032	50.000	027.000
		2.76	05.018	002.43	2.73							
												.02000
'CL'	17.	035.5	02.820	00.000	0194	10.71	1.0000	0.117	04.728	04.403	50.000	031.700
		2.76	04.377	002.98	2.73							
												.02125
'AR'	18.	039.9	03.200	00.000	0209	10.43	1.0000	0.138	04.193	03.886	50.000	036.900
		2.75	03.871	003.62	2.73							
												.02250
'K'	19.	039.1	03.610	00.295	0224	10.17	2.0000	0.159	03.742	03.454	50.000	042.500
		2.75	03.437	004.31	2.73							
												.02375
'CA'	20.	040.1	04.040	00.349	0239	09.92	1.0000	0.180	03.359	03.090	36.330	048.400
		2.74	03.070	005.10	2.73							
												.02500
'SC'	21.	045.1	04.490	00.000	0255	09.68	1.0000	0.202	03.032	02.780	31.350	055.100
		2.74	02.7620005.99	2.73								
												.02625

TABLE I--(continued)

FL	NO.	WGT.	VC(K)	VC(L)	-J--	-RK--	-NXY--	OMGAK	KALPHA	KRETA-	LALPHA	--CK--			
					-NK-	KFNGF-	--CKL-	NKL-	--L1--	--CL1--	--L2--	--CL2--	--L3--	--CLM--	--M1--
					-CM1-	-M2--	CM2-	--M3--	CM3-	--M4--	--M5--	CMN--	--N1--	OMGAL	
'TI'	22.	047.9	04.970	00.450	0270	09.45	1.0000	0.224	02.748	02.514	27.420	062.100			
		2.73	02.497	007.00	2.73						27.3				
											.02750				
'V'	23.	051.0	05.470	00.000	0286	09.24	1.0000	0.246	02.504	02.284	24.250	069.800			
		2.73	02.269	008.02	2.73										
											.02875				
'CR'	24.	052.0	05.990	00.598	0301	09.03	1.0000	0.268	02.291	02.085	21.640	078.000			
		2.73	02.070	009.18	2.73	16.7	7.80	17.9	4.8	20.7					
											.03000				
'MN'	25.	054.9	06.540	00.000	0316	08.83	1.0000	0.290	02.103	01.910	19.500	086.700			
		2.72	01.896	010.45	2.73	8.85		5.9	19.1						
											.03125				
'FF'	26.	055.9	07.110	00.707	0332	08.64	1.0000	0.320	01.937	01.757	17.600	095.800			
		2.72	01.743	011.75	2.73	17.2	10.0	17.2	7.0	17.525					
											.03250				
'CN'	27.	058.9	07.710	00.770	0347	08.46	1.0000	0.350	01.790	01.621	16.000	105.500			
		2.71	01.608	013.25	2.73	15.61	11.25	15.61	8.0	15.92					
											.03375				
'NI'	28.	058.9	08.330	00.854	0362	08.29	1.0000	0.380	01.659	01.500	14.600	115.900			
		2.71	01.488	014.80	2.73	14.2	12.6	14.2	9.0	14.52					
											.03500				
'CI'	29.	063.5	08.980	00.932	0377	08.12	1.0000	0.410	01.542	01.392	13.400	126.800			
		2.71	01.380	016.45	2.73	11.270	014.06	13.01	10.1	13.29					
											.03625				
'ZN'	30.	065.4	09.660	01.020	0392	07.96	1.0000	0.440	01.436	01.295	12.300	138.000			
		2.70	01.283	018.25	2.73	10.330	015.60	11.870	011.20	12.13					
											.03750				
'GA'	31.	069.7	10.400	01.120	0407	07.81	1.0000	0.480	01.341	01.208	11.313	149.800			
		2.70	01.196	020.20	2.73	09.520	017.25	10.830	12.40	11.100					
											.03875				
'GF'	32.	072.6	11.100	01.220	0422	07.66	1.0000	0.520	01.255	01.128	10.456	162.200			
		2.70	01.117	022.15	2.73	08.730	018.95	09.940	013.60	10.190					
											.04000				
'AS'	33.	074.9	11.900	01.320	0437	07.51	1.0000	0.560	01.177	01.056	09.671	175.400			
		2.69	01.045	024.25	2.73	08.107	20.750	09.124	014.90	09.390	004.28				
											.04125				
'SF'	34.	079.0	12.600	01.430	0451	07.38	1.0000	0.600	01.106	00.990	08.990	184.400			
		2.69	00.980	026.40	2.73	07.506	22.550	08.416	16.200	08.670	004.69				
											.04250				
'BR'	35.	079.9	13.500	01.550	0466	07.24	1.0000	0.640	01.041	00.931	08.375	205.000			
		2.69	00.920	028.80	2.73	06.970	24.600	007.80	17.700	008.00	005.13				
											.04375				
'KR'	36.	083.7	14.300	01.680	0481	07.12	1.0000	0.666	00.981	00.876	00.000	219.300			
		2.68	00.866	031.25	2.73	6.4600	26.700	007.21	019.15	007.43	005.60				
											.04500				
'RR'	37.	085.5	15.200	01.810	0495	00.00	1.0000	0.692	00.927	00.827	07.318	235.500			
		2.68	00.816	033.90	2.73	05.998	028.80	06.643	020.80	004.89	006.10				
											.04625				
'SR'	38.	087.6	16.110	01.940	0510	00.00	1.0000	0.718	00.876	00.781	06.863	251.300			
		2.68	00.770	036.50	2.73	05.583	031.20	06.172	022.40	06.387	006.62				
											.04750				
'Y'	39.	088.9	17.040	02.080	0524	00.001	0.0000	0.744	00.830	00.738	06.449	268.100			
		2.67	00.728	039.30	2.73	05.232	033.50	05.755	024.10	05.961	007.18				
											.04875				
'ZR'	40.	091.2	00.000	02.220	0538	00.00	0.5000	0.770	00.787	00.700	06.070				
											.05000				
											.05724				
'NR'	41.	092.9	00.000	02.370	0553	00.00	1.0000	0.796	00.746						
											.05600				
											.05600				

TABLE I-(continued)

FL	NO.	WGT.-	VC(K)-	VC(L)-	-J--	-RK--	-OXY--	OMGAK	KALPHA	KBETA-	LALPHA	--CK---	
			-NK-	KFDGF-	--CKL-	NKL-	--L1--	-CL1--	--L2--	-CL2--	--L3--	-CLM--	--M1--
			-CM1-	-M2---	CM2-	--M3--	CM3-	--M4--	--M5--	CMN--	--N1--	OMGAI.	
'MO'	42.	096.0	00.000	02.520	0567	00.00	1.0000	0.822	00.710			05.406	
			048.50	2.72	04.298	041.45	04.718	029.73	04.913	009.05			
												.062	
'TC'	43.	099.0	00.000	02.680	0581	00.00	1.0000	0.848				00.000	
			052.10	2.72	04.060	044.50	04.436	031.95	04.632	009.74			
												.068	
'RH'	44.	101.7	00.000	02.840	0595	00.00	1.0000	0.874				04.846	
			055.60	2.72	003.83	047.50	04.180	034.10	04.369	010.47			
												.074	
'RH'	45.	102.9	00.000	03.000	0609	00.00	1.0000	0.900				04.597	
			059.30	2.72	03.626	050.70	03.942	036.40	04.130	011.24			
												.080	
'PD'	46.	106.4	00.000	03.170	0623	00.00	1.0000	0.926				04.368	
			063.10	2.72	03.428	053.95	03.724	038.70	03.908	012.03			
												.086	
'AC'	47.	107.9	00.000	03.350	0637	00.00	2.0000	0.952				04.156	
			067.30	2.71	03.254	057.50	03.514	041.30	03.698	012.87			
												.092	
'GD'	48.	112.4	00.000	03.540	0651	00.00	1.0000	0.978				03.956	
			071.00	2.71	03.085	060.70	03.326	043.60	03.504	013.76			
												.098	
'JM'	49.	114.8	00.000	03.730	0665	00.00	1.0000	1.004				03.772	
			075.50	2.71	02.926	064.50	03.147	046.30	03.324	014.68			
												.104	
'SN'	50.	118.7	00.000	03.930	0679	00.00	1.0000	1.030				03.600	
			079.60	2.71	02.777	068.00	02.982	048.80	03.156	015.65			
												.110	
'SR'	51.	121.8	00.000	04.130	0692	00.00	0.6667	1.056				03.439	
			084.20	2.70	02.639	071.95	02.830	051.70	3.0000	016.66			
												.116	
'TF'	52.	127.6	00.000	04.340	0706	00.00	1.0000	1.082				03.289	
			088.50	2.70	02.511	075.65	02.687	054.30	02.855	017.72			
												.122	
'J'	53.	126.9	00.000	04.560	0720	00.00	1.0000	1.108				03.148	
			094.00	2.70	02.389	080.35	02.553	057.70	02.719	018.82			
												.128	
'XF'	54.	131.3	00.000	04.780	0734	00.00	1.0000	1.134				00.000	
			098.30	2.69	02.274	084.00	02.429	060.30	02.592	019.97			
												.134	
'CS'	55.	132.9	00.000	05.010	0747	00.00	1.0000	1.160				02.892	
			103.40	2.69	02.167	088.40	02.314	063.40	02.474	021.16	10.202		
			018.2									.140	
'RA'	56.	137.4	00.000	05.250	0761	00.00	1.0000	1.186				02.775	
			109.00	2.69	02.068	093.15	02.204	066.90	02.363	022.41	09.557		
			019.3	10.845	16.0							.146	
'LA'	57.	138.9	00.000	05.480	0775	00.00	1.0000	1.212				02.665	
			114.40	2.68	01.973	097.80	02.103	070.20	02.258	023.70	09.042		
			020.4	10.321	16.9							.152	
'CF'	58.	140.1	00.000	05.72	0788	00.00	0.6667	1.238				02.561	
			120.00	2.68	01.889	102.50	02.011	073.60	02.164	025.05	08.614		
			021.6	09.692	17.9	10.408	15.5					.158	
'PR'	59.	140.9	00.000	05.960	0802	00.00	1.0000	1.264				02.463	
			125.50	2.68	01.811	107.30	01.924	77.000	02.077	026.47	08.188		
			022.8	09.258	18.9	09.957	16.3					.164	
'ND'	60.	144.3	00.000	06.210	0815	00.00	1.0000	1.290				02.370	
			132.00	2.67	01.735	112.80	01.843	81.000	01.995	027.93	07.831		
			024.1	08.773	20.0	09.499	17.2					.170	
'PM'	61.	145.0	00.000	06.460	0829	00.00	1.0000	1.316				02.283	
			138.50	2.67	01.665	118.50	01.767	085.00	01.918	029.44	07.513		
			025.4	08.376	21.0	09.115	18.2					.176	

TABLE I-(continued)

FL NO. WGT.- VC(K)- VC(L)- -J-- -RK-- -OXY-- OMGAK KALPHA KRFTA- LALPHA --CK---  
 -NK- KFDGE- --CKL- NKL- --L1-- -CL1-- --L2-- -CL2-- --L3-- -CLM-- --M1--  
 -CM1- -M2-- -CM2- --M3-- CM3- --M4-- --M5-- CMN-- --N1-- OMGAL

'SN'62. 150.4 00.000 06.720 0843 00.00 1.0000 1.342 02.199  
 143.50 2.67 01.599 122.50 01.703 088.00 01.845 031.01 07.178  
 026.7 08.023 22.2 08.705 19.1 .182  
 'EU'63. 152.0 00.000 06.980 0856 00.00 1.0000 1.368 02.121  
 150.00 2.66 01.536 128.00 01.626 092.00 01.775 032.63 06.856  
 028.1 07.642 23.3 08.331 20.1 10.604 10.893 13.85 .188  
 'GD'64. 156.9 00.000 07.240 0870 00.00 1.0000 1.394 02.046  
 157.00 2.66 01.477 134.00 01.561 096.30 01.710 034.31 06.566  
 029.6 07.322 24.5 07.997 21.2 10.111 10.408 14.30 .194  
 'TR'65. 158.9 00.000 07.520 0883 00.00 1.0000 1.420 01.975  
 163.50 2.65 01.421 140.00 01.501 100.30 01.649 036.06 06.292  
 031.1 07.011 25.8 07.685 22.3 09.692 09.957 14.80 .200  
 'DY'66. 162.5 00.000 07.790 0897 00.00 1.0000 1.446 01.909  
 170.50 2.65 01.365 146.00 01.438 104.50 01.579 037.86 06.088  
 032.6 06.715 27.0 07.383 23.4 09.258 09.528 15.30 .206  
 'HN'67. 164.9 00.000 08.070 0910 00.00 1.0000 1.472 01.845  
 176.50 2.65 01.317 151.00 01.390 108.50 01.535 039.76 05.820  
 034.3 06.453 28.5 07.128 24.5 08.899 09.155 15.85 .212  
 'ER'68. 167.2 00.000 08.360 0923 00.00 1.0000 1.498 01.784  
 184.00 2.64 01.268 157.00 01.33, 113.00 01.482 041.69 05.581  
 035.9 06.179 29.8 06.834 25.7 08.508 08.773 16.30 .218  
 'TM'69. 168.9 00.000 08.650 0937 00.00 1.0000 1.524 01.726  
 192.50 2.64 01.222 165.00 01.288 118.00 01.433 043.69 05.366  
 037.7 05.931 31.2 06.559 27.0 08.155 08.433 16.85 .224  
 'YR'70. 173.0 00.000 08.940 0950 00.00 1.0000 1.550 01.672  
 199.00 2.63 01.182 170.00 01.243 122.00 01.386 045.79 05.161  
 039.6 05.686 32.7 06.331 28.3 07.836 08.086 17.45 .230  
 'LI'71. 175.0 00.000 09.250 0964 00.00 1.0000 1.576 01.619  
 206.00 2.63 1.1400 176.00 01.199 126.00 01.341 047.92 04.972  
 041.3 05.475 34.2 06.112 29.6 07.545 07.777 18.00 .245  
 'HF'72. 178.6 00.000 09.560 0997 00.00 1.0000 1.602 01.570  
 214.00 2.62 01.100 183.00 01.155 131.00 01.297 050.12 04.753  
 043.2 05.211 35.8 05.861 30.9 07.190 07.427 18.55 .260  
 'TA'73. 180.9 00.000 09.880 0991 00.00 1.0000 1.628 01.522  
 222.00 2.62 01.061 190.00 01.114 136.00 01.255 052.44 04.569  
 045.2 05.015 37.5 05.637 32.4 06.898 07.128 19.05 .275  
 'W'74. 183.9 00.000 10.200 1004 00.00 1.0000 1.654 01.476  
 231.00 2.61 01.025 197.00 01.075 142.00 01.216 054.79 04.405  
 047.2 04.823 39.1 05.451 33.8 06.640 06.871 19.60 .290  
 'RF'75. 186.2 00.000 10.500 1017 00.00 1.0000 1.680  
 239.00 2.61 00.990 204.00 01.037 147.00 01.177 057.21 04.231  
 049.3 04.625 40.9 05.253 35.3 06.357 06.594 20.10 .305  
 'DS'76. 190.2 00.000 10.900 1031 00.00 1.0000 1.706 01.391  
 247.00 2.60 00.956 211.00 01.001 152.00 01.140 059.70 04.064  
 051.5 04.443 42.6 05.060 36.9 06.106 06.357 20.45 .320  
 'IR'77. 193.1 00.000 11.200 1044 00.00 1.0000 1.732 01.351  
 256.00 2.60 00.923 219.00 00.967 157.00 01.106 062.27 03.915  
 053.7 04.273 44.5 04.873 38.4 05.880 06.094 21.35 .335  
 'PT'78. 195.1 00.000 11.600 1057 00.00 1.0000 1.758 01.313  
 263.00 2.59 00.893 225.00 00.934 161.00 01.072 064.92 03.774  
 056.0 04.114 46.4 04.710 40.1 05.665 05.880 21.95 .350  
 'AU'79. 197.2 00.000 11.900 1071 00.00 1.0000 1.784 01.276  
 272.00 2.59 00.863 232.00 00.903 167.00 01.040 067.64 03.620  
 058.3 03.939 48.3 04.522 41.8 05.415 05.629 22.50 .365  
 'HG'80. 200.6 00.000 12.300 1084 00.00 1.0000 1.810 01.241  
 281.00 2.58 00.835 240.00 00.872 172.00 01.008 070.55 03.482  
 060.8 03.779 50.4 04.349 43.6 05.280 05.413 23.10 .380  
 'TL'81. 204.4 00.000 12.600 1097 00.00 1.0000 1.836 01.207  
 289.00 2.58 00.808 247.00 00.843 177.00 00.979 073.60 03.349  
 063.5 03.632 52.6 04.201 45.4 04.998 05.200 23.80 .385



TABLE I--(Continued)

```

FL  NO. WGT.- VC(K)- VC(L)- -J-- -RK-- -OXY-- OMGAK KALPHA KBETA- LALPHA --CK---
-NK- KEDGE- --CKL- NKL- --L1-- -CL1-- --L2-- -CL2-- --L3-- -CLM-- --M1--
-CM1- -M2--- CM2- --M3-- CM3- --M4-- --M5-- CMN-- --N1-- OMGAL

'PB'82. 207.2 00.000 13.000 1111 00.00 1.0000 1.862 . 01.175
          298.00 2.57 00.782 255.00 00.815 183.00 00.950 076.64 03.218
          066.1 03.484 54.7 04.042 47.3 04.797 04.994 24.40 .390
'AI'83. 209.0 00.000 13.400 1124 00.00 0.6667 1.888 01.144
          307.00 2.56 00.757 262.00 00.789 188.00 00.923 079.94 03.099
          068.9 03.355 57.1 03.902 49.3 04.612 04.808 25.10 .395
'TH'90. 232.1 00.000 16.300 1217 00.00 1.0000 2.070 00.956
          094.3 02.577 78.1 03.080 67.5 03.573 03.745 31.30 9.492 .430
          00.761 109.40 02.401
'II'92. 238.1 00.000 17.200 1244 00.00 1.0000 2.122 00.911
          00.722 120.70 02.235
          104.0 02.394 86.2 02.882 74.5 03.300 03.496 33.60 8.614 .440

```

The order in which the Element File is used as input is most important. If we run in MODE = 1, the order of the element cards as submitted to the computer, must be the same as the order in which the probe measured the elements; for example, if data set 1 consisted of elements E<sub>1</sub>, E<sub>2</sub>, E<sub>3</sub>, measured on spectrometers 1, 2, and 3, respectively, and set 2 consisted of E<sub>4</sub>, E<sub>5</sub>, measured on spectrometers 1 and 2 respectively, then the order of the element cards must be E<sub>1</sub>, E<sub>2</sub>, E<sub>3</sub>, E<sub>4</sub>, E<sub>5</sub>. If other elements are present in the unknown and their weight percentages are known, their element cards come next. If oxygen is present in the unknown, its element cards are next. The last elements considered in the Element File are the file cards of elements present in the standards but not in the unknown. The element file is input following the DATA File. Figure 9 shows data for Input Files A and B for one of the sample problems (Appendix VI).

### C-Composition Files

The composition file consists of the following types of data:

(1) Title card for the unknown sample. On this card may be punched up to 80 characters of alphanumeric information to aid the experimenter in identifying his problem. This information is stored in the UTITLE array. The statement in MAIN;

Read (IN,40) (UTITLE(I), I = 1,20) accomplishes the data transmittal.

(2) Composition card(s) for the unknown sample. The composition of each element in the unknown (wt. %), if known before the analysis, is input here. If it is not known it is set to zero. For example in columns 1-10 is punched the weight % of E(1) present in the unknown; in columns 11-20, the weight % of

E(2), . . . , in columns 71-80, the weight % of E(8). On the second card in columns 1-10 goes the weight % of E(9), and so forth, until the weight percentage of each of the NEU elements in the unknown, if available before electron probe analysis, has been entered ( $1 \leq \text{NEU} \leq 10$ ). Note: The order of the elements in the unknown should be the same as in the E list. If none of the sample compositions is known, they are all set to zero. The statement in MAIN

```
READ (IN,2) (C(I), I = 1, NEU)
```

2 FORMAT (8F10.2), accomplishes the data transmittal.

(3) Cards for the Standards. These cards are input using the following statements:

```
DO 600 K = 1, NM
```

```
READ (IN,40) (STITLE(I,K), I = 1,20)
```

```
READ (IN,41) NEST, (ESTD(I,K), I = 1, NEST)
```

```
READ (IN,2) (CS(I,K), I = 1, NEST)
```

```
600 CONTINUE
```

The first READ inputs the title card for standard #K. The second READ inputs the second card which contains the number of elements in standard #K and the chemical symbols of each of these elements. On this card columns 1-4 contain the number of elements, 5-8, the name of ESTD (1,K), 9-12, the name of ESTD (2,K), . . . ., 41-44, the name of ESTD (10,K). All entries on this card must be right-adjusted. The third READ statement inputs the next cards which give the wt. % of ESTD (1, K), . . . ., ESTD(NEST,K), respectively. The same format (8F10.2), as that used to input any known compositions in the unknown sample, is also used here. The order of the standards,  $K = 1, 2, \dots, \text{NM}$ , must be the order in which their intensities were measured. Thus, if STANDARD #12 was measured first, then information regarding it will be found whenever  $K = 1$ ; if STANDARD #10 read second, then information regarding it stored for  $K = 2$ .

(4) Cards for the backgrounds. If NBGS > 0, then for each background standard a title card of identifying information is input. This is accomplished by the statements in MAIN:

```
DO 800 K = 1, NBGS
```

```
800 READ (IN,40) (BTITLE(I,K), K = 1,20).
```

Figure 10 shows data from the Composition File as used in the sample problem illustrated in Appendix VI.

#### D-Intensity Data File

Whenever the program is operating in MODE 1, subroutine DATPTS inputs D. The following procedure may have to be changed according to the type of output data obtained from each particular electron probe. The type of identification scheme used at GSFC is outlined here. To change the scheme, DATPTS, Appendix III, may have to be reprogrammed.

Each probe measurement is recorded on a punched card, which is divided into 10 fields. These fields contain

- (1) the x-ray counting time (TIME)
- (2-4) the intensity counts for each spectrometer (XINT(I)), I = 1,2,3)
- (5) the sample-type code (SAME) for unknown, standard, or background,
- (6) the set-number (NUMRDG), which tells whether element group 1 (E(1) - E(3)), 2 (E(4)-E(6)) or 3 (E(7)-E(9)) is being measured
- (7) the group - number (INDIC) which tells the number of times the standards, as a group, have been measured. INDIC is used so that the drift calculation can be made several times during a run
- (8) the number or identification of the standard or background being measured (NOST). As discussed in the description of the Composition File, the background and standards are remembered according to the order in which their intensities were measured (1,2,3. . . .). Therefore this re-alignment is used for NOST
- (9) the dummy variable (IUNK) for unspecified usage,
- (10) the end of data card (LAST). This field is used on the last data card of a problem to indicate the end of the data. Its usage allows a second problem to be run.

The statements:

```
READ(IN,1) TIME, (XINT(J), J = 1,3), SAME, NUMRDG, INDIC, NOST, IUNK, LAST  
1 FORMAT (F6.0, 3F8.0, 4I1, I2,I1)
```

input one card. The data are then processed according to the value of SAME. Table II presents a capsule review of the use of the 10 fields with some additional comments.

**TABLE II**  
**Format for Intensity Data File**

Field	Columns	Format	Parameter	Comment
1	1-6	F6.0	TIME	duration of measurement
2	7-14	F8.0	XINT(1)	intensity in counts from spectrometer #1
3	15-22	F8.0	XINT(2)	intensity from spectrometer #2
4	23-30	F8.0	XINT(3)	intensity from spectrometer #3
5	31	I1	SAME	SAME = 1 unknown measured SAME = 2 standard measured SAME = 3 background measured
6	32	I1	NUMRDG	NUMRDG = 1 measurements taken on E(1),E(2),E(3) (Set 1) NUMRDG = 2 measurements taken on E(4),E(5),E(6) (Set 2) NUMRDG = 3 measurements taken on E(7),E(8),E(9) (Set 3)
7	33	I1	INDIC	INDIC = i, i = 1,2,.. . . , 9 ith time readings are taken on standards
8	34	I1	NOST	NOST = i, i = 1,2,.. . . , NM measurement taken on the ith standard
9	35-36	I2	IUNK	IUNK = i, i = 1, 2, . . . . , 99 Dummy variable for other usage
10	37	I1	LAST	LAST = 0 except on last data card where LAST = 1. This indicates the end of the data for the problem.

## APPENDIX V

### Output from the Program

The format of the data output and the FORTRAN statements which control this output are given below. The FORTRAN statements are referred to by their sequence numbers (SN) which appear to the right of each statement in the program listing (Appendix I).

Data set B, that is the E file, and selected parameters from data set A are printed out by MAIN, SN 15200: 19700. The matrix of mass absorption coefficients is calculated in SUBROUTINE ABSCO, SN 6500: 8100. The raw probe data are output in SUBROUTINE DATPTS at SN 4700: 4900, SN 6200: 6300, SN 7100: 7200, according to whether the measurement was of an unknown, a standard, or a background, respectively. For each standard, the title and composition cards are printed by MAIN, SN 24100: 24300; and the atomic number correction is also printed by MAIN, SN 24400: 24400. For each element in each standard, the absorption correction is printed by MAIN, SN 25700: 25700 and the fluorescence correction is output in FLUCOR, SN 5200: 5300, SN 5800: 6200.

Whenever a new standard combination is used, a message is written by MAIN, SN 27500: 27500 telling which standard is used for each measured element in the unknown. The sample compositions are printed at each iteration step: MAIN, SN 32300: 32600 only for the first set of intensity data. The weight % sum is also printed out. No attempt is made to normalize the sum to 100%.

After the final iteration on the first data point in a MODE 1 problem and for MODE 2 problems, the atomic number corrections, the absorption corrections and the fluorescence corrections are printed by statements MAIN, SN 31000: 31600, FLUCOR, SN 5200: 5300, FLUCOR, SN 5800: 6200 respectively. For subsequent data points the compositions of the first and last iterations are written, MAIN 29200 and 32600. For MODE 2 problems intensity ratios are printed by MAIN, SN 33300: 33300. Sample output data are given in Appendix VI, where several sample problems are considered. The output format can be easily changed by altering the FORTRAN statements previously indicated.

## APPENDIX VI Sample Problems

Three sample problems will be described in order to illustrate the use of this program. A copy of the input data (Files A-D) and the output for each of the problems is supplied as a guide for the successful application of the program.

Problem 1 – (Geological Application)\* A sample is analyzed in MODE 1 for the following 8 elements: Ca, Mg, Si, Al, Na, Mn, Cr, and Fe. Oxygen is also present as is 0.1 wt % Ti. Six standards are used: pure diopside for Ca, Mg and Si, and 5 other standards, one for Al, Na, Mn, Cr and Fe. The data were taken on an electron probe with 3 spectrometers and a takeoff angle of 52.5°. The operating voltage was 20 KV; only K lines were measured. Input Files A-C have already been illustrated (Figures 8-10). Input File D, the Intensity Data File, is shown in Figure 11. A shortened version of the Output Data is given in Figure 12.

The initial estimate of the unknown composition using a linear relation (Equation 15) is given in Table III as well as the final calculated compositions determined after 3 iterations.

TABLE III  
Calculated Compositions Problem 1

Element	Initial Estimate (wt. %)	Final calculated composition (wt. %)
Ca	11.54	11.57
Mg	12.55	12.87
Si	25.13	26.1
Al	1.15	1.08
Na	0.88	0.93
Mn	0.09	0.09
Cr	0.59	0.58
Fe	2.57	2.59
Ti(known)	0.10	0.10
O	43.78	45.0
Sum	98.4	100.9

For all elements measured in this problem, the only correction factor of any significance was the absorption correction.

\*This problem was kindly loaded to us by Dr. J. Boyd of the Geophysical Lab., Washington, D.C.

**Problem 2 – (Metallurgical Application)\*** A binary sample is analyzed in MODE 1 for Ti and Nb only. Pure standards of the elements are used. The data was taken on an electron probe with 2 spectrometers and a takeoff angle of 52.5°. The operating voltage was 20 KV and Ti-K and Nb-L radiations were measured. Input Files A-D are illustrated in Figure 13 and the Output Data is given in Figure 14.

The initial estimate of the unknown composition using a linear relation (Equation 15) is given in Table IV as well as the final calculated composition determined after 3 iterations.

TABLE IV  
Calculated Composition Problem 2

Element	Initial Estimates (wt. %)	Final Calculated Composition (wt. %)
Ti	32.7	34.8
Nb	62.2	65.5
Sum	94.9	100.3

For Ti, an atomic number correction of +7.6% is necessary while for Nb a negative atomic number correction of 5% is calculated. A significant absorption correction (~13%) was calculated for the Ti. The absorption correction for Nb was insignificant.

**Problem 3 – (Metallurgical Application)** A sample containing 4 elements has the following composition Al-10 wt. %, Mg-50 wt. %, Au-40 wt. %, Cu-0.10 wt. %. It is of interest to calculate using MODE 2 the expected concentration ratios ( $I/I(A)$ ) for the 4 elements if pure standards of each element are used. The data is calculated for an electron probe with a takeoff angle of 52.5° and an operating voltage of 20 KV. The following x-ray lines were considered Al-K, Mg-K, Au-L and Cu-K. Input Files A-C are illustrated in Figure 15 and a shortened version of the Output Data is given in Figure 16.

The calculated intensity ratios are given in Table V.

This particular combination of elements exhibits a large atomic number effect (up to 30%) and a large absorption effect (up to 50%). Only Cu is noticeably fluoresced, namely by the Au-L line. Large corrections of this order, compounded with  $f(x)$  values less than 0.8 for Al and Mg, argue quite strongly for the use of intermediate, 4 element, standards.

---

\*This problem was kindly loaned to us by D. Beaman of the Dow Chemical Company, Midland, Michigan.

**TABLE V**  
**Calculated Intensity Ratios, Problem 3**

Eloment	Composition (wt. %)	ANC = ATN/ATD	ABS = ABN/ABD	FLU = FLN/FLD	Intensity Ratio, Sample to Standard = ANC·ABS·FLU·C(U)
Al	10	1.14	0.48	1.0	$5.5 \times 10^{-2}$
Mg	50	1.17	0.69	1.0	$41. \times 10^{-2}$
Au	40	0.71	1.01	1.0	$29. \times 10^{-2}$
CU	0.1	0.95	0.98	1.07	$.101 \times 10^{-2}$



## **APPENDIX VII**

### **Figures**

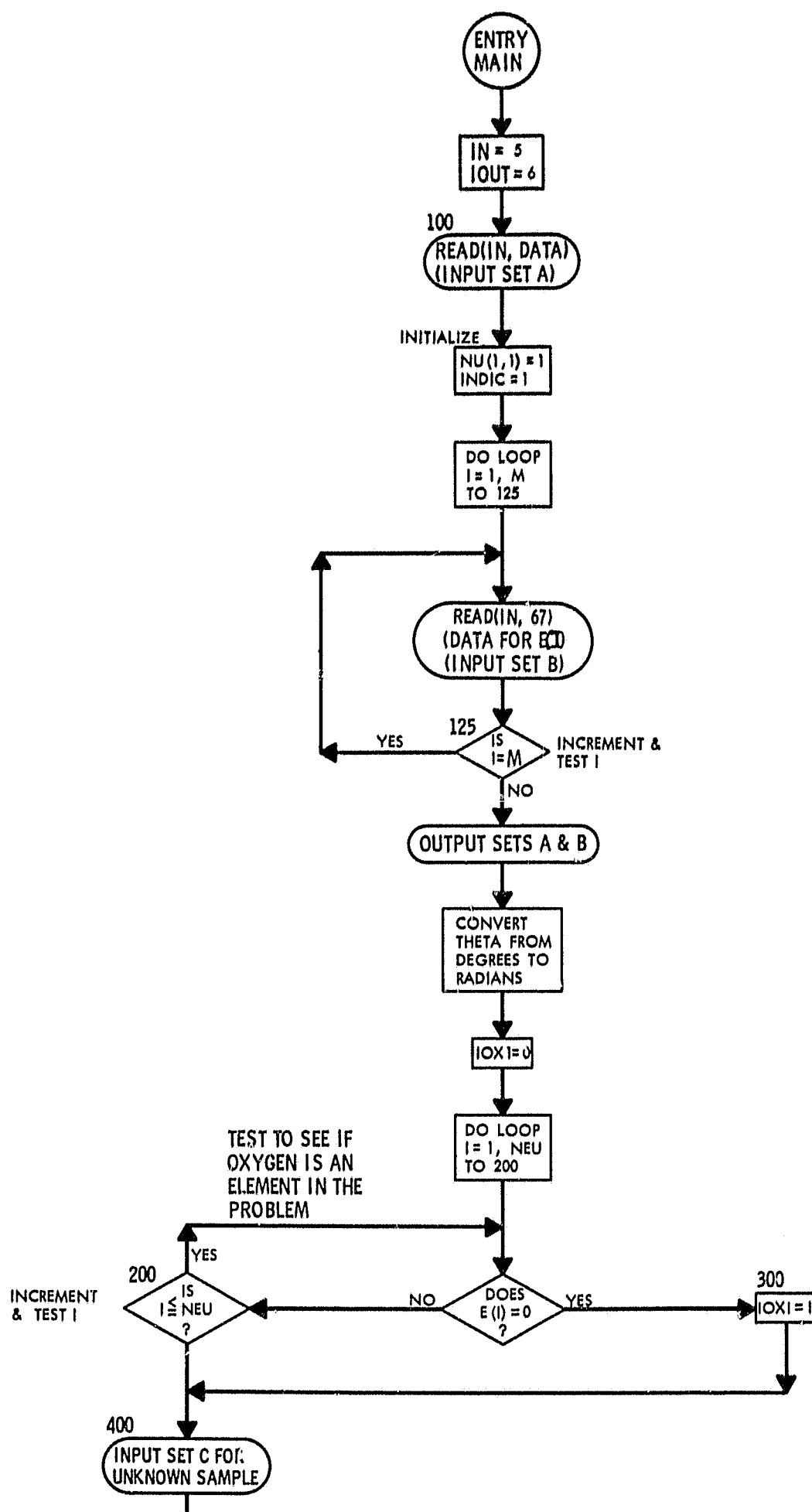


FIGURE 1 : MAIN PROGRAM

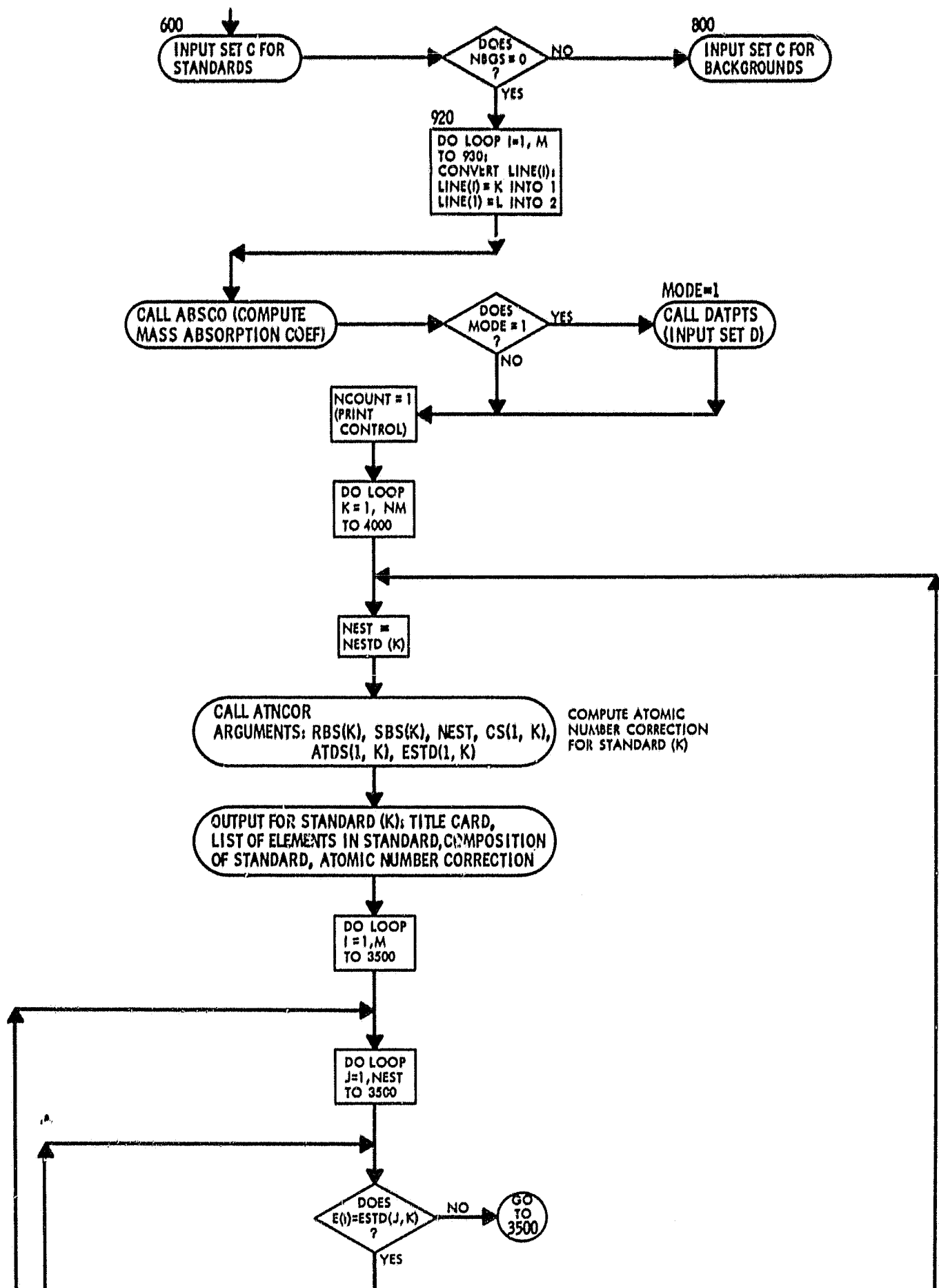


FIGURE 1 : MAIN PROGRAM (CONTINUED)

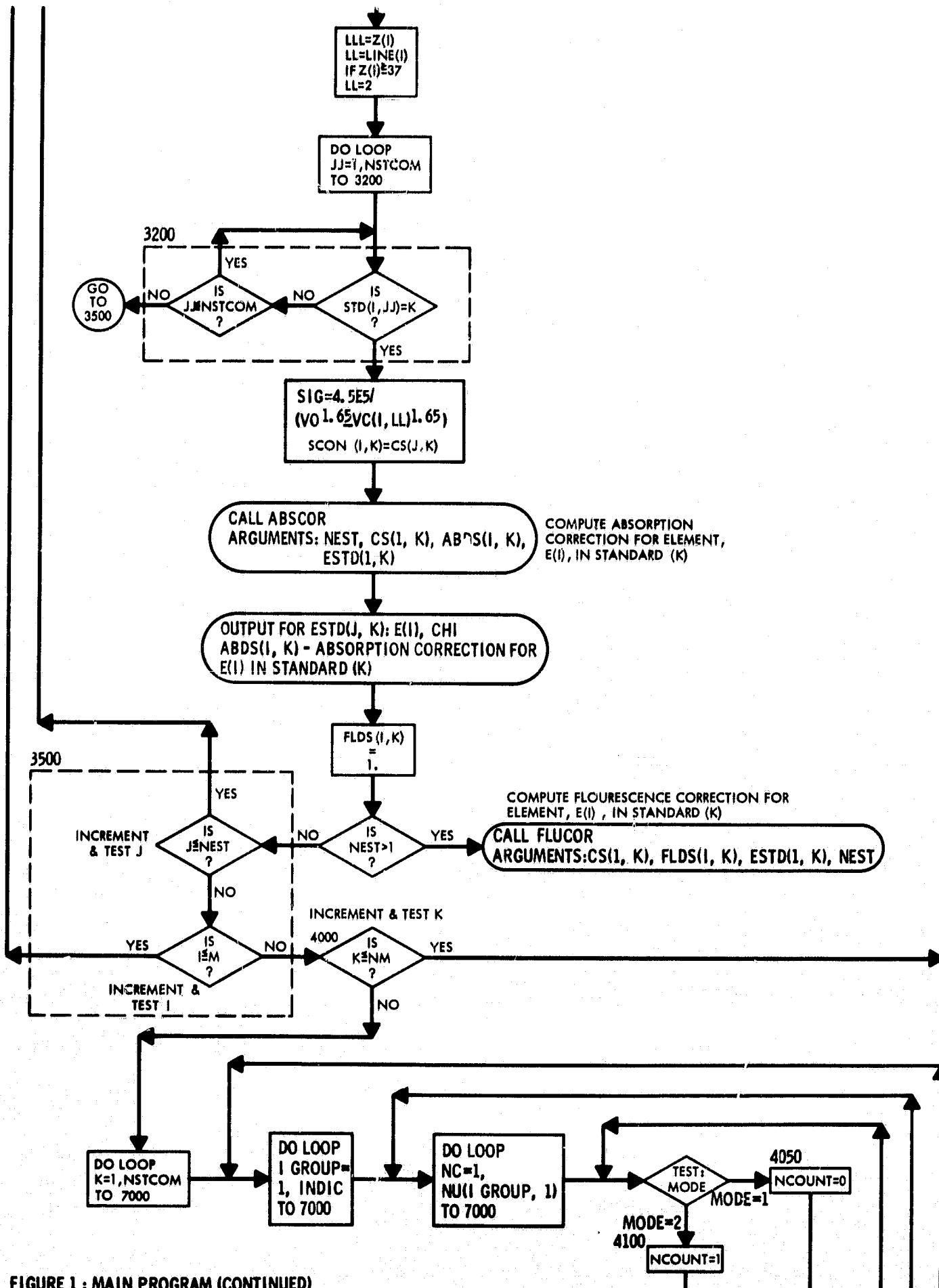


FIGURE 1 : MAIN PROGRAM (CONTINUED)

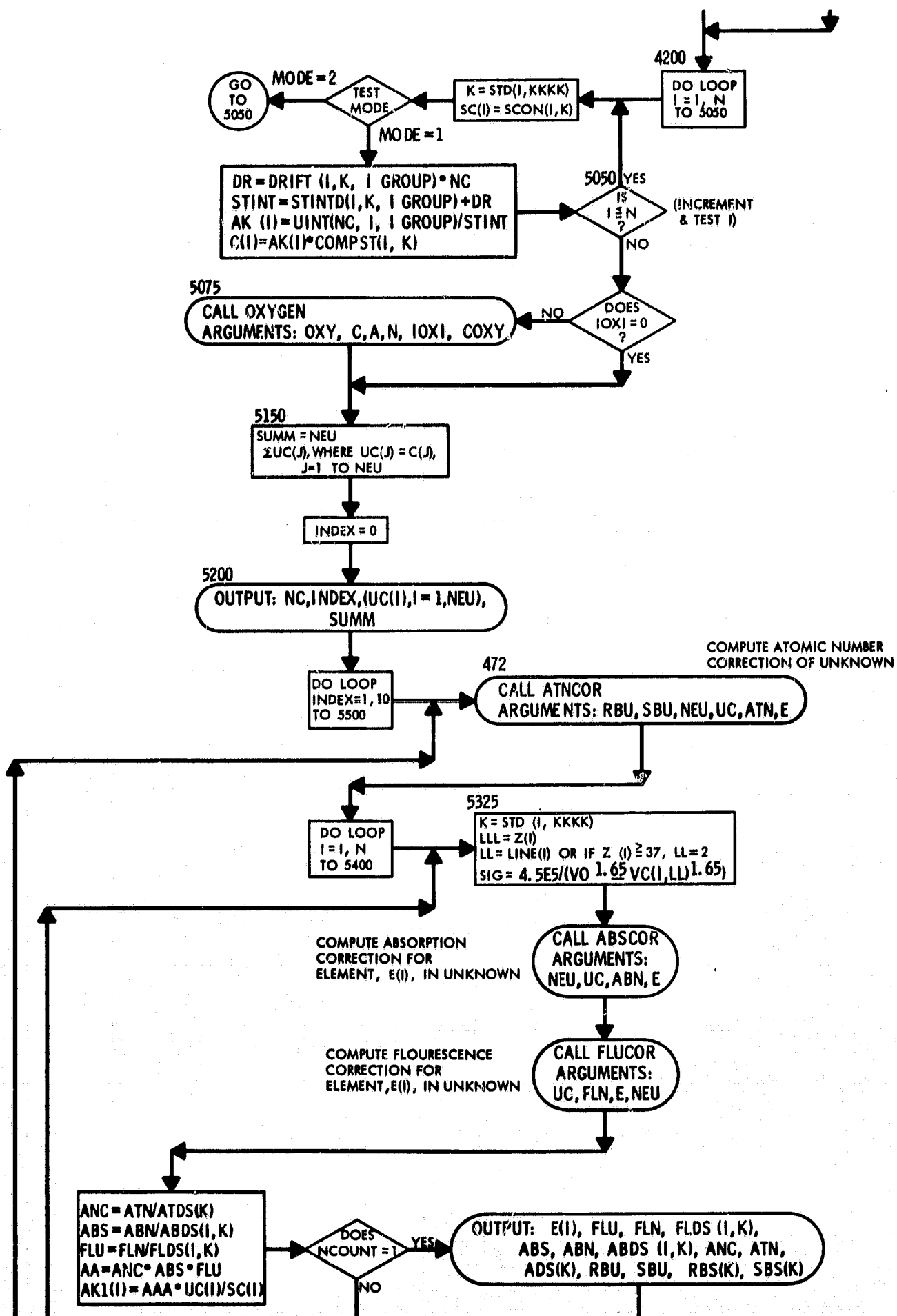


FIGURE 1 : MAIN PROGRAM (CONTINUED)



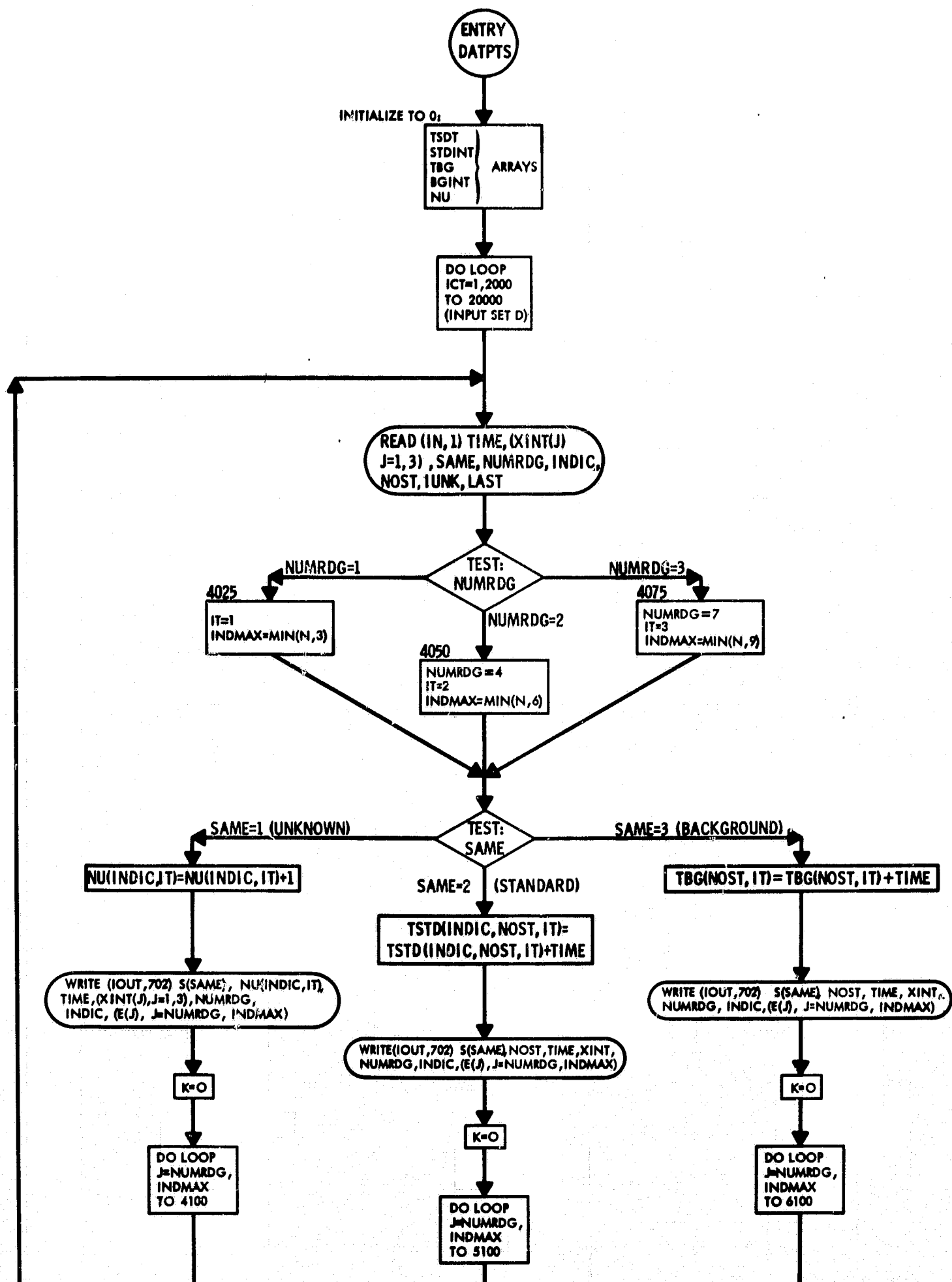


FIGURE 2: SUBROUTINE DATPTS

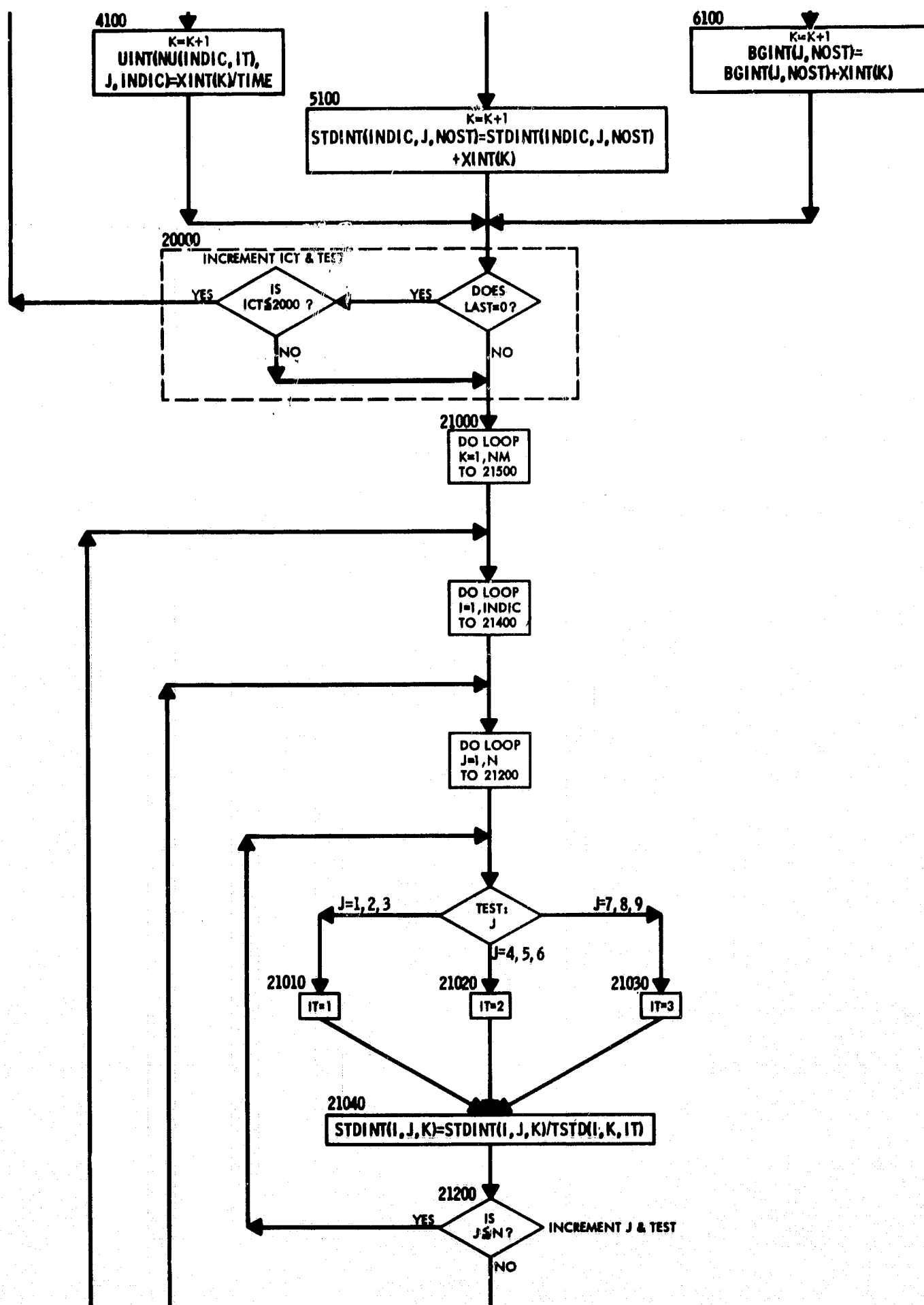
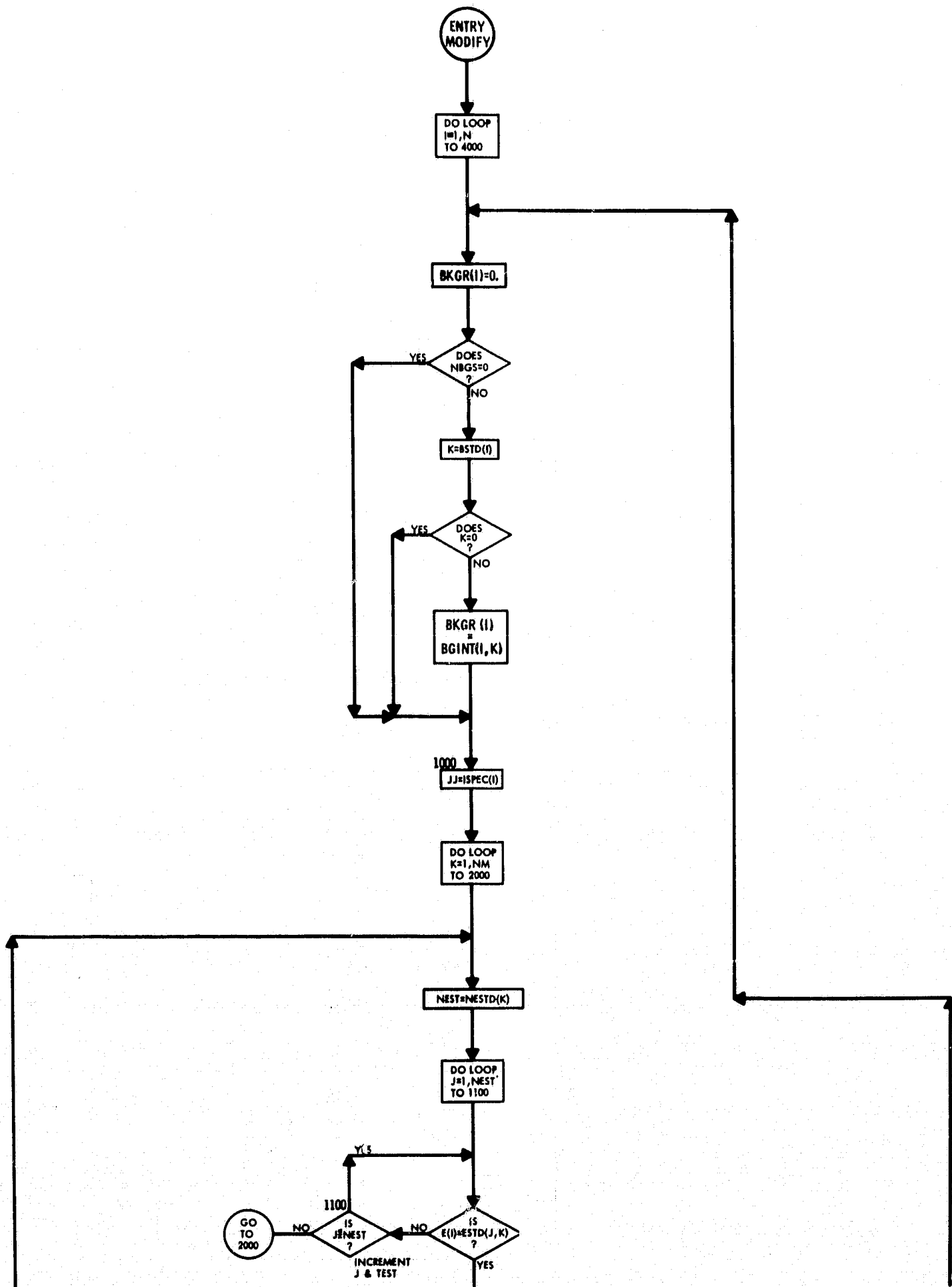


FIGURE 2 : SUBROUTINE DATPTS (CONTINUED)





FIGURE 1: SUBROUTINE MODIFY



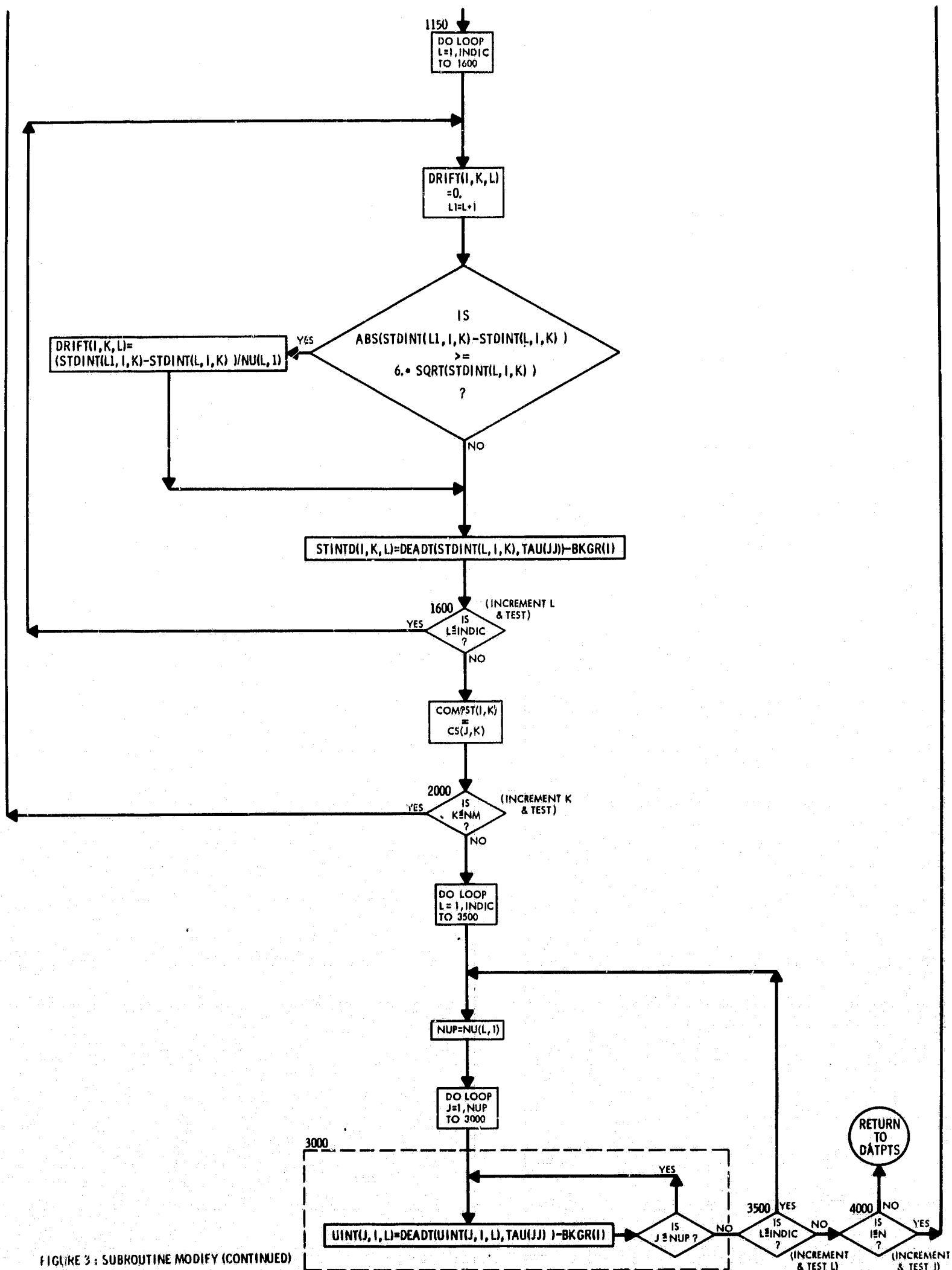


FIGURE 3 : SUBROUTINE MODIFY (CONTINUED)

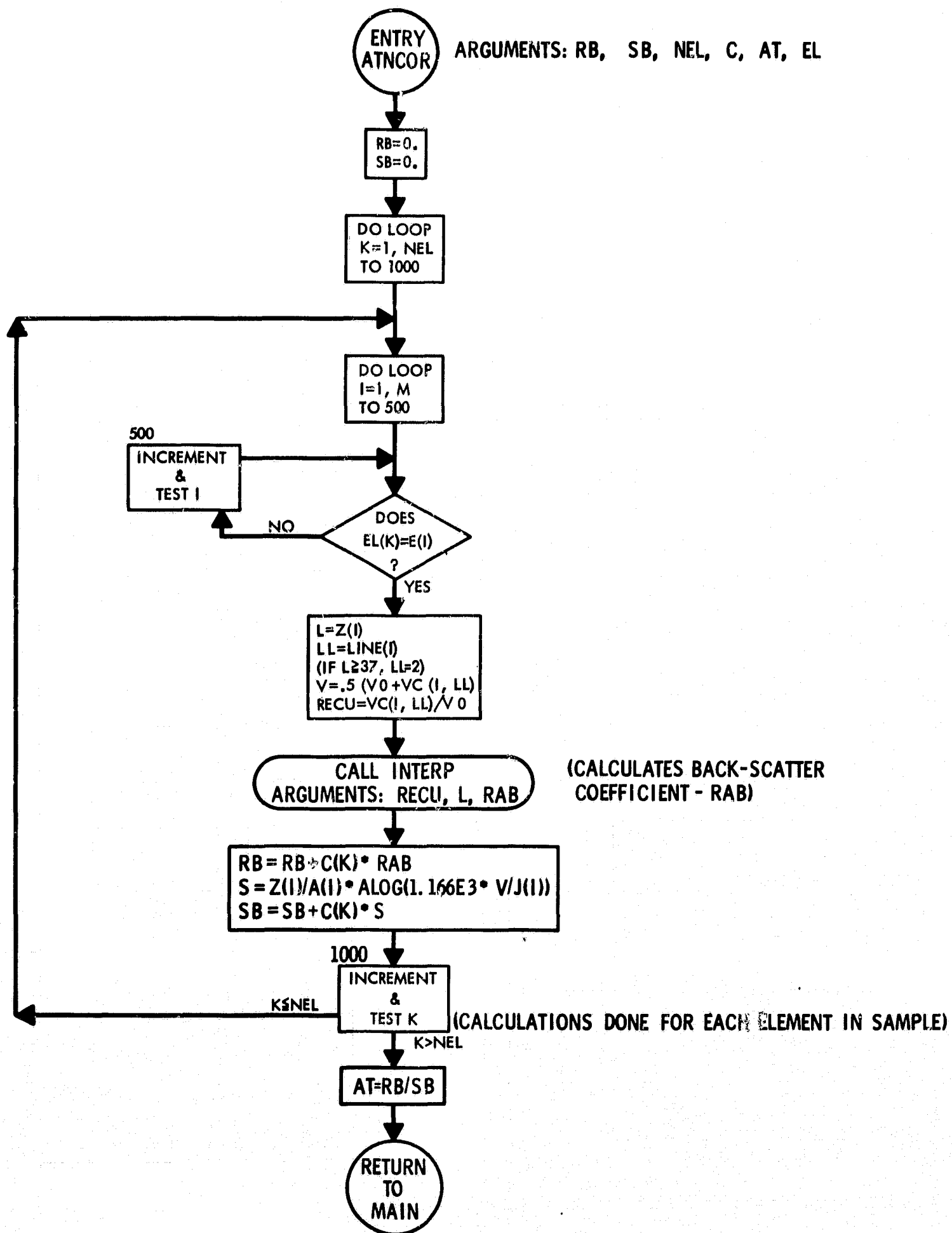


FIGURE 4: SUBROUTINE ATNCOR

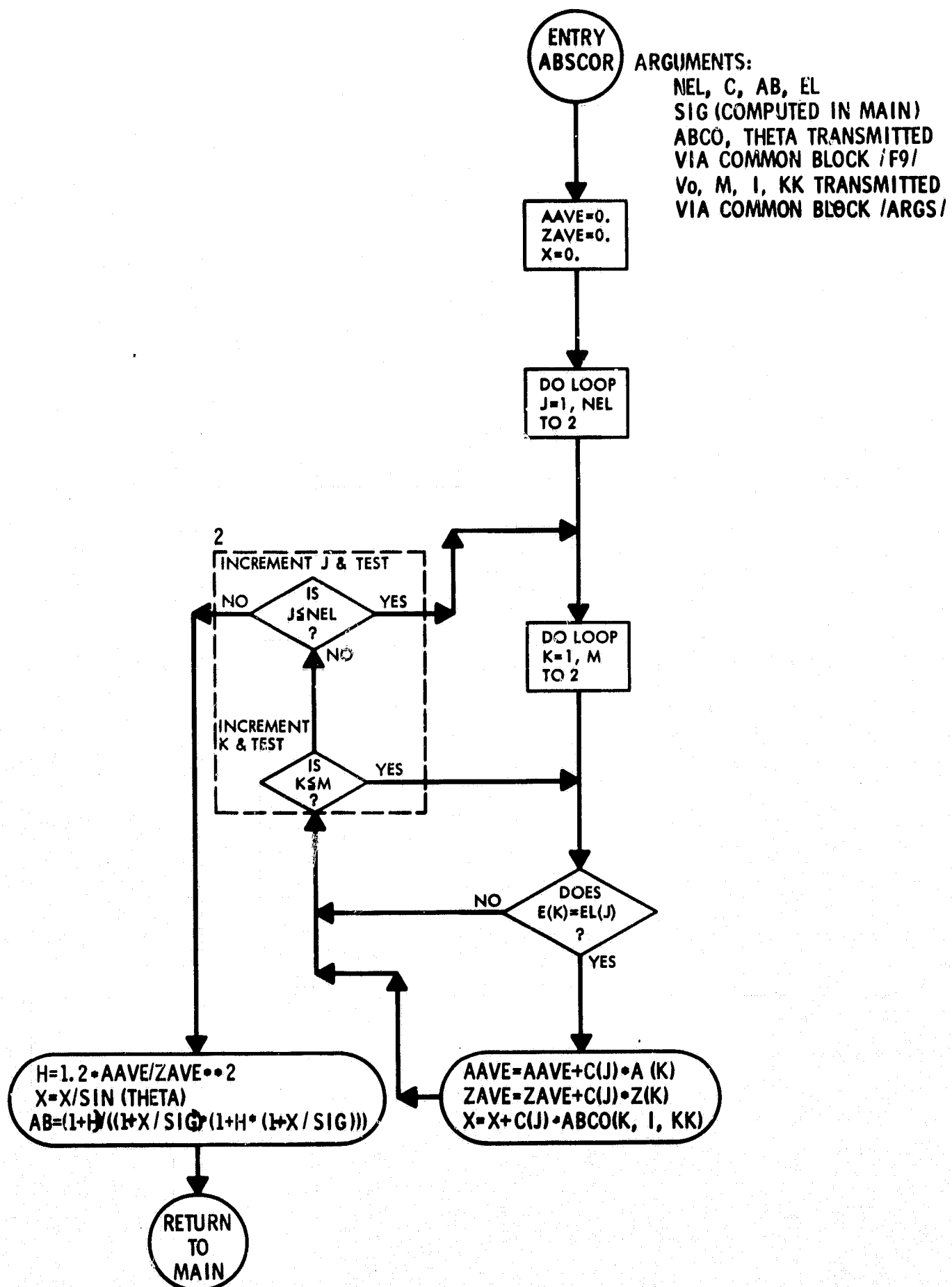


FIGURE 5: SUBROUTINE ABSCOR

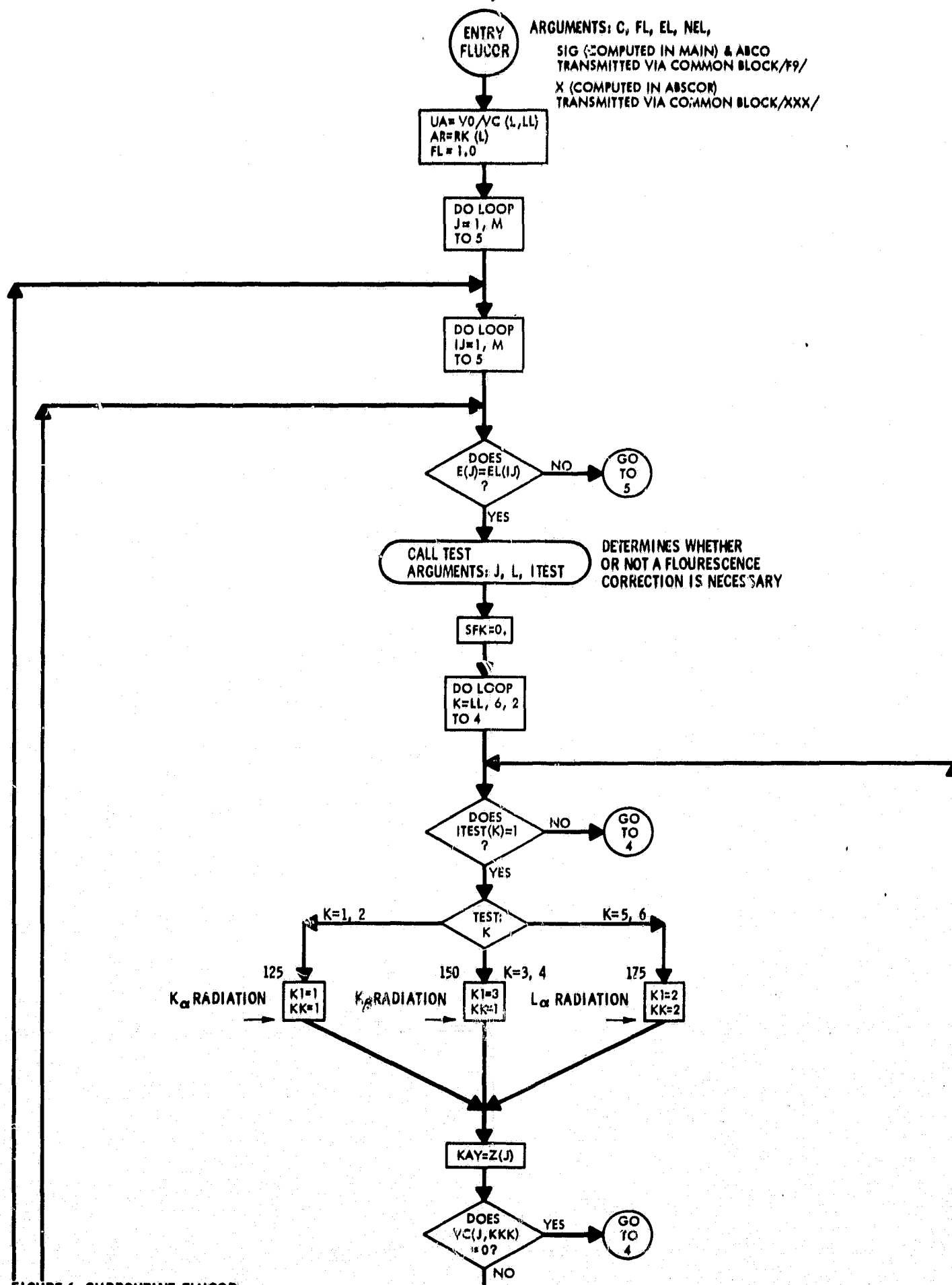


FIGURE 6: SUBROUTINE FLUCOR

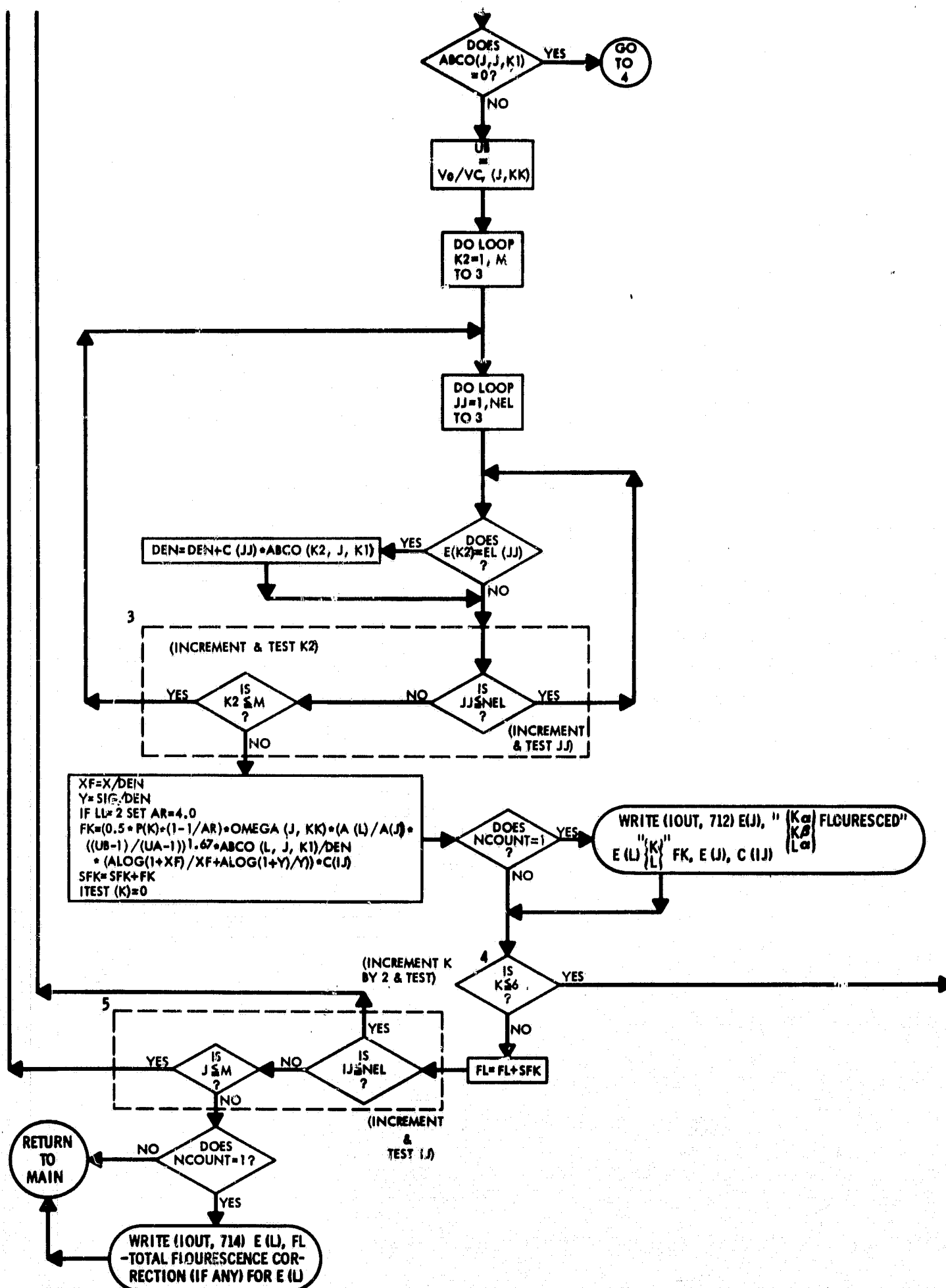
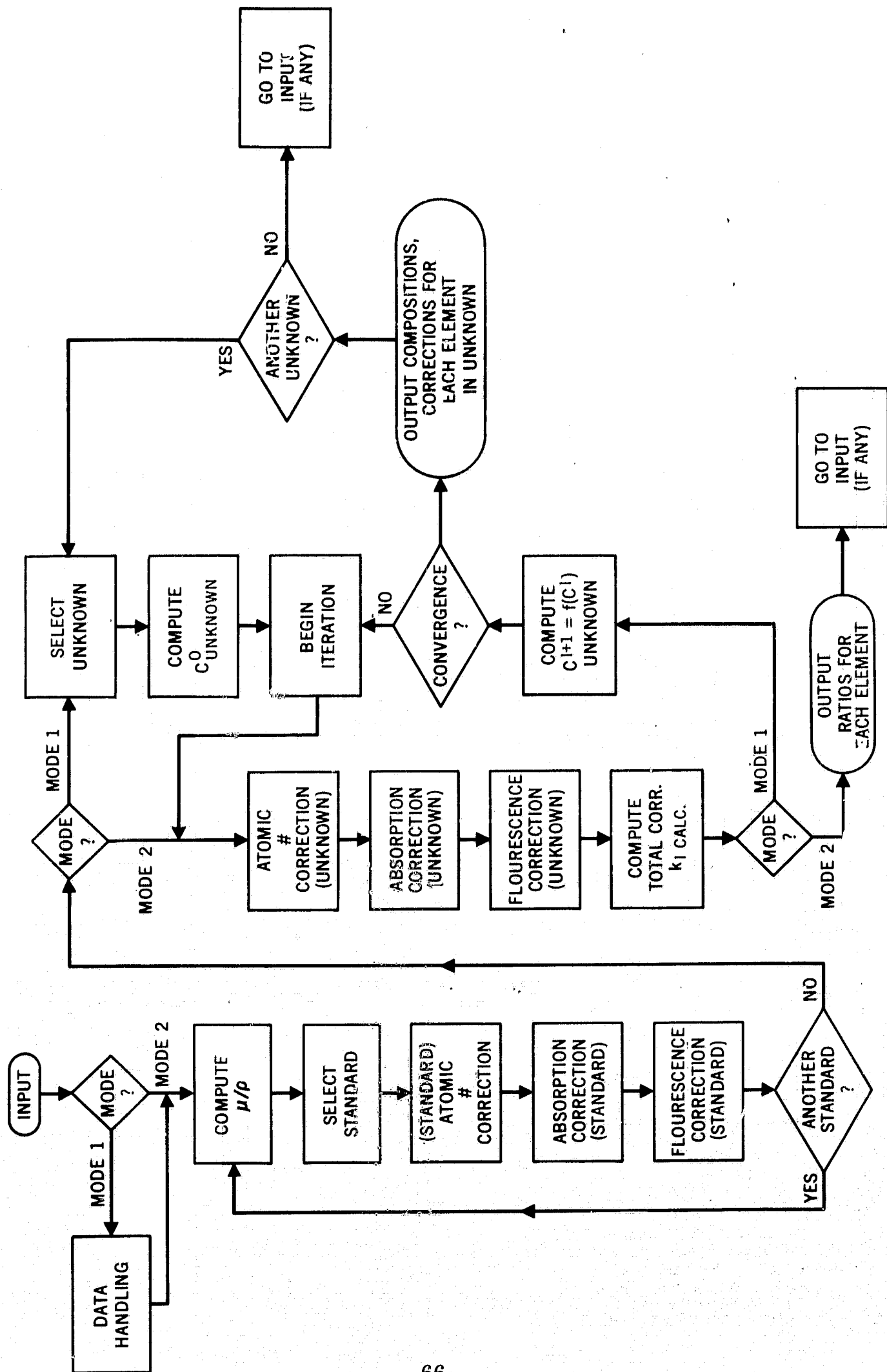


FIGURE 6 : SUBROUTINE FLUCOR (CONTINUED)

# MICROPROBE CORRECTION PROGRAM





```

&DATA THETA=52.5,TAU=.15E-5,.30E-5,.15E-5,STD=3*1,2,3,4,5,6.82*0,
N=8,NFH=10,M=10,NM=6,NBGS=3,MODE=1,NPROR=2,NSTCOM=1,LJNF=10*1K,
JSPEC=1,2,3,1,2,3,1,2,0,BSTD=3*0,1,2*0,3,2.0,VO=20.0,
&END

```

Figure 8--Input Data -- Sample Problem 1 -- DATA Files

```

ENDATA THETA=52.5,TAU=.15E-5,.30E-5,.15E-5,STD=3*1,2,3,4,5,6,82*0,
N=8,NFU=10,M=10,NM=6,NRGS=3,MODE=1,NPROB=2,NSTCONM=1,LINE=10*'K',
ISPFC=1,2,3,1,2,3,1,2,3,1,2,3,2,0,3,2,0,V0=20.0,
&END
'CA'20. 040.1 04.040 00.349 0239 09.92 1.0000 0.180 03.350 03.090 36.330 048.400
2.74 03.070 005.10 2.73
25.5
'MG'12. 024.2 01.300 00.045 0123 12.35 1.0000 0.030 09.889 09.558 50.000 011.750
2.79 09.512 000.89 2.73
250.7
'SI'14. 028.1 01.840 00.100 0154 11.64 0.5000 0.056 07.126 06.768 50.000 018.500
2.77 06.738 001.54 2.73
123.0
'AL'13. 027.0 01.560 00.075 0142 11.99 0.6667 0.042 08.327 07.981 50.000 014.870
2.78 07.951 001.18 2.73
170.4
'ND'11. 023.0 01.070 00.000 0126 12.74 2.0000 0.020 11.910 11.617 50.000 009.050
2.79 11.480 0.62 2.73
.01625
'MN'25. 054.6 06.540 00.000 0316 08.83 1.0000 0.290 02.103 01.910 19.500 086.700
2.72 01.896 010.45 2.73
8.85 5.9 10.1
.01375
'CR'24. 052.0 05.990 00.598 0301 09.03 1.0000 0.268 02.291 02.085 21.640 078.000
2.73 02.070 009.18 2.73
7.80 17.9 4.8 20.7
.03125
'FF'26. 055.9 07.110 00.707 0332 08.64 1.0000 0.320 01.937 01.757 17.600 095.800
2.72 01.743 011.75 2.73 17.2 10.0 17.2 7.0 17.525
.03000
'TI'22. 047.9 04.970 00.450 0270 09.45 1.0000 0.224 02.748 02.514 27.420 062.100
2.73 02.497 007.00 2.73
27.3
.03250
'N' 8. 016.0 00.532 00.000 0127 14.08 1.0000 0.004 23.700 00.000 50.000 003.800
2.82 23.320
.02750

```

Figure 9-Input Data -- Sample Problem 1 -- DATA and Element File



```

DATA THETA=52.5,TAI=1.5E-5,.30E-5,.15E-5,STD=3*1,2,3,4,5,6,82*0,
N=8,NFI=10,M=10,NM=6,NRGS=3,MIND=1,NPROR=2,NSTCOM=1,IJNF=10*1K1,
ISPEC=1,2,3,1,2,3,1,2,0,8STD=3*0,1,2*0,3,2,0,VN=20,0,
&END
'CA'20. 040.1 04.040 00.349 0239 04.92 1.0000 0.180 03.359 03.090 36.330 048.400
2.74 03.070 005.10 2.73 35.5
'MG'12. 024.2 01.300 00.040 0123 12.35 1.0000 0.030 09.889 09.558 50.000 011.750
2.79 09.512 000.89 2.73 250.7
'SI'14. 028.1 01.840 00.100 0154 11.64 0.5000 0.056 07.126 06.768 50.000 018.500
2.77 06.738 001.54 2.73 123.0
'AL'13. 027.0 01.560 00.075 0142 11.99 0.6667 0.042 08.337 07.981 50.000 014.870
2.78 07.951 001.18 2.73 170.4
'NA'11. 023.0 01.070 00.000 0126 12.74 2.0000 0.020 11.910 11.617 50.000 009.050
2.79 11.480 0.62 2.73 01625
'MN'25. 054.9 06.540 00.000 0316 08.83 1.0000 0.290 02.103 01.910 19.500 086.700
2.72 01.896 010.45 2.73 8.85 5.9 19.1
'CR'24. 052.0 05.990 00.598 0301 09.03 1.0000 0.268 02.291 02.085 21.640 078.000
2.73 02.070 009.18 2.73 16.7 7.80 17.9 4.8 20.7
'FF'26. 055.9 07.110 00.707 0332 08.64 1.0000 0.320 01.937 01.757 17.600 095.800
2.72 01.743 011.75 2.73 17.2 10.0 17.2 7.0 17.525
'TI'22. 047.9 04.970 00.450 0270 09.45 1.0000 0.224 02.748 02.514 27.420 062.100
2.73 02.497 007.00 2.73 03250
'N' 8. 016.0 00.532 00.000 0127 14.08 1.0000 0.004 23.700 00.000 50.000 003.800
2.82 23.320

```

```

&END
RNDYD'S PROBLEM
0. 0. 0. 0. 0. 0. 0.
0.1 .0334
STANDARD #1.....PURE DIOPSIDE
4 CA MG SI N
18.51 11.23 25.93 44.33
STANDARD #2.....ENAL 5
4 MG AL SI N
23.02 2.65 26.57 47.76
STANDARD #3.....DJ 65 JD 25
6 CA MG SI NA AL N
12.03 7.30 26.59 3.98 4.67 45.16
STANDARD #4.....PYROPE GLASS
5 MN MG AL SI N
4.19 16.29 12.05 20.95 46.52
STANDARD #5.....CR IN MD 75 GLASS
5 CA MG CR SI N
13.62 14.20 1.31 25.93 44.94
STANDARD #6.....P-140 FOSTERITE
4 MG FE SI N
31.04 5.66 19.27 44.03
BACKGROUND FOR AL
BACKGROUND FOR FE
BACKGROUND FOR CR
000010 0010000 0010000 0010000 211100
000010 0006240 0011170 0009690 111000
000010 0010000 0010000 0010000 212100
000010 0010000 0000000 0000000 221200
000010 0000000 0010000 0000000 221300
000010 0000000 0000000 0010000 221400
000010 0004350 0002210 0000215 121000
000010 0010000 0000000 0000000 222200
000010 0000000 0010000 0000000 222300
000010 0000000 0000000 0010000 222400
000010 0000010 0000000 0000000 322100
000010 0010000 0000000 0000000 231500
000010 0000000 0010000 0000000 231600
000010 0004500 0004550 0000000 121000
000010 0010000 0000000 0000000 232500
000010 0000000 0010000 0000000 232600
000010 0000010 0000000 0000000 332300
000010 0000000 0000010 0000000 3322001

```

Figure 11—Input Data — Sample Problem 1 — Intensity Data File

PARAMETERS USED										
ELEMENT	CA	MS	SI	AL	NA	MN	CR	FE	TI	O
ATOM.NN	20.0000	12.0000	14.0000	13.0000	11.0000	25.0000	24.0000	26.0000	22.0000	8.0000
ATOM.WT	40.1000	24.3000	28.1000	27.0000	23.0000	54.9000	52.0000	55.9000	47.5000	16.0000
K ALPHA	3.3590	6.8890	7.1260	8.3370	11.9100	2.1030	2.2910	1.9370	2.7480	23.7000
K BETA	3.0490	9.5590	6.7680	7.9810	11.6170	1.9100	2.0850	1.7570	2.5140	0.0
K EDGE	36.0700	9.5120	6.7380	7.9510	11.4800	1.8960	2.0700	1.7430	2.4970	23.3200
L ALPHA	36.0700	50.0000	50.0000	50.0000	50.0000	19.5000	21.6400	17.6000	27.4200	50.0000
L3 EDGE	35.5000	250.7000	123.0000	170.4000	0.0	19.1000	20.7000	17.5250	27.3000	0.0
CK	48.4000	11.7500	18.5000	14.8700	9.0500	86.7000	79.0000	95.8000	62.1000	3.9000
NK	2.7400	2.7900	2.7700	2.7800	2.7500	2.7200	2.7300	2.7200	2.7300	2.8200
CKL	5.1000	0.8900	1.5400	1.1900	0.6200	10.4500	9.1800	11.7500	7.6000	0.0
NKL	2.7300	2.7300	2.7300	2.7300	2.7300	2.7300	2.7300	2.7300	2.7300	0.0
L1	0.0	0.0	0.0	0.0	0.0	0.0	16.7000	17.2000	0.0	0.0
CL1	0.0	0.0	0.0	0.0	0.0	8.8500	7.8000	10.0000	0.0	0.0
L2	0.0	0.0	0.0	0.0	0.0	0.0	17.9000	17.2000	0.0	0.0
CL2	0.0	0.0	0.0	0.0	0.0	5.9000	4.8000	7.0000	0.0	0.0
CLM	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
M1	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
CM1	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
M2	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
CM2	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
M3	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
CM3	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
M4	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
M5	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
CMN	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
N1	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
VC(7,K)	4.7400	1.3000	1.8400	1.5600	1.0700	6.5400	5.9900	7.1100	4.5700	0.5320
VC(7,L)	0.7400	0.4900	0.1000	0.0750	0.0	0.0	0.5980	0.7070	0.4500	0.0
J1(7)	239	133	154	142	125	316	301	332	270	127
RK(7)	9.9200	12.3500	11.6400	11.9900	12.7400	8.8300	9.0300	8.6400	9.4500	14.0800
PHY(Z)	1.0000	1.0000	0.5000	0.6667	2.0000	1.0000	1.0000	1.0000	1.0000	1.0000
OMEGA-K	0.1900	0.0200	0.0560	0.0420	0.0200	0.2900	0.2680	0.3200	0.2240	0.0040
OMEGA-L	0.0250	0.0150	0.0175	0.0162	0.0137	0.0312	0.0300	0.0325	0.0275	0.0100
LINE	K	K	K	K	K	K	K	K	K	K
SPEC.NO	1	2	3	1	2	3	1	2		
VO	20.0000									
THETA	52.5000									
TAU	0.15000E-05	0.30000E-05	0.15000E-05							
N	8									
NM	6									
NBGS	3									
MODE	1									
NPRNR	1									

Figure 12-Output Data - Sample Problem 1

MASS ABSORPTION COEFFICIENTS (EMITTER AT TOP) - KALPHA RADIATION											
CA	MG	SI	AL	NA	MN	CR	FE	TI	O		
139.4	2656.7	1086.0	1667.0	4413.7	371.0	469.2	296.2	772.2	2882.3		
345.3	463.6	2915.0	4361.7	770.2	93.5	118.7	74.3	197.2	5040.2		
530.6	802.2	327.9	503.4	1322.8	145.0	183.8	115.5	304.3	8721.3		
431.7	614.7	3493.2	385.7	1021.2	117.4	149.0	93.4	247.0	6682.6		
265.9	5409.1	2168.1	3359.4	536.6	72.0	91.4	57.2	151.9	3511.2		
285.5	5443.5	2225.3	3415.6	5043.9	79.5	100.5	63.5	165.1	59160.5		
250.8	4782.0	1054.8	3000.5	7944.7	69.9	88.2	474.2	145.0	0.0		
221.1	6120.7	2502.1	3840.5	10168.9	89.4	113.0	71.4	185.6	0.0		
101.3	3646.4	1490.6	2288.0	6058.1	472.5	597.0	377.5	110.6	39642.4		
115.8	2432.8	965.6	1503.2	4110.1	30.9	35.4	24.5	65.7	0.0		
MASS ABSORPTION COEFFICIENTS (EMITTER AT TOP) - LALPHA RADIATION											
CA	MG	SI	AL	NA	MN	CR	FE	TI	O		
92703.0	0.0	0.0	0.0	0.0	16957.6	22533.1	12817.9	43002.5	0.0		
26177.6	0.0	0.0	0.0	0.0	2059.7	3932.2	2236.9	7504.3	0.0		
27992.7	0.0	0.0	0.0	0.0	5120.5	6804.1	3870.5	12985.1	0.0		
21448.0	0.0	0.0	0.0	0.0	3923.5	5213.5	2965.7	9949.6	0.0		
11269.8	0.0	0.0	0.0	0.0	2061.5	2739.3	1558.3	5227.7	0.0		
*****	0.0	0.0	0.0	0.0	34746.4	46170.8	26264.2	88112.9	0.0		
0.0	0.0	0.0	0.0	0.0	15960.1	0.0	19603.9	0.0	0.0		
0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0		
*****	0.0	0.0	0.0	0.0	23275.1	30927.8	17593.2	59023.0	0.0		
0.0	0.0	0.0	0.0	0.0	16507.5	22141.5	12363.1	0.0	0.0		
MASS ABSORPTION COEFFICIENTS (EMITTER AT TOP) - KBETA RADIATION											
CA	MG	SI	AL	NA	MN	CR	FE	TI	O		
111.0	2420.9	943.5	1470.7	4123.6	285.0	362.4	226.7	605.1	0.0		
273.5	422.5	2477.5	3861.7	719.6	71.5	91.3	56.6	153.8	0.0		
421.1	731.9	284.0	446.8	1245.2	111.1	141.6	88.1	237.8	0.0		
342.3	560.1	3026.8	342.4	554.1	89.9	114.7	71.2	192.9	0.0		
219.7	4512.9	1877.7	2574.3	501.3	55.0	70.3	42.6	118.5	0.0		
227.4	4960.4	1933.2	3032.0	8449.3	61.1	77.7	401.6	129.5	0.0		
109.7	4357.6	1698.2	2663.5	7422.5	456.4	68.2	363.3	113.7	0.0		
255.6	5577.5	2172.7	3409.2	5500.4	68.7	87.3	54.7	145.6	0.0		
152.3	3322.2	1295.0	2031.0	5659.8	363.3	461.6	289.3	65.7	0.0		
0.0	2210.3	835.0	1329.2	3831.3	23.6	30.2	18.6	51.1	0.0		

Figure 12-Output Data - Sample Problem 1 (Continued)

ELEMENT CHART - WEIGHT PERCENT

DATA PT	NO ITER	CA	MG	SI	AL	NA	MM	CR	FE	TI	O	SUM
1	1	11.544	12.544	25.135	1.150	0.878	0.050	0.500	2.500	0.100	42.074	56.275
2	2	11.535	12.742	25.924	1.166	0.917	0.091	0.500	2.505	0.100	44.715	100.258
3	3	11.551	12.842	26.037	1.073	0.926	0.091	0.501	2.507	0.100	44.574	100.772
4	4	11.557	12.845	26.064	1.075	0.928	0.091	0.501	2.507	0.100	44.574	100.772

FINAL CORRECTIONS: CA & FE GIVEN AS EXAMPLES

CA	MM-K ALPHA	FLUORESCENCE	CA-K	KF	0.500000E-04	CC (MN) = 0.500000E-03
	MM-K BETA	FLUORESCENCE	CA-K	KF	0.475000E-04	CC (MN) = 0.475000E-03
	CP-K ALPHA	FLUORESCENCE	CA-K	KF	0.475000E-04	CC (CP) = 0.475000E-03
	CP-K BETA	FLUORESCENCE	CA-K	KF	0.421000E-04	CC (CP) = 0.421000E-03
	FF-K ALPHA	FLUORESCENCE	CA-K	KF	0.421000E-04	CC (FF) = 0.421000E-03
	FF-K BETA	FLUORESCENCE	CA-K	KF	0.410000E-04	CC (FF) = 0.410000E-03
	TI-K ALPHA	FLUORESCENCE	CA-K	KF	0.410000E-04	CC (TI) = 0.410000E-03
	TI-K BETA	FLUORESCENCE	CA-K	KF	0.410000E-04	CC (TI) = 0.410000E-03
	TOTAL FLUORESCENCE CORRECTION FOR CA				0.100000E-01	

FLUORESCENCE CORRECTION (FLU)

FLU = 1.07205  
FLN = 1.07205  
FLD = 1.07205

ABSORPTION CORRECTION (ABS)

APS = 0.95000  
APN = 0.95000  
APD = 0.95000

ATOMIC NUMBER CORRECTION (ANC)

ANCE = 1.00176  
ANCE = 0.95000  
ATNE = 0.95000  
ATNE = 0.95000  
BRIE = 0.95000  
BRIE = 0.95000  
BRIE = 0.95000  
BRIE = 0.95000  
BRIE = 0.95000  
BRIE = 0.95000

FE

NO ELEMENT FLUORESCED FE... KF (TOTAL) = 1.0

FLUORESCENCE CORRECTION (FLU)

FLU = 1.00000  
FLN = 1.00000  
FLD = 1.00000

ABSORPTION CORRECTION (ABS)

APS = 0.95000  
APN = 0.95000  
APD = 0.95000

ATOMIC NUMBER CORRECTION (ANC)

ANCE = 1.00176  
ANCE = 0.95000  
ATNE = 0.95000  
ATNE = 0.95000  
BRIE = 0.95000  
BRIE = 0.95000  
BRIE = 0.95000  
BRIE = 0.95000  
BRIE = 0.95000  
BRIE = 0.95000

Figure 12-Output Data - Sample Problem 1 (Continued)

STANDARD #1  
THERE ARE 4 ELEMENTS IN THIS SAMPLE: CA MG SI O  
WEIGHT PERCENTAGES 11% FOR THE ELEMENTS: 0.1081 0.1057 0.4433

ATOMIC NUMBER CORRECTIONS: 0.41888 0.2

ABSORPTION CORRECTION FOR CA.....CHIK 0.11888 0.2 FICH1) 0.90108E 00  
NO ELEMENT FLUORESCED (A.....K(TOTAL)) 0.2

ABSORPTION CORRECTION FOR MG.....CHIK 0.03777 0.4 FICH1) 0.58333E 00  
CA-K ALPHA FLUORESCED MG-K : KFR 0.07178E-03 C( CA ) 0.16710E 00  
CA-K BETA FLUORESCED MG-K : KFR 0.07746E-02 C( CA ) 0.16710E 00  
SI-K ALPHA FLUORESCED MG-K : KFR 0.07178E-03 C( SI ) 0.28030E 00  
SI-K BETA FLUORESCED MG-K : KFR 0.07746E-02 C( SI ) 0.28030E 00  
TOTAL FLUORESCENCE CORRECTION FOR MG 0.10097E 01

ABSORPTION CORRECTION FOR SI.....CHIK 0.05244E 0.4 FICH1) 0.67214E 00  
CA-K ALPHA FLUORESCED SI-K : KFR 0.02597E-02 C( CA ) 0.16510E 00  
CA-K BETA FLUORESCED SI-K : KFR 0.02746E-02 C( CA ) 0.16510E 00  
TOTAL FLUORESCENCE CORRECTION FOR SI 0.10097E 01

STANDARD #2  
THERE ARE 4 ELEMENTS IN THIS SAMPLE: MG AL SI O  
WEIGHT PERCENTAGES 11% FOR THE ELEMENTS: 0.0303 0.0245 0.2057 0.4776

ATOMIC NUMBER CORRECTIONS: 0.01417 0.2

ABSORPTION CORRECTION FOR AL.....CHIK 0.11888 0.2 FICH1) 0.51243E 00  
SI-K ALPHA FLUORESCED AL-K : KFR 0.01417E-02 C( SI ) 0.26570E 00  
SI-K BETA FLUORESCED AL-K : KFR 0.01398E-02 C( SI ) 0.26570E 00  
TOTAL FLUORESCENCE CORRECTION FOR AL 0.10097E 01

STANDARD #3  
THERE ARE 5 ELEMENTS IN THIS SAMPLE: CA MG SI NA AL C  
WEIGHT PERCENTAGES 11% FOR THE ELEMENTS: 0.1081 0.1057 0.4433 0.4516

ATOMIC NUMBER CORRECTIONS: 0.41888 0.2

ABSORPTION CORRECTION FOR NA.....CHIK 0.03777 0.4 FICH1) 0.39648E 00  
CA-K ALPHA FLUORESCED NA-K : KFR 0.01417E-02 C( CA ) 0.16710E 00  
CA-K BETA FLUORESCED NA-K : KFR 0.01417E-02 C( CA ) 0.16710E 00  
MG-K ALPHA FLUORESCED NA-K : KFR 0.01417E-02 C( MG ) 0.72000E-01  
MG-K BETA FLUORESCED NA-K : KFR 0.01417E-02 C( MG ) 0.72000E-01  
SI-K ALPHA FLUORESCED NA-K : KFR 0.01417E-02 C( SI ) 0.28030E 00  
SI-K BETA FLUORESCED NA-K : KFR 0.01417E-02 C( SI ) 0.28030E 00  
AL-K ALPHA FLUORESCED NA-K : KFR 0.01417E-02 C( AL ) 0.46700E-01  
AL-K BETA FLUORESCED NA-K : KFR 0.01417E-02 C( AL ) 0.46700E-01  
TOTAL FLUORESCENCE CORRECTION FOR NA 0.10097E 01

STANDARD #4  
THERE ARE 4 ELEMENTS IN THIS SAMPLE: NA MG AL SI O  
WEIGHT PERCENTAGES 11% FOR THE ELEMENTS: 0.0303 0.0245 0.2057 0.4776

ATOMIC NUMBER CORRECTIONS: 0.41888 0.2

ABSORPTION CORRECTION FOR MN.....CHIK 0.07657E 0.2 FICH1) 0.87050E 00  
NO ELEMENT FLUORESCED MN.....K(TOTAL) 0.2

STANDARD #5  
THERE ARE 5 ELEMENTS IN THIS SAMPLE: CA MG CR SI O  
WEIGHT PERCENTAGES 11% FOR THE ELEMENTS: 0.1081 0.1057 0.4433 0.4494

ATOMIC NUMBER CORRECTIONS: 0.41888 0.2

ABSORPTION CORRECTION FOR CR.....CHIK 0.10543E 0.3 FICH1) 0.94454E 00  
NO ELEMENT FLUORESCED CR.....K(TOTAL) 0.2

STANDARD #6  
THERE ARE 4 ELEMENTS IN THIS SAMPLE: MG FE SI O  
WEIGHT PERCENTAGES 11% FOR THE ELEMENTS: 0.0303 0.0245 0.2057 0.4776

ATOMIC NUMBER CORRECTIONS: 0.41888 0.2

ABSORPTION CORRECTION FOR FE.....CHIK 0.07657E 0.2 FICH1) 0.87050E 00  
NO ELEMENT FLUORESCED FE.....K(TOTAL) 0.2

Figure 12-Output Data - Sample Problem 1 (Continued)



DATA FILE	{ <pre> DATA THETA=52.5, TAU=3.0, V0=20.0, M=2, NRCS=2, N=2, NFII=2, NM=2, RSTD(1)=1, RSTD(2)=2, STD(1,1)=1, STD(2,1)=2, NSTCMM=1, MODE=1, NPROR=1, LIME(1)=K, LIME(2)=L, ISPEC(1)=1, ISPEC(2)=2, \$END </pre>	
	{ <pre> TI 22. 047.9 04.970 00.450 0270 09.45 1.0000 0.224 02.743 02.514 27.120 062.100 2.73 02.497 007.00 2.73 \$END </pre>	
ELEMENT FILE	{ <pre> NR 41. 002.5 00.000 02.370 0543 00.00 1.0000 0.706 00.745 045.50 2.72 04.581 028.90 05.026 027.00 05.222 008.20 \$END </pre>	
	{ <pre> \$END </pre>	
COMPLETION FILE	{ <pre> 0. STANDARD-PURE TI 1 TI 100. STANDARD-PURE NR 1 NR 100. BACKGROUND FOR TI BACKGROUND FOR NR </pre>	
	{ <pre> 000001 0025100 0000000 0000000 2111000 000001 0000000 0010100 0000000 2112000 000001 0000160 0000000 0000000 2110000 000001 0000140 0000000 0000000 2111000 000001 0000000 0000135 0000000 2112000 000001 0000000 0000125 0000000 2112000 000001 0000300 0000230 0000000 1111000 000001 0025100 0000000 0000000 2121000 000001 0000000 0010100 0000000 2122000 </pre>	
INTENSITY DATA FILE	{ <pre> 000001 0025100 0000000 0000000 2111000 000001 0000000 0010100 0000000 2112000 000001 0000160 0000000 0000000 2110000 000001 0000140 0000000 0000000 2111000 000001 0000000 0000135 0000000 2112000 000001 0000000 0000125 0000000 2112000 000001 0000300 0000230 0000000 1111000 000001 0025100 0000000 0000000 2121000 000001 0000000 0010100 0000000 2122000 </pre>	

PARAMETERS USED

ELEMENT	TI	NR
ATOM.NO	22.0000	41.0000
ATOM.WT	47.9200	92.9000
K ALPHA	2.7490	0.7450
K BETA	2.5140	0.0
K EDGE	2.4970	0.0
L ALPHA	27.4200	5.7240
L3 EDGE	27.3000	5.2230
CK	62.1000	0.0
NK	2.7300	0.0
CKL	7.0000	45.5000
NKL	2.7300	2.7300
L1	0.0	4.5810
CL1	0.0	38.9000
L2	0.0	5.0260
CL2	0.0	27.2000
CLM	0.0	8.7500
M1	0.0	0.0
CM1	0.0	0.0
M2	0.0	0.0
CM2	0.0	0.0
M3	0.0	0.0
CM3	0.0	0.0
M4	0.0	0.0
M5	0.0	0.0
CMN	0.0	0.0
N1	0.0	0.0
VC(7,K)	4.6700	0.0
VC(7,L)	0.4500	2.3700
J1(7)	270	550
RK(7)	9.4500	0.0
ORV(7)	1.0000	0.0000
OMEGA-K	0.2240	0.7460
OMEGA-L	0.0275	0.0560
LINE	K	L
SPEC.NO	1	2
VO	20.0000	
THETA	52.5000	
TAU	0.0	0.0
N	2	
NM	2	
NRGS	2	
MODE	1	
NPROB	2	

#### MASS ABSORPTION COEFFICIENTS

MASS ABSORPTION COEFFICIENTS (EMITTER AT TOP) -- KALPHA RADIATION		
TI	NR	
TI	110.6	27.9
NR	718.7	20.4
MASS ABSORPTION COEFFICIENTS (EMITTER AT TOP) -- LALPHA RADIATION		
TI	NR	
TI	59027.0	819.6
NR	45907.7	783.0
MASS ABSORPTION COEFFICIENTS (EMITTER AT TOP) -- KBETA RADIATION		
TI	NR	
TI	86.7	0.0
NR	567.6	0.0

Figure 14—Output Data — Sample Problem 2

IN STANDARD COMBINATION # 1 TI USES STANDARD # 1  
 IN STANDARD COMBINATION # 1 NB USES STANDARD # 2

#### ELEMENT CHART - WEIGHT PERCENT

DATA PT	NO ITER	TI	NB	SUM
1	0	32.665	62.187	94.852
2	1	34.491	60.616	95.107
3	2	34.712	65.332	100.044
4	3	34.762	65.462	100.224

#### FINAL CORRECTIONS

NO ELEMENT FLOURESCED TI...KF(TOTAL)= 1.0

##### FLUORESCENCE CORRECTION (FLU)

FLU = 1.00000  
 FLN = 1.00000  
 FLD = 1.00000

##### ABSORPTION CORRECTION (ABS)

ABS = 0.87307  
 ARN = 0.87675  
 ABD = 0.95845

##### ATOMIC NUMBER CORRECTION (ANC)

ANC= 1.07604  
 ATN= 0.51920  
 ATD= 0.48251  
 PRIJ= 0.80462E 00  
 SRIJ= 0.15497E 01  
 PRS= 0.98366E 00  
 SRS= 0.18714E 01

TI-K ALPHA FLOURESCED NB-L . KF= 0.36801E-02  
 TI-K BETA FLOURESCED NB-L . KF= 0.32807E-03  
 TOTAL FLUORESCENCE CORRECTION FOR NB= 0.10040E 01

C( TI )= 0.34762E 00  
 C( TI )= 0.34762E 00

##### FLUORESCENCE CORRECTION (FLU)

FLU = 1.00401  
 FLN = 1.00401  
 FLD = 1.00300

##### ABSORPTION CORRECTION (ABS)

ABS = 0.99238  
 ARN = 0.75041  
 ABD = 0.75617

##### ATOMIC NUMBER CORRECTION (ANC)

ANC= 0.95305  
 ATN= 0.51920  
 ATD= 0.54476  
 PRIJ= 0.80462E 00  
 SRIJ= 0.15497E 01  
 PRS= 0.75989E 00  
 SRS= 0.13949E 01

Figure 14-Output Data - Sample Problem 2 (continued)

<b>DATA FILE</b>	<pre> &amp;DATA THETA=52.5,TAU=.15E-5,.30E-5,.15E-5,MODE=2,V0=20.0, STD(1,1)=3,STD(2,1)=4,STD(3,1)=2,STD(4,1)=1, N=4,NEU=4,M=4, LINE(1)='K',LINE(2)='K',LINE(3)='L',LINE(4)='K', ISPEC=1,2,3,1,5#0,NPROR=3,NSTCOM=1,NM=4,NRGS=0, &amp;END </pre>
<b>ELEMENT FILE</b>	<pre> AL 13. 027.0 01.560 00.075 0142 11.99 0.6667 0.042 08.337 07.981 50.000 014.870 2.78 07.951 001.18 2.73 MG 12. 024.2 01.300 00.040 0123 12.35 1.0000 0.030 09.880 09.558 50.000 011.750 2.79 09.512 000.89 2.73 AU 179. 197.2 00.000 11.900 1071 00.00 1.0000 1.784 272.00 2.59 00.863 232.00 00.903 167.00 01.040 067.64 03.620 058.3 03.939 48.3 04.522 41.8 05.415 05.629 22.50 CU 129. 063.5 08.980 00.923 0377 08.12 1.0000 0.410 01.542 01.302 12.400 126.800 2.71 01.380 016.45 2.73 11.270 014.06 13.01 10.1 13.20 </pre>
<b>COMPOSITION FILE</b>	<pre> MODE 2 PRORLEF 10. 50. 40. 0.1 STANDARD #1.....PURE COPPER 1 CU 100. STANDARD #2.....PURE GOLD 1 AU 100. STANDARD #3.....PURE ALUMINIUM 1 AL 100. STANDARD #4.....PURE MAGNESIUM 1 MG 100. </pre>

Figure 15-Input Data - Sample Problem 3

ELEMENT	AL	MG	AU	CU
ATOM.NO	13.0000	12.0000	79.0000	25.0000
ATOM.WT	27.0000	24.3000	197.2000	63.5000
K ALPHA	8.3370	9.8890	0.0	1.5420
K BETA	7.9810	9.5520	0.0	1.3920
K EDGE	7.9510	9.5120	0.0	1.3800
L ALPHA	50.0000	50.0000	1.2760	13.4000
L3 EDGE	170.4000	250.7000	1.0400	13.2900
CK	14.8700	11.7500	0.0	126.8000
NK	2.7800	2.7900	0.0	2.7100
CKL	1.1800	0.8900	272.0000	16.4500
NKL	2.7300	2.7300	2.5900	2.7300
L1	0.0	0.0	0.8630	11.2700
L2	0.0	0.0	232.0000	14.0600
L3	0.0	0.0	0.5030	13.0100
CL2	0.0	0.0	167.0000	10.1000
CLM	0.0	0.0	67.6400	0.0
M1	0.0	0.0	3.6200	0.0
M2	0.0	0.0	58.3000	0.0
M3	0.0	0.0	7.9390	0.0
M4	0.0	0.0	48.3000	0.0
M5	0.0	0.0	4.5220	0.0
CM1	0.0	0.0	41.8000	0.0
CM2	0.0	0.0	5.4150	0.0
CM3	0.0	0.0	5.6250	0.0
CM4	0.0	0.0	22.5000	0.0
CM5	0.0	0.0	0.0	0.0
VC(7,K)	1.5600	1.3000	0.0	8.9800
VC(7,L)	0.6750	0.0490	11.5000	0.9330
JK(7)	1.02	133	1071	377
PK(7)	11.9500	12.3500	0.0	8.1200
OKY(7)	0.6667	1.0000	1.0000	1.0000
OMEGA-K	0.0420	0.0300	1.7840	0.4100
OMEGA-L	0.0142	0.0150	0.3650	0.0362
LINE	K	K	L	K
SPEC.NO	1	2	3	1
V0	20.0000			
THETA	52.5000			
TAU	0.15000E-05	0.30000E-05	0.15000E-05	
N	4			
NM	4			
NBGS	0			
MODE	2			
NPRNR	3			

MASS ABSORPTION COEFFICIENTS

MASS ABSORPTION	COEFFICIENTS (EMITTER AT TCP)	MASS ABSORPTION	COEFFICIENTS (EMITTER AT TCP)
AL	AU	AL	AU
285.7	0.0	285.7	0.0
4251.7	0.0	4251.7	0.0
16781.4	0.0	16781.4	0.0
5375.8	0.0	5375.8	0.0

MASS ABSORPTION COEFFICIENTS (EMITTER AT TCP) -- KALPHA RADIATION

MASS ABSORPTION	COEFFICIENTS (EMITTER AT TCP)	MASS ABSORPTION	COEFFICIENTS (EMITTER AT TCP)
AL	AU	AL	AU
0.0	29.3	0.0	1408.9
0.0	23.2	0.0	1062.6
0.0	127.5	0.0	57633.6
0.0	245.5	0.0	0.0

MASS ABSORPTION COEFFICIENTS (EMITTER AT TCP) -- KALPHA RADIATION

MASS ABSORPTION	COEFFICIENTS (EMITTER AT TCP)	MASS ABSORPTION	COEFFICIENTS (EMITTER AT TCP)
AL	AU	AL	AU
342.4	0.0	342.4	0.0
3861.7	0.0	3861.7	0.0
14581.4	0.0	14581.4	0.0
4772.9	0.0	4772.9	0.0

MASS ABSORPTION COEFFICIENTS (EMITTER AT TCP) -- KBETA RADIATION

Figure 16-Output Data - Sample Problem 3

AL=L ALPHA FLUORESCENCE AL=K : KFA 0.18917E-04  
 CU=K ALPHA FLUORESCENCE AL=V : KFA 0.12737E-06  
 CU=V BETA FLUORESCENCE AL=K : KFA 0.18667E-07  
 TOTAL FLUORESCENCE CORRECTION FOR AL 0.18917E-04  
 CI AU 3# 0.40501E-00  
 CI CU 3# 0.10000E-02  
 CI CU 3# 0.10000E-02

FLUORESCENCE CORRECTION (FLU): AL  
 FLU = 1.00000  
 FLN = 0.00000  
 FLV = 0.00000

ABSORPTION CORRECTION (ABS): AL  
 ABS = 0.00000  
 APN = 0.00000  
 APV = 0.00000

ATOMIC NUMBER CORRECTION (ANC): AL  
 ANC = 1.00000  
 ATN = 0.00000  
 ATV = 0.00000  
 BRN = 0.00000  
 BRV = 0.00000  
 BRN = 0.00000  
 BRV = 0.00000

AL=V ALPHA FLUORESCENCE AL=V : KFA 0.18917E-04  
 AL=V BETA FLUORESCENCE AL=V : KFA 0.12737E-06  
 AL=L ALPHA FLUORESCENCE AL=V : KFA 0.18667E-07  
 CU=K ALPHA FLUORESCENCE AL=V : KFA 0.18917E-04  
 CU=V BETA FLUORESCENCE AL=V : KFA 0.12737E-06  
 TOTAL FLUORESCENCE CORRECTION FOR V 0.18917E-04  
 CI AL 3# 0.00000E-00  
 CI AL 3# 0.00000E-00  
 CI AU 3# 0.40501E-00  
 CI CU 3# 0.10000E-02  
 CI CU 3# 0.10000E-02

FLUORESCENCE CORRECTION (FLU): MG  
 FLU = 1.00000  
 FLN = 0.00000  
 FLV = 0.00000

ABSORPTION CORRECTION (ABS): MG  
 ABS = 0.00000  
 APN = 0.00000  
 APV = 0.00000

ATOMIC NUMBER CORRECTION (ANC): MG  
 ANC = 1.00000  
 ATN = 0.00000  
 ATV = 0.00000  
 BRN = 0.00000  
 BRV = 0.00000  
 BRN = 0.00000  
 BRV = 0.00000

NO ELEMENT FLUORESCED AU...KFA(TOTAL)= 1.0

FLUORESCENCE CORRECTION (FLU): AU  
 FLU = 1.00000  
 FLN = 0.00000  
 FLV = 0.00000

ABSORPTION CORRECTION (ABS): AU  
 ABS = 0.00000  
 APN = 0.00000  
 APV = 0.00000

ATOMIC NUMBER CORRECTION (ANC): AU  
 ANC = 0.70547  
 ATN = 0.00000  
 ATV = 0.00000  
 BRN = 0.00000  
 BRV = 0.00000  
 BRN = 0.00000  
 BRV = 0.00000

AU=L ALPHA FLUORESCENCE CU=K : KFA 0.72031E-01  
 TOTAL FLUORESCENCE CORRECTION FOR CU 0.72031E-01  
 CI AU 3# 0.40501E-00

FLUORESCENCE CORRECTION (FLU): CU  
 FLU = 1.00000  
 FLN = 0.00000  
 FLV = 0.00000

ABSORPTION CORRECTION (ABS): CU  
 ABS = 0.00000  
 APN = 0.00000  
 APV = 0.00000

ATOMIC NUMBER CORRECTION (ANC): CU  
 ANC = 0.00000  
 ATN = 0.00000  
 ATV = 0.00000  
 BRN = 0.00000  
 BRV = 0.00000  
 BRN = 0.00000  
 BRV = 0.00000

MODE 2 PROBLEM: INPUT-COMPOSITION RATIOS...OUTPUT-INTENSITY RATIOS  
 PROBLEM 3  
 INPUT...INTENSITY RATIO-SAMPLE TO STANDARD-ATKINS-FLUORESCENCE  
 AL 0.24227E-01  
 MG 0.10000E-00  
 AU 0.40501E-00  
 CU 0.10000E-02

Figure 16-Output Data – Sample Problem 3 (Continued)

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